# Measurement of genuine three-particle Bose-Einstein correlations in hadronic Z decay 

L3 Collaboration

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#### Abstract

We measure three-particle Bose-Einstein correlations in hadronic $Z$ decay with the L3 detector at LEP. Genuine three-particle Bose-Einstein correlations are observed. By comparing two- and three-particle correlations we find that the data are consistent with fully incoherent pion production. © 2002 Elsevier Science B.V. All rights reserved.


## 1. Introduction

So far, no theory exists which can describe the non-perturbative process of hadron production in general and Bose-Einstein (BE) effects in particular. The latter are expected from general spin statistics considerations. To help understand these phenomena, studies of identical-boson correlations in $\mathrm{e}^{+} \mathrm{e}^{-}$collisions at LEP have been performed in terms of the absolute four-momentum difference $Q$ [1], as well as in two- and three-dimensional distributions in components of $Q[2,3]$.

It has long been realized that the shape and size in spacetime of a source of pions can be determined, as a consequence of the interference of identical bosons, from the shape and size of the correlation function of two identical pions in energy-momentum space [4]. Additional information can be derived from higher-order correlations. Furthermore, such correlations constitute an important theoretical issue for the understanding of BoseEinstein correlations (BEC) [5].

In this Letter three-particle correlations are analysed. These correlations are sensitive to asymmetries in the particle production mechanism [6,7] which cannot be studied by two-particle correlations. In addition, the combination of two- and three-particle correlation analyses gives access to the degree of coherence of pion production [8,9], which is very difficult to investigate from two-particle correlations alone due to the effect of long-lived resonances on the correlation function. The DELPHI [10] and OPAL [11] Collaborations have both studied three-particle correlations but did not investigate the degree of coherence.

## 2. The data and Monte Carlo

The data used in this analysis were collected by the L3 detector [12] in 1994 at a centre-of-mass energy of 91.2 GeV and correspond to a total integrated luminosity of $48.1 \mathrm{pb}^{-1}$. The Monte Carlo (MC) event generators JETSET [13] and HERWIG [14] are used to simulate the signal process. Within JETSET, BEC are simulated using the $\mathrm{BE}_{0}$ algorithm $[15,16] .{ }^{7}$ The generated events are passed through the L 3 detector simulation program, which is based on the GEANT [17] and GHEISHA [18] programs, reconstructed and subjected to the same selection criteria as the data.

The event selection is identical to that presented in Ref. [2], resulting in about one million hadronic Z decay events, with an average track multiplicity of about 12 . Two additional cuts are performed in order to reduce the dependence of the detector correction on the MC model used: tracks with measured momentum greater than 1 GeV are rejected, as are pairs of like-sign tracks with opening angle below $3^{\circ}$. This results in an average track multiplicity of about 7. For the computation of three-particle correlations, each possible triplet of like-sign tracks is used to compute the variable $Q_{3} \equiv \sqrt{Q_{12}^{2}+Q_{23}^{2}+Q_{31}^{2}}$, where $Q_{i j} \equiv \sqrt{-\left(p_{i}-p_{j}\right)^{2}}$ is the absolute four-momentum difference between particles $i$ and $j$. Since $Q_{i j}$, and thus $Q_{3}$, depends both on the energy of the particles and on the

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Fig. 1. Normalized distributions of the sum of the difference in (a) azimuthal and (b) polar angle of pairs of tracks in a triplet, $\sum \delta \phi$ and $\sum \delta \theta$, and of (c) $Q_{3}$. Data, JETSET, with and without BE effects, and HERWIG are displayed. The ratios between the data and MC distributions are shown in (d), (e) and (f).
angle between them, small $Q_{i j}$ can be due to small angles or low energies. In a MC generator with BE effects, the fraction of pairs at small $Q_{i j}$ with small angle is larger than in one without. Consequently, the estimated detection efficiency depends on the MC model used. The momentum and opening angle cuts reduce this model dependence. After selection, the average triplet multiplicity is about 6. In the region of interest, $Q_{3}<1 \mathrm{GeV}$, the loss of triplets by the momentum and opening angle cut is about $40 \%$.

The momentum cut improves the resolution of $Q_{3}$ by a factor three with respect to that for the full momentum spectrum. Using MC events, its average is estimated to be 26 MeV for triplets of tracks with $Q_{3}<0.8 \mathrm{GeV}$. We choose a bin size of 40 MeV , somewhat larger than this resolution.

In Fig. 1, the data are compared to JETSET (with and without BE effects) and HERWIG (not having a BE option) at the detector level, after performing all the cuts mentioned above, in the three-particle distributions $\sum \delta \phi$, $\sum \delta \theta$, and $Q_{3}$. The sums run over the three pairs of like-sign tracks in the triplet and $\delta \phi$ and $\delta \theta$ are the absolute differences in azimuthal and polar angle between two tracks, respectively. Within $10 \%$, the angular distributions of the MC models agree with those of the data. None of the models describes the $Q_{3}$ distribution: JETSET with BE
effects overestimates the data by approximately $20 \%$ at low $Q_{3}$, even though we found good agreement for $\delta \phi, \delta \theta$, and $Q$ [2]. JETSET without BE effects and HERWIG grossly underestimate the data at low $Q_{3}$. The statistics for $Q_{3}<160 \mathrm{MeV}$ are so poor, that this region is rejected from the analysis.

## 3. The analysis

The three-particle number density $\rho_{3}\left(p_{1}, p_{2}, p_{3}\right)$ of particles with four-momenta $p_{1}, p_{2}$ and $p_{3}$ can be described in terms of single-particle, two-particle and genuine three-particle densities as

$$
\begin{equation*}
\rho_{3}\left(p_{1}, p_{2}, p_{3}\right)=\rho_{1}\left(p_{1}\right) \rho_{1}\left(p_{2}\right) \rho_{1}\left(p_{3}\right)+\sum_{(3)}\left\{\rho_{1}\left(p_{1}\right)\left[\rho_{2}\left(p_{2}, p_{3}\right)-\rho_{1}\left(p_{2}\right) \rho_{1}\left(p_{3}\right)\right]\right\}+C_{3}\left(p_{1}, p_{2}, p_{3}\right), \tag{1}
\end{equation*}
$$

where the sum is over the three possible permutations and $C_{3}$ is the third-order cumulant, which measures the genuine three-particle correlations. The $\rho_{1} \rho_{2}$ terms contain all the two-particle correlations. In order to focus on the correlation due to BE interference, we replace products of single-particle densities by the corresponding twoor three-particle density, $\rho_{0}$, which would occur in the absence of BEC, and define the correlation functions

$$
\begin{equation*}
R_{2}\left(p_{1}, p_{2}\right) \equiv \frac{\rho_{2}\left(p_{1}, p_{2}\right)}{\rho_{0}\left(p_{1}, p_{2}\right)}, \quad R_{3}\left(p_{1}, p_{2}, p_{3}\right) \equiv \frac{\rho_{3}\left(p_{1}, p_{2}, p_{3}\right)}{\rho_{0}\left(p_{1}, p_{2}, p_{3}\right)} \tag{2}
\end{equation*}
$$

Assuming the absence of two-particle correlations, i.e., $\rho_{2}\left(p_{1}, p_{2}\right)=\rho_{1}\left(p_{1}\right) \rho_{1}\left(p_{2}\right)$, results in

$$
\begin{equation*}
R_{3}^{\text {genuine }}\left(p_{1}, p_{2}, p_{3}\right) \equiv 1+\frac{C_{3}\left(p_{1}, p_{2}, p_{3}\right)}{\rho_{0}\left(p_{1}, p_{2}, p_{3}\right)} \tag{3}
\end{equation*}
$$

The kinematical variable $Q_{3}$ is used to study three-particle correlations. For a three-pion system, $Q_{3}=$ $\sqrt{M_{123}^{2}-9 m_{\pi}^{2}}$, with $M_{123}$ the invariant mass of the pion triplet and $m_{\pi}$ the mass of the pion. In this Letter, $\rho_{3}$ is defined as

$$
\begin{equation*}
\rho_{3}\left(Q_{3}\right) \equiv \frac{1}{N_{\mathrm{ev}}} \frac{\mathrm{~d} n_{\text {triplets }}}{\mathrm{d} Q_{3}} \tag{4}
\end{equation*}
$$

with $N_{\mathrm{ev}}$ the number of selected events and $n_{\text {triplets }}$ the number of triplets of like-sign tracks, and $\rho_{2}$ is defined analoguously.

Assuming totally incoherent production of particles and a source density $f(x)$ in spacetime with no dependence on the four-momentum of the emitted particle, the BE correlation functions is related to the source density by $[8$, 19]

$$
\begin{align*}
& R_{2}\left(Q_{i j}\right)=1+\left|F\left(Q_{i j}\right)\right|^{2}  \tag{5}\\
& R_{3}\left(Q_{12}, Q_{23}, Q_{31}\right)=1+\left|F\left(Q_{12}\right)\right|^{2}+\left|F\left(Q_{23}\right)\right|^{2}+\left|F\left(Q_{31}\right)\right|^{2}+2 \operatorname{Re}\left\{F\left(Q_{12}\right) F\left(Q_{23}\right) F\left(Q_{31}\right)\right\},  \tag{6}\\
& R_{3}^{\text {genuine }}\left(Q_{12}, Q_{23}, Q_{31}\right)=1+2 \operatorname{Re}\left\{F\left(Q_{12}\right) F\left(Q_{23}\right) F\left(Q_{31}\right)\right\}, \tag{7}
\end{align*}
$$

where $F\left(Q_{i j}\right)$ is the Fourier transform of $f(x)$.
$R_{2}$ does not depend on the phase $\phi_{i j}$ contained in $F\left(Q_{i j}\right) \equiv\left|F\left(Q_{i j}\right)\right| \exp \left(\iota \phi_{i j}\right)$. However, this phase survives in the three-particle BE correlation functions, Eqs. (6) and (7). Assuming fully incoherent particle production, the phase $\phi_{i j}$ can be non-zero only if the spacetime distribution of the source is asymmetric and $Q_{i j}>0$. Defining

$$
\begin{equation*}
\omega\left(Q_{12}, Q_{23}, Q_{31}\right)=\frac{R_{3}^{\text {genuine }}\left(Q_{12}, Q_{23}, Q_{31}\right)-1}{2 \sqrt{\left(R_{2}\left(Q_{12}\right)-1\right)\left(R_{2}\left(Q_{23}\right)-1\right)\left(R_{2}\left(Q_{31}\right)-1\right)}}, \tag{8}
\end{equation*}
$$

then for an incoherent source Eqs. (5) and (7) imply that $\omega=\cos \phi$, where $\phi \equiv \phi_{12}+\phi_{23}+\phi_{31}$. Furthermore, as $Q_{i j} \rightarrow 0$, then $\phi_{i j} \rightarrow 0$, and hence $\omega \rightarrow 1$. For $Q_{i j}>0$, a deviation from unity can be caused by an asymmetry
in the production. However, this will only result in a small (a few percent) reduction of $\omega$ [6,7], and this only in the case where the asymmetry occurs around the point of highest emissivity. It is important to emphasize that for (partially) coherent sources, $\omega$ can still be defined by Eq. (8), but Eqs. (5)-(7) are no longer valid, in which case more complicated expressions are needed [7], and one can no longer deduce that $\omega=\cos \phi$ or that $\omega \rightarrow 1$ as $Q_{i j} \rightarrow 0$. In at least one type of model, one can make the stronger statement that the limit $\omega=1$ at $Q_{i j} \rightarrow 0$ can only be reached if the source is fully incoherent [20].

## 4. Determination of $\boldsymbol{R}_{\mathbf{3}}$ and $\boldsymbol{R}_{\mathbf{3}}^{\text {genuine }}$

The reference sample, from which $\rho_{0}$ is determined, is formed by mixing particles from different data events in the following way. Firstly, 1000 events are rotated to a system with the $z$-axis along the thrust axis and are stored in a "pool". Then, tracks of each new event outside the pool are exchanged with tracks of the same charge from events in the pool having about the same (within about 20\%) multiplicity, under the condition that all tracks originate from different events. Thus, after this procedure the new event consists of tracks originating from different events in the pool, and its original tracks have entered the pool. This updating process prevents any regularities in the reference sample. Finally, $Q_{3}$ is calculated for each triplet of like-sign tracks, resulting in the density $\rho_{\text {mix }}$.

This mixing procedure removes more correlations than just those of BE, e.g., those from energy-momentum conservation and from resonances. This effect is estimated using a MC model with no BE effects (JETSET or HERWIG) at generator level and using pions only. Thus, in the absence of BEC, the corrected three-particle density is given by

$$
\begin{equation*}
\rho_{0}\left(Q_{3}\right)=\rho_{\text {mix }}\left(Q_{3}\right) \mathcal{C}_{\text {mix }}\left(Q_{3}\right), \quad \text { where } \mathcal{C}_{\text {mix }}\left(Q_{3}\right)=\left[\frac{\rho_{3}\left(Q_{3}\right)}{\rho_{\text {mix }}\left(Q_{3}\right)}\right]_{\mathrm{MC}, \text { noBE }} \tag{9}
\end{equation*}
$$

The density $\rho_{3}$, measured in data, must be corrected for detector resolution, acceptance, efficiency and for particle misidentification. For this we use a multiplicative factor, $\mathcal{C}_{\text {det }}$, derived from MC studies. Since no hadrons are identified in the analysis, $\mathcal{C}_{\text {det }}$ is given by the ratio of the three-pion correlation function found from MC events at generator level to the three-particle correlation function found using all particles after full detector simulation, reconstruction and selection. Combining this correction factor with Eqs. (2) and (9) results in

$$
\begin{equation*}
R_{3}\left(Q_{3}\right)=\frac{\rho_{3}\left(Q_{3}\right) \mathcal{C}_{\operatorname{det}}\left(Q_{3}\right)}{\rho_{\text {mix }}\left(Q_{3}\right) \mathcal{C}_{\text {mix }}\left(Q_{3}\right)} \tag{10}
\end{equation*}
$$

The genuine three-particle BE correlation function, $R_{3}^{\text {genuine }}$, is obtained via

$$
\begin{equation*}
R_{3}^{\text {genuine }}=R_{3}-R_{1,2}+1, \tag{11}
\end{equation*}
$$

where $R_{1,2} \equiv\left(\sum \rho_{1} \rho_{2}\right) / \rho_{0}-2$ is the contribution due to two-particle correlations, as may be seen from Eqs. (1) and (2). The product of densities $\sum \rho_{1}\left(p_{1}\right) \rho_{2}\left(p_{2}, p_{3}\right)$ is determined by a similar mixing procedure, as defined earlier, where two like-sign tracks from the same event are combined with one track having the same charge from another event with the same multiplicity. Finally, the variable $Q_{3}$ is calculated from these three tracks. This procedure is similar to that given in Ref. [21]. The ratio $\left(\sum \rho_{1} \rho_{2}\right) / \rho_{0}$ is also corrected for detector effects as $\rho_{3} / \rho_{\text {mix }}$.

In our analysis, we use JETSET without BEC and HERWIG to determine $\mathcal{C}_{\text {mix }}$ and JETSET with and without BEC as well as HERWIG to determine $\mathcal{C}_{\text {det }}$. These six MC combinations serve to estimate systematic uncertainties. The corrections are largest at small $Q_{3}$. At $Q_{3}=0.16 \mathrm{GeV}$, these corrections to $R_{3}$ are $\mathcal{C}_{\text {mix }} \approx 5-30 \%$ and $\mathcal{C}_{\text {det }} \approx 20-30 \%$, depending on which MC is used. These corrections for $R_{3}$ and $R_{1,2}$ are correlated and largely cancel in calculating $R_{3}^{\text {genuine }}$ by Eq. (11).

To correct the data for two-pion Coulomb repulsion in calculating $\rho_{2}$, each pair of pions is weighted by the inverse Gamow factor [22]

$$
\begin{equation*}
G_{2}^{-1}\left(\eta_{i j}\right)=\frac{\exp \left(2 \pi \eta_{i j}\right)-1}{2 \pi \eta_{i j}}, \quad \text { where } \eta_{i j}=\frac{m_{\pi} \alpha}{Q_{i j}} \tag{12}
\end{equation*}
$$

and $\alpha$ is the fine-structure constant. It has been shown [23] that this Gamow factor is an approximation suitable for our purposes. For $\rho_{3}$, the weight of each triplet is taken as the product of the weights of the three pairs within it. For $\sum \rho_{2} \rho_{1}$ we use the same weight but with $G_{2}\left(Q_{i j}\right) \equiv 1$ when particles $i$ and $j$ come from different events. At the lowest $Q_{3}$ values under consideration, the Coulomb correction is approximately $10 \%, 3 \%$ and $2 \%$, for $\rho_{3}$, $\sum \rho_{1} \rho_{2}$ and $\rho_{2}$, respectively.

## 5. Results

The measurements of $R_{3}, R_{1,2}$ and $R_{2}$ are shown in Fig. 2. The full circles correspond to the averages of the data points obtained from the six possible MC combinations used to determine $\mathcal{C}_{\text {mix }}$ and $\mathcal{C}_{\text {det }}$. The error bars, $\sigma_{1}$, include both the statistical uncertainty and the systematic uncertainty of the MC modeling, which is taken as the r.m.s. of the values obtained using the different MC combinations. This dominant systematic uncertainty is, for $Q_{3}<0.8 \mathrm{GeV}$, about a factor 5 to 7 larger than the statistical uncertainty and is correlated between the $R_{3}, R_{1,2}$ and $R_{2}$ distributions of Fig. 2 and between bins. Fig. 2(a) shows the existence of three-particle correlations and from Fig. 2(b) it is clear that about half is due to two-particle correlations. Fig. 2(c) shows the two-particle correlations.

As a check, $R_{3}, R_{1,2}$ and $R_{2}$ are also computed for MC models without BEC, both HERWIG and JETSET, after detector simulation, reconstruction and selection. For the mixing and detector corrections all possible MC combinations, giving non-trivial results, are studied. The results of this check are shown in Fig. 2 as open circles and, as expected, flat distributions around unity are observed.

Fig. 3(a) shows the genuine three-particle BE correlation function $R_{3}^{\text {genuine }}$. The data points show the existence of genuine three-particle BE correlations. The MC systematic uncertainty is highly correlated from bin to bin. At $Q_{3}<0.8 \mathrm{GeV}$, it is about a factor 1.5 to 3.5 larger than the statistical uncertainty, the higher value corresponding to the lowest $Q_{3}$ value used. The open circles correspond to MC without BEC and form a flat distribution around unity, as expected.

### 5.1. Gaussian parametrizations

A fit from $Q_{3}=0.16$ to 1.40 GeV using the covariance matrix including both the statistical uncertainty and the systematic uncertainty due to the MC modeling, $\sigma_{1}$, is performed on the data points with the commonly used [8, 10,11,21] parametrization

$$
\begin{equation*}
R_{3}^{\text {genuine }}\left(Q_{3}\right)=\tilde{\gamma}\left[1+2 \tilde{\lambda}^{1.5} \exp \left(-\widetilde{R}^{2} Q_{3}^{2} / 2\right)\right]\left(1+\tilde{\varepsilon} Q_{3}\right) \tag{13}
\end{equation*}
$$

where $\tilde{\gamma}$ is an overall normalization factor, $\tilde{\lambda}$ measures the strength of the correlation, $\widetilde{R}$ is a measure for the effective source size in spacetime and the term $\left(1+\tilde{\varepsilon} Q_{3}\right)$ takes into account possible long-range momentum correlations. The form of this parametrization is a consequence of the assumptions that $\omega=1$ and that $\left|F\left(Q_{i j}\right)\right|=$ $\sqrt{\lambda} \exp \left(-\widetilde{R}^{2} Q_{i j}^{2} / 2\right)$, as would be expected for a Gaussian source density. The fit results are given in the first column of Table 1 and shown as the full line in Fig. 3(a).

In addition to the MC modeling, we investigate four other sources of systematic uncertainties on the fit parameters. Firstly, the influence of a different mixing sample is studied by removing the conditions that tracks are replaced by tracks with the same charge and coming from events with approximately the same multiplicity. For each of the six MC combinations, the difference in the fit results between the two mixing methods is taken as an


Fig. 2. (a) The three-particle BE correlation function, $R_{3}$, from Eq. (10), (b) the contribution of two-particle correlations, $R_{1,2} \equiv\left(\sum \rho_{2} \rho_{1}\right) / \rho_{0}-2$, and (c) $R_{2}$ from Eq. (5). The full circles correspond to the data and the error bars to $\sigma_{1}$ (see text). The open circles correspond to the results from MC models without BEC. In (c) the dashed and full lines show the fits of Eqs. (14) and (15), respectively.
estimate of the systematic uncertainty. The square root of the mean of the squares of these differences is assigned as the systematic uncertainty from this source. In the same way, systematic uncertainties related to track and event selection and to the choice of the fit range are evaluated. The analysis is repeated with stronger and weaker selection criteria, changing the number of events by about $\pm 11 \%$ and the number of tracks by about $\pm 12 \%$. The fit range is varied by removing the first point of the fit and varying the end point by $\pm 200 \mathrm{MeV}$. Finally, we study the influence of removing like-sign track pairs with small polar and azimuthal opening angles. The maximum deviation that is found by varying the cuts on these angles up to $6^{\circ}$, is taken as the systematic uncertainty from this source. The total systematic uncertainty due to these four sources is obtained by adding the four uncertainties in quadrature. We refer to this systematic uncertainty as $\sigma_{2}$. For all fit parameters, the largest part of the total uncertainty is due to the six possible combinations of mixing and detector MC corrections and amounts to $50-90 \%$. Table 2 shows the uncertainties for each of the sources for the fit parameters of Eq. (13).

As a cross-check, the analysis is repeated without the momentum cut of 1 GeV and without the cut of $3^{\circ}$ on the opening angle of like-sign track pairs. The results agree with those given in Table 1 well within quoted uncertainties, but the systematic uncertainties are approximately twice as large.


Fig. 3. The genuine three-particle BE correlation function $R_{3}^{\text {genuine }}$, Eq. (11). The full circles correspond to the data and the error bars to $\sigma_{1}$. The open circles correspond to results from MC models without BEC. In (a) the full line shows the fit of Eq. (13), the dashed line the prediction of completely incoherent pion production and a Gaussian source density in spacetime, derived from parametrizing $R_{2}$ with Eq. (14). In (b) Eqs. (16) and (15) are used, respectively.

Table 1
Values of the fit parameters for the genuine three-particle BE correlation function $R_{3}^{\text {genuine }}$, using the parametrizations of Eqs. (13) and (16). The first uncertainty corresponds to $\sigma_{1}$, the second to $\sigma_{2}$, defined in the text

| Parameter | Eq. (13) | Eq. (16) |
| :--- | :---: | :---: |
| $\tilde{\gamma}$ | $0.96 \pm 0.03 \pm 0.02$ | $0.95 \pm 0.03 \pm 0.02$ |
| $\tilde{\lambda}$ | $0.47 \pm 0.07 \pm 0.03$ | $0.75 \pm 0.10 \pm 0.03$ |
| $\widetilde{R}, \mathrm{fm}^{2}$ | $0.65 \pm 0.06 \pm 0.03$ | $0.72 \pm 0.08 \pm 0.03$ |
| $\tilde{\varepsilon}, \mathrm{GeV}^{-1}$ | $0.02 \pm 0.02 \pm 0.02$ | $0.02 \pm 0.02 \pm 0.02$ |
| $\tilde{\kappa}$ | - | $0.79 \pm 0.26 \pm 0.15$ |
| $\chi^{2} / \mathrm{NDF}$ | $29.9 / 27$ | $17.7 / 26$ |

To measure the ratio $\omega$, we also need to determine the two-particle BE correlation function $R_{2}(Q)$. This is done in the same way as the three-particle BE correlation function. The correlation function $R_{2}$ is parametrized as a Gaussian:

$$
\begin{equation*}
R_{2}(Q)=\gamma\left[1+\lambda \exp \left(-R^{2} Q^{2}\right)\right](1+\varepsilon Q) \tag{14}
\end{equation*}
$$

Table 2
Contribution to the uncertainty on the fit parameters of the parametrizations of Eqs. (13) and (16), respectively. The first uncertainty corresponds to $\sigma_{1}$, the others added in quadrature give $\sigma_{2}$

| Parametrization | Eq. (13) |  |  |  | Eq. (16) |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Fit parameter | $\tilde{\gamma}$ | $\tilde{\lambda}$ | $\widetilde{R}, \mathrm{fm}$ | $\tilde{\varepsilon}, \mathrm{GeV}^{-1}$ | $\tilde{\gamma}$ | $\tilde{\lambda}$ | $\widetilde{R}, \mathrm{fm}$ | $\tilde{\varepsilon}, \mathrm{GeV}^{-1}$ | $\tilde{\kappa}$ |
| $\sigma_{1}$ (stat. + modeling) | 0.029 | 0.071 | 0.056 | 0.022 | 0.031 | 0.103 | 0.078 | 0.024 | 0.26 |
| Mixing | 0.004 | 0.006 | 0.009 | 0.007 | 0.010 | 0.009 | 0.011 | 0.011 | 0.04 |
| Fit range | 0.008 | 0.019 | 0.020 | 0.013 | 0.010 | 0.022 | 0.017 | 0.019 | 0.14 |
| Track/event sel. | 0.010 | 0.013 | 0.012 | 0.008 | 0.011 | 0.016 | 0.012 | 0.007 | 0.10 |
| $\delta \phi+\delta \theta$ cut | 0.013 | 0.014 | 0.012 | 0.009 | 0.014 | 0.017 | 0.010 | 0.008 | 0.11 |
| $\sigma_{2}$ | 0.019 | 0.028 | 0.028 | 0.020 | 0.023 | 0.033 | 0.026 | 0.024 | 0.15 |



Fig. 4. The ratio $\omega$ as a function of $Q_{3}$ assuming $R_{2}$ is described (a) by the Gaussian, Eq. (14), and (b) by the first-order Edgeworth expansion of the Gaussian, Eq. (15). The error bars correspond to $\sigma_{1}$. For completely incoherent production, $\omega=1$.

The parametrization starts at $Q=0.08 \mathrm{GeV}$, consistent with the study of $R_{3}$ from $Q_{3}=0.16 \mathrm{GeV}$. The fit results ${ }^{8}$ are given in the first column of Table 3 and in Fig. 2(c).

If the spacetime structure of the pion source is Gaussian and the pion production mechanism is completely incoherent, $\tilde{\lambda}$ and $\widetilde{R}$ as derived from the fit to Eq. (13) measure the same correlation strength and effective source size as $\lambda$ and $R$ of Eq. (14). The values of $\lambda$ and $R$ are consistent with $\tilde{\lambda}$ and $\widetilde{R}$, as expected for fully incoherent production of pions ( $\omega=1$ ). Using the values of $\lambda$ and $R$ instead of $\tilde{\lambda}$ and $\widetilde{R}$ in Eq. (13), which is justified if $\omega=1$, results in the dashed line in Fig. 3(a). It is only slightly different from the result of the fit to Eq. (13), indicating that $\omega$ is indeed near unity.

Another way to see how well $R_{3}^{\text {genuine }}$ corresponds to a completely incoherent pion production interpretation and a Gaussian source density in spacetime, is to compute $\omega$ with Eq. (8), for each bin in $Q_{3}$ (from 0.16 to 0.80 GeV ), using the measured $R_{3}^{\text {genuine }}$ and $R_{2}$ derived from the parametrization of Eq. (14). The result is shown in Fig. 4(a). At low $Q_{3}, \omega$ appears to be higher than unity.

[^3]Table 3
Values of the fit parameters for the two-particle BE correlation function, $R_{2}$, using the parametrizations of Eqs. (14) and (15). The first uncertainty corresponds to $\sigma_{1}$, the second to $\sigma_{2}$

| Parameter | Eq. (14) | Eq. (15) |
| :--- | :---: | ---: |
| $\gamma$ | $0.98 \pm 0.03 \pm 0.02$ | $0.96 \pm 0.03 \pm 0.02$ |
| $\lambda$ | $0.45 \pm 0.06 \pm 0.03$ | $0.72 \pm 0.08 \pm 0.03$ |
| $R, \mathrm{fm}$ | $0.65 \pm 0.03 \pm 0.03$ | $0.74 \pm 0.06 \pm 0.02$ |
| $\varepsilon, \mathrm{GeV}^{-1}$ | $0.01 \pm 0.01 \pm 0.02$ | $0.01 \pm 0.02 \pm 0.02$ |
| $\kappa$ | - | $0.74 \pm 0.21 \pm 0.15$ |
| $\chi^{2} / \mathrm{NDF}$ | $60.2 / 29$ | $26.0 / 28$ |

### 5.2. Extended Gaussian parametrizations

However, the assumption of a Gaussian source density is only an approximation, as observed in Ref. [2] and confirmed by the $\chi^{2}$ of the fit to Eq. (14). Deviations from a Gaussian can be studied by expanding in terms of derivatives of the Gaussian, which are related to Hermite polynomials. Taking only the lowest-order non-Gaussian term into account, this so-called Edgeworth expansion [25] replaces the parametrization of Eq. (14) by

$$
\begin{equation*}
R_{2}(Q)=\gamma\left[1+\lambda \exp \left(-R^{2} Q^{2}\right)\left(1+\kappa H_{3}(\sqrt{2} R Q) / 6\right)\right](1+\varepsilon Q), \tag{15}
\end{equation*}
$$

where $\kappa$ measures the deviation from the Gaussian and $H_{3}(x) \equiv x^{3}-3 x$ is the third-order Hermite polynomial. The fit results for the two-particle BE correlation function with this parametrization are given in the second column of Table 3.

Using the first-order Edgeworth expansion of the Gaussian, Eq. (15), and using Eq. (8), assuming $\omega=1$, the parametrization of Eq. (13) becomes

$$
\begin{align*}
R_{3}^{\text {genuine }}\left(Q_{3}\right) & =\tilde{\gamma}\left(1+2 \tilde{\lambda}^{1.5} \exp \left(-\widetilde{R}^{2} Q_{3}^{2} / 2\right)\left[\prod_{i, j=1, j>i}^{3} \sqrt{1+\frac{H_{3}\left(\sqrt{2} \widetilde{R} Q_{i j}\right)}{6} \tilde{\kappa}}\right]\right)\left(1+\tilde{\varepsilon} Q_{3}\right) \\
& \simeq \tilde{\gamma}\left(1+2 \tilde{\lambda}^{1.5} \exp \left(-\widetilde{R}^{2} Q_{3}^{2} / 2\right)\left[1+\frac{H_{3}\left(\sqrt{2} \widetilde{R} Q_{3} / 2\right)}{6} \tilde{\kappa}\right]^{1.5}\right)\left(1+\tilde{\varepsilon} Q_{3}\right) \tag{16}
\end{align*}
$$

In the second line the approximation is made that $Q_{i j}=Q_{3} / 2$. The effect of this approximation on $R_{3}^{\text {genuine }}$ is small compared to the statistical uncertainty. The results of a fit to Eq. (16) are given in the second column of Table 1. The uncertainties are summarized in Table 2.

For both $R_{3}^{\text {genuine }}$ and $R_{2}$, a better $\chi^{2} / \mathrm{NDF}$ is found using the Edgeworth expansion, and the values of $\tilde{\lambda}$ and $\lambda$ are significantly higher, as shown in Tables 1 and 3 and in Figs. 3(b) and 2(c). The values for $\tilde{\lambda}$ and $\widetilde{R}$ are still consistent with the corresponding $\lambda$ and $R$, as would be expected for a fully incoherent production mechanism of pions.

In Fig. 3(b), as in Fig. 3(a), we observe good agreement between the fit of $R_{3}^{\text {genuine }}$ using the parametrization of Eq. (16) and the prediction of a completely incoherent pion production mechanism, derived from parametrizing $R_{2}$ with Eq. (15), over the full range of $Q_{3}$. In Fig. 4(b), no deviation from unity is observed for the ratio $\omega$. This indicates that the data agree with the assumption of fully incoherent pion production.

Fits to samples generated with JETSET with BE effects modelled by $\mathrm{BE}_{0}$ or $\mathrm{BE}_{32}{ }^{9}$ [16] result in values of $\widetilde{R}$ in agreement with the data but in significantly higher values of $\tilde{\lambda}$. This confirms the observation in Fig. 1(f) that the standard BE implementations of JETSET overestimate the genuine three-particle BEC.

[^4]
## 6. Summary

Three-particle, as well as two-particle Bose-Einstein correlations of like-sign charged pions have been measured in hadronic Z decay. Genuine three-particle BE correlations are observed. The correlation functions are better parametrized by an Edgeworth expansion of a Gaussian than by a simple Gaussian. Combining the two- and threeparticle correlations shows that the data are consistent with a fully incoherent production mechanism of pions.

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## References

[1] DELPHI Collaboration, P. Abreu et al., Phys. Lett. B 286 (1992) 201; DELPHI Collaboration, P. Abreu et al., Z. Phys. C 63 (1994) 17; ALEPH Collaboration, D. Decamp et al., Z. Phys. C 54 (1992) 75; OPAL Collaboration, G. Alexander et al., Z. Phys. C 72 (1996) 389.
[2] L3 Collaboration, M. Acciarri et al., Phys. Lett. B 458 (1999) 517.
[3] OPAL Collaboration, G. Abbiendi et al., Eur. Phys. J. C 16 (2000) 423; DELPHI Collaboration, P. Abreu et al., Phys. Lett. B 471 (2000) 460.
[4] G. Goldhaber et al., Phys. Rev. 120 (1960) 300; D.H. Boal, C.K. Gelbke, B.K. Jennings, Rev. Mod. Phys. 62 (1990) 553.
[5] M. Biyajima et al., Progr. Theor. Phys. 84 (1990) 931.
[6] H. Heiselberg, A.P. Vischer, Phys. Rev. C 55 (1997) 874, nucl-th/9707036.
[7] U. Heinz, Q. Zhang, Phys. Rev. C 56 (1997) 426.
[8] B. Lörstad, Int. J. Mod. Phys. A 4 (1989) 2861.
[9] I.V. Andreev, M. Plümer, R.M. Weiner, Int. J. Mod. Phys. A 8 (1993) 4577.
[10] DELPHI Collaboration, P. Abreu et al., Phys. Lett. B 355 (1995) 415.
[11] OPAL Collaboration, K. Ackerstaff et al., Eur. Phys. J. C 5 (1998) 239.
[12] L3 Collaboration, B. Adeva et al., Nucl. Instrum. Methods A 289 (1990) 35; J.A. Bakken et al., Nucl. Instrum. Methods A 275 (1989) 81;
O. Adriani et al., Nucl. Instrum. Methods A 302 (1991) 53;
O. Adriani et al., Phys. Rep. 236 (1993) 1;
B. Adeva et al., Nucl. Instrum. Methods A 323 (1992) 109;
K. Deiters et al., Nucl. Instrum. Methods A 323 (1992) 162;
M. Chemarin et al., Nucl. Instrum. Methods A 349 (1994) 345;
M. Acciarri et al., Nucl. Instrum. Methods A 351 (1994) 300;
I.C. Brock et al., Nucl. Instrum. Methods A 381 (1996) 23;
A. Adam et al., Nucl. Instrum. Methods A 383 (1996) 342;
G. Basti et al., Nucl. Instrum. Methods A 374 (1996) 293.
[13] T. Sjöstrand, Comput. Phys. Commun. 82 (1994) 74.
[14] G. Marchesini, B.R. Webber, Nucl. Phys. B 310 (1988) 461;
G. Marchesini et al., Comput. Phys. Commun. 67 (1992) 465.
[15] L. Lönnblad, T. Sjöstrand, Phys. Lett. B 351 (1995) 293.
[16] L. Lönnblad, T. Sjöstrand, Eur. Phys. J. C 2 (1998) 165.
[17] R. Brun et al., CERN report CERN DD/EE/84-1 (1984); revised 1987.
[18] H. Fesefeldt, RWTH Aachen report PITHA 85/02 (1985).
[19] V.L. Lyuboshitz, Sov. J. Nucl. Phys. 53 (1991) 514.
[20] T. Csörgő et al., Eur. Phys. J. C 9 (1999) 275.
[21] NA22 Collaboration, N.M. Agababyan et al., Z. Phys. C 68 (1995) 229.
[22] M. Gyulassy, S. Kauffmann, L.W. Wilson, Phys. Rev. C 20 (1979) 2267.
[23] E.O. Alt et al., Eur. Phys. J. C 13 (2000) 663.
[24] L3 Collaboration, P. Achard et al., Phys. Lett. B 524 (2002) 55.
[25] F.Y. Edgeworth, Trans. Cambridge Phil. Soc. 20 (1905) 36;
See also, e.g., H. Cramér, Mathematical Methods of Statistics, Princeton Univ. Press, Princeton, NJ, 1946;
T. Csörgő, S. Hegyi, Phys. Lett. B 489 (2000) 15.


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    ${ }^{6}$ Supported by the National Natural Science Foundation of China.
    ${ }^{7}$ The Bose-Einstein simulation is done by the subroutine LUBOEI, with the values $\operatorname{PARJ}(92)=1.5$ and $\operatorname{PARJ}(93)=0.33 \mathrm{GeV}$.

[^3]:    ${ }^{8}$ Due to the use of a different fit range, these fit results differ from those found in Ref. [24]. The same fit range gives similar results.

[^4]:    ${ }^{9}$ The $\mathrm{BE}_{32}$ algorithm uses the values $\operatorname{PARJ}(92)=1.68$ and $\operatorname{PARJ}(93)=0.38 \mathrm{GeV}$.

