

Search for diphoton events with large missing transverse momentum in 7 TeV proton–proton collision data with the ATLAS detector [☆]

ATLAS Collaboration ^{*}

ARTICLE INFO

Article history:

Received 4 September 2012

Received in revised form 11 October 2012

Accepted 24 October 2012

Available online 29 October 2012

Editor: H. Weerts

ABSTRACT

A search for diphoton events with large missing transverse momentum has been performed using proton–proton collision data at $\sqrt{s} = 7$ TeV recorded with the ATLAS detector, corresponding to an integrated luminosity of 4.8 fb^{-1} . No excess of events was observed above the Standard Model prediction and model-dependent 95% confidence level exclusion limits are set. In the context of a generalised model of gauge-mediated supersymmetry breaking with a bino-like lightest neutralino of mass above 50 GeV, gluinos (squarks) below 1.07 TeV (0.87 TeV) are excluded, while a breaking scale Λ below 196 TeV is excluded for a minimal model of gauge-mediated supersymmetry breaking. For a specific model with one universal extra dimension, compactification scales $1/R < 1.40$ TeV are excluded. These limits provide the most stringent tests of these models to date.

© 2012 CERN. Published by Elsevier B.V. Open access under CC BY license.

1. Introduction

This Letter reports on a search for diphoton ($\gamma\gamma$) events with large missing transverse momentum (E_T^{miss}) in 4.8 fb^{-1} of proton–proton (pp) collision data at $\sqrt{s} = 7$ TeV recorded with the ATLAS detector at the Large Hadron Collider (LHC) in 2011, extending and superseding a prior study performed with 1 fb^{-1} [1]. The results are interpreted in the context of three models of new physics: a general model of gauge-mediated supersymmetry breaking (GGM) [2–4], a minimal model of gauge-mediated supersymmetry breaking (SPS8) [5], and a model with one universal extra dimension (UED) [6–8].

2. Supersymmetry

Supersymmetry (SUSY) [9–17] introduces a symmetry between fermions and bosons, resulting in a SUSY partner (sparticle) with identical quantum numbers except a difference by half a unit of spin for each Standard Model (SM) particle. As none of these sparticles have been observed, SUSY must be a broken symmetry if realised in nature. Assuming R -parity conservation [18–22], sparticles are produced in pairs. These would then decay through cascades involving other sparticles until the lightest SUSY particle (LSP), which is stable, is produced.

In gauge-mediated SUSY breaking (GMSB) models [23–28] the LSP is the gravitino \tilde{G} . GMSB experimental signatures are largely determined by the nature of the next-to-lightest SUSY particle (NLSP). In this study the NLSP is assumed to be the lightest

neutralino $\tilde{\chi}_1^0$. Should the lightest neutralino be a bino (the SUSY partner of the SM U(1) gauge boson), the final decay in the cascade would predominantly be $\tilde{\chi}_1^0 \rightarrow \gamma \tilde{G}$, with two cascades per event, leading to final states with $\gamma\gamma + E_T^{\text{miss}}$, where E_T^{miss} results from the undetected gravitinos.

Two different classes of gauge-mediated models, described in more detail below, are considered as benchmarks to evaluate the reach of this analysis: the minimal GMSB model (SPS8) as an example of a complete SUSY model with a full particle spectrum and two different variants of the GGM model as examples of phenomenological models with reduced particle content.

In the SPS8 model, the only free parameter is the SUSY-breaking mass scale Λ that establishes the nature of the observable phenomena exhibited by the low-energy sector. The other model parameters are fixed to the following values: the messenger mass $M_{\text{mess}} = 2\Lambda$, the number of SU(5) messengers $N_5 = 1$, the ratio of the vacuum expectation values of the two Higgs doublets $\tan\beta = 15$, and the Higgs sector mixing parameter $\mu > 0$. The bino NLSP is assumed to decay promptly ($c\tau_{\text{NLSP}} < 0.1$ mm). For $\Lambda \simeq 200$ TeV, the direct production of gaugino pairs such as $\tilde{\chi}_2^0 \tilde{\chi}_1^\pm$ or $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ pairs is expected to dominate at a LHC centre-of-mass energy of $\sqrt{s} = 7$ TeV. The contribution from gluino and/or squark pairs is below 10% of the production cross section due to their high masses. The sparticle pair produced in the collision decays via cascades into two photons and two gravitinos. Further SM particles such as gluons, quarks, leptons and gauge bosons may be produced in the cascade decays. The current best limit on Λ in this model is 145 TeV [1].

Two different configurations of the GGM SUSY model are considered in this study, for which the neutralino NLSP, chosen to be the bino, and either the gluino or the squark masses are treated as

[☆] © CERN for the benefit of the ATLAS Collaboration.

^{*} E-mail address: atlas.publications@cern.ch.

free parameters. For the squark–bino GGM model all squark masses are treated as degenerate except the right-handed up-type squarks whose mass is decoupled (set to inaccessible large values). For the gluino–bino model all squark masses are decoupled. For both configurations all other sparticle masses are also decoupled, leading to a dominant production mode at $\sqrt{s} = 7$ TeV of a pair of squarks in one case and a pair of gluinos in the other case. These would decay via short cascades into the bino-like neutralino NLSP. Jets may be produced in the cascades from the gluino and squark decays. Further model parameters are fixed to $c\tau_{\text{NLSP}} < 0.1$ mm and $\tan\beta = 2$; for this GGM scenario, restricted to the region of parameter space for which the NLSP is the bino-like neutralino, the final-state phenomenology relevant to this search is only weakly dependent on the value of $\tan\beta$ [4]. The decay into the wino-like neutralino NLSP is possible and was studied by the CMS Collaboration [29].

3. Extra dimensions

UED models postulate the existence of additional spatial dimensions in which all SM particles can propagate, leading to the existence of a series of excitations for each SM particle, known as a Kaluza–Klein (KK) tower. This analysis considers the case of a single UED, with compactification radius (size of the extra dimension) $R \approx 1 \text{ TeV}^{-1}$. At the LHC, the main UED process would be the production via the strong interaction of a pair of first-excitation-level KK quarks and/or gluons [30]. These would decay via cascades involving other KK particles until reaching the lightest KK particle (LKP), i.e. the first-excitation-level KK photon γ^* . SM particles such as quarks, gluons, leptons and gauge bosons may be produced in the cascades. If the UED model is embedded in a larger space with N additional eV $^{-1}$ -sized dimensions accessible only to gravity [31], with a $(4 + N)$ -dimensional Planck scale (M_D) of a few TeV, the LKP would decay gravitationally via $\gamma^* \rightarrow \gamma + G$. G represents a tower of eV-spaced graviton states, leading to a graviton mass between 0 and $1/R$. With two decay chains per event, the final state would contain $\gamma\gamma + E_T^{\text{miss}}$, where E_T^{miss} results from the escaping gravitons. Up to $1/R \sim 1$ TeV, the branching ratio to the diphoton and E_T^{miss} final state is close to 100%. As $1/R$ increases, the gravitational decay widths become more important for all KK particles and the branching ratio into photons decreases, e.g. to 50% for $1/R = 1.5$ TeV [7].

The UED model considered here is defined by specifying R and Λ , the ultraviolet cut-off used in the calculation of radiative corrections to the KK masses. This analysis sets Λ such that $\Lambda R = 20$ [32]. The γ^* mass is insensitive to Λ , while other KK masses typically change by a few per cent when varying ΛR in the range 10–30. For $1/R = 1.4$ TeV, the masses of the first-excitation-level KK photon, quark and gluon are 1.40 TeV, 1.62 TeV and 1.71 TeV, respectively [33].

4. Simulated samples

For the GGM model, the SUSY mass spectra were calculated using **SUSPECT** 2.41 [34] and **SDECAY** 1.3 [35]; for the SPS8 model, the SUSY mass spectra were calculated using **ISAJET** 7.80 [36]. The Monte Carlo (MC) SUSY signal samples were produced using **Herwig++** 2.5.1 [37] with **MRST2007 LO*** [38] parton distribution functions (PDFs). Signal cross sections were calculated to next-to-leading order (NLO) in the strong coupling constant, including the resummation of soft gluon emission at next-to-leading-logarithmic accuracy [39–43]. The nominal cross sections and the uncertainties were taken from an envelope of cross-section predictions using different PDF sets and factorisation and renormalisation scales, as described in Ref. [44]. In the case of the UED model, cross

sections were estimated and MC signal samples generated using the UED model as implemented at leading order (LO) in **PYTHIA** 6.423 [45,33] with **MRST2007 LO*** PDFs.

The “irreducible” background from $W(\rightarrow \ell\nu) + \gamma\gamma$ and $Z(\rightarrow \nu\bar{\nu}) + \gamma\gamma$ production was simulated at LO using **MadGraph 4** [46] with the **CTEQ6L1** [47] PDFs. Parton showering and fragmentation were simulated with **PYTHIA**. NLO cross sections and scale uncertainties were implemented via multiplicative constants (K -factors) that relate the NLO and LO cross sections. These have been calculated for several restricted regions of the overall phase space of the $Z(\rightarrow \nu\bar{\nu}) + \gamma\gamma$ and $W(\rightarrow \ell\nu) + \gamma\gamma$ processes [48,49], and are estimated to be 2.0 ± 0.3 and 3 ± 3 for the $Z(\rightarrow \nu\bar{\nu}) + \gamma\gamma$ and $W(\rightarrow \ell\nu) + \gamma\gamma$ contributions to the signal regions of this analysis, respectively. As described below, all other background sources are estimated through the use of control samples derived from data.

All samples were processed through the **GEANT4**-based simulation of the ATLAS detector [50,51]. The variation of the number of pp interactions per bunch crossing (pile-up) as a function of the instantaneous luminosity is taken into account by overlaying simulated minimum bias events according to the observed distribution of the number of pile-up interactions in data, with an average of ~ 10 interactions.

5. ATLAS detector

The ATLAS detector [52] is a multi-purpose apparatus with a forward-backward symmetric cylindrical geometry and nearly 4π solid angle coverage. Closest to the beamline are tracking devices comprising layers of silicon-based pixel and strip detectors covering $|\eta| < 2.5^1$ and straw-tube detectors covering $|\eta| < 2.0$, located inside a thin superconducting solenoid that provides a 2 T magnetic field. Outside the solenoid, fine-granularity lead/liquid-argon electromagnetic (EM) calorimeters provide coverage for $|\eta| < 3.2$ to measure the energy and position of electrons and photons. A presampler, covering $|\eta| < 1.8$, is used to correct for energy lost upstream of the EM calorimeter. An iron/scintillating-tile hadronic calorimeter covers the region $|\eta| < 1.7$, while a copper/liquid-argon medium is used for hadronic calorimeters in the end-cap region $1.5 < |\eta| < 3.2$. In the forward region $3.2 < |\eta| < 4.9$ liquid-argon calorimeters with copper and tungsten absorbers measure the electromagnetic and hadronic energy. A muon spectrometer consisting of three superconducting toroidal magnet systems each comprising eight toroidal coils, tracking chambers, and detectors for triggering surrounds the calorimeter system.

6. Reconstruction of candidates and observables

The reconstruction of converted and unconverted photons and of electrons is described in Refs. [53] and [54], respectively. Photon candidates were required to be within $|\eta| < 1.81$, and to be outside the transition region $1.37 < |\eta| < 1.52$ between the barrel and end-cap calorimeters. Identified on the basis of the characteristics of the longitudinal and transverse shower development in the EM calorimeter, the analysis made use of both “loose” and “tight” photons [53]. In the case that an EM calorimeter deposition was identified as both a photon and an electron, the photon candidate was discarded and the electron candidate retained. In addition,

¹ ATLAS uses a right-handed coordinate system with its origin at the nominal interaction point (IP) in the centre of the detector and the z -axis along the beam pipe. The x -axis points from the IP to the centre of the LHC ring, and the y -axis points upward. Cylindrical coordinates (r, ϕ) are used in the transverse plane, ϕ being the azimuthal angle around the beam pipe. The pseudorapidity is defined in terms of the polar angle θ as $\eta = -\ln \tan(\theta/2)$.

converted photons were re-classified as electrons if one or more candidate conversion tracks included at least one hit from the pixel layers. Giving preference to the electron selection in this way reduced the electron-to-photon fake rate by 50–60% (depending on the value of η) relative to that of the prior 1 fb^{-1} analysis [1], while preserving over 70% of the signal efficiency. Finally, an “isolation” requirement was imposed. After correcting for contributions from pile-up and the deposition ascribed to the photon itself, photon candidates were removed if more than 5 GeV of transverse energy was observed in a cone of $\sqrt{(\Delta\eta)^2 + (\Delta\phi)^2} < 0.2$ surrounding the energy deposition in the calorimeter associated with the photon.

The measurement of the two-dimensional transverse momentum vector $\mathbf{p}_T^{\text{miss}}$ (and its magnitude E_T^{miss}) was based on energy deposits in calorimeter cells inside three-dimensional clusters with $|\eta| < 4.9$ and was corrected for contributions from muons, if any [55]. The cluster energy was calibrated to correct for the different response to electromagnetically- and hadronically-induced showers, energy loss in dead material, and out-of-cluster energy. The contribution from identified muons was accounted for by adding in the energy derived from the properties of reconstructed muon tracks.

Jets were reconstructed using the anti- k_t jet algorithm [56] with radius parameter $R = 0.4$. They were required to have $p_T > 20 \text{ GeV}$ and $|\eta| < 2.8$ [57].

Two additional observables of use in discriminating SM backgrounds from potential GMSB and UED signals were defined. The total visible transverse energy H_T was calculated as the sum of the magnitude of the transverse momenta of the two selected photons and any additional leptons and jets in the event. The photon- E_T^{miss} separation $\Delta\phi(\gamma, E_T^{\text{miss}})$ was defined as the azimuthal angle between the missing transverse momentum vector and either of the two selected photons, with $\Delta\phi_{\min}(\gamma, E_T^{\text{miss}})$ the minimum value of $\Delta\phi(\gamma, E_T^{\text{miss}})$ of the two selected photons.

7. Data analysis

The data sample, corresponding to an integrated luminosity of $(4.8 \pm 0.2) \text{ fb}^{-1}$ [58,59], was selected by a trigger requiring two loose photon candidates with $E_T > 20 \text{ GeV}$. To ensure the event resulted from a beam collision, events were required to have at least one vertex with five or more associated tracks. Events were then required to contain at least two tight photon candidates with $E_T > 50 \text{ GeV}$, which MC studies suggested would provide the greatest separation between signal and SM background for a broad range of the parameter space of the new physics scenarios under consideration in this search. A total of 10455 isolated $\gamma\gamma$ candidate events passing these selection requirements were observed in the data sample. The E_T distributions² of the leading and sub-leading photon for events in this sample are shown in Figs. 1 and 2. Also shown are the E_T spectra obtained from GGM MC samples for $m_{\tilde{g}} = 1000 \text{ GeV}$ and $m_{\tilde{\chi}_1^0} = 450 \text{ GeV}$, from SPS8 MC samples with $\Lambda = 190 \text{ TeV}$, and from UED MC samples for $1/R = 1.3 \text{ TeV}$, representing model parameters near the expected exclusion limit. Figs. 3 and 4 show the H_T and $\Delta\phi_{\min}(\gamma, E_T^{\text{miss}})$

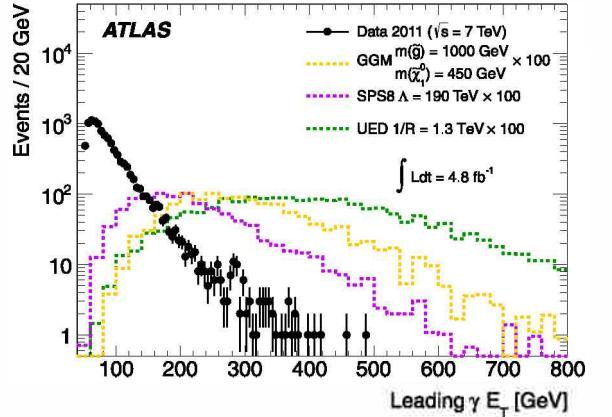


Fig. 1. The E_T spectrum of the leading photon in the $\gamma\gamma$ candidate events in the data (points, statistical uncertainty only) together with the spectra from simulated GGM ($m_{\tilde{g}} = 1000 \text{ GeV}$, $m_{\tilde{\chi}_1^0} = 450 \text{ GeV}$), SPS8 ($\Lambda = 190 \text{ TeV}$), and UED ($1/R = 1.3 \text{ TeV}$) samples after the diphoton requirement. The signal samples are scaled by a factor of 100 for clarity.

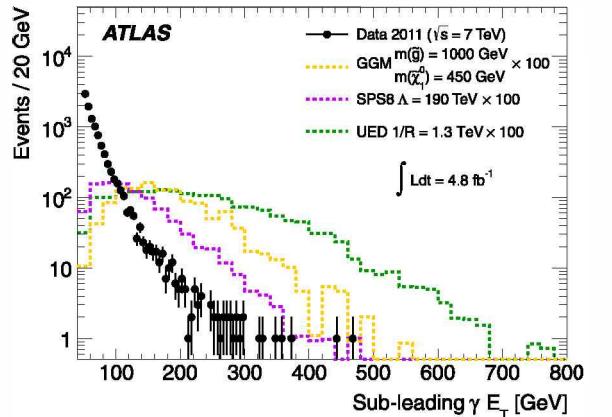


Fig. 2. The E_T spectrum of the sub-leading photon in the $\gamma\gamma$ candidate events in the data (points, statistical uncertainty only) together with the spectra from simulated GGM ($m_{\tilde{g}} = 1000 \text{ GeV}$, $m_{\tilde{\chi}_1^0} = 450 \text{ GeV}$), SPS8 ($\Lambda = 190 \text{ TeV}$), and UED ($1/R = 1.3 \text{ TeV}$) samples after the diphoton requirement. The signal samples are scaled by a factor of 100 for clarity.

distributions of selected diphoton events, with those of the same signal models overlaid.

To maximise the sensitivity of this analysis over a wide range of model parameters that may lead to different kinematic properties, three different signal regions (SRs) were defined based on the observed values of E_T^{miss} , H_T and $\Delta\phi_{\min}(\gamma, E_T^{\text{miss}})$. SR A, optimised for gluino/squark production with a subsequent decay to a high-mass bino, requires large E_T^{miss} and moderate H_T . SR B, optimised for gluino/squark production with a subsequent decay to a low-mass bino, requires moderate E_T^{miss} and large H_T . SR C, optimised for the electroweak production of intermediate-mass gaugino pairs that dominates the SPS8 cross section in this regime, requires moderate E_T^{miss} but makes no requirement on H_T . In addition, a requirement of $\Delta\phi_{\min}(\gamma, E_T^{\text{miss}}) > 0.5$ was imposed on events in SR A and C; for the low-mass bino targeted by SR B, the separation between the photon and gravitino daughters of the bino is too slight to allow for the efficient separation of signal from background through the use of this observable. The selection requirements of the three SRs are summarised in Table 1. Of the three SRs, SR A provides the greatest sensitivity to the UED model, and is thus the SR used to test this model.

² An excess of events relative to a smoothly-falling distribution of the leading-photon spectrum was observed for $E_T \sim 285 \text{ GeV}$. Searching over the range $100 \text{ GeV} < E_T < 500 \text{ GeV}$, a significance of 1.9σ was found using BumpHunter [60], while the local significance was found to be 3.1σ . No correlation between the excess and the LHC running period or luminosity was observed. A comparison of other observables (e.g. diphoton mass, E_T^{miss} , leading-photon η , $\Delta\phi(\gamma_1, \gamma_2)$) between the excess and sideband regions exhibited no appreciable differences. It was concluded that the observed excess of events is compatible with a statistical fluctuation.

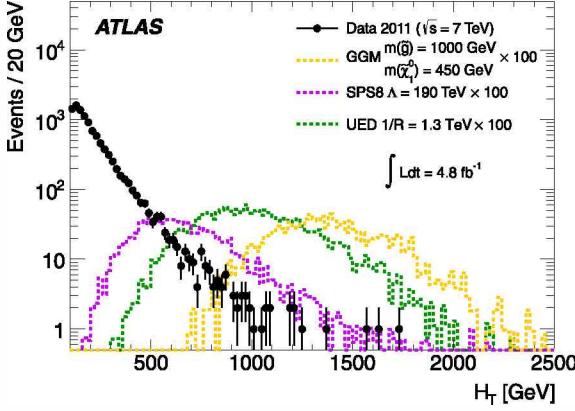


Fig. 3. The H_T spectrum of $\gamma\gamma$ candidate events in the data (points, statistical uncertainty only) together with the spectra from simulated GGM ($m_{\tilde{g}} = 1000$ GeV, $m_{\tilde{\chi}_1^0} = 450$ GeV), SPS8 ($\Lambda = 190$ TeV), and UED ($1/R = 1.3$ TeV) samples after the diphoton requirement. The signal samples are scaled by a factor of 100 for clarity.

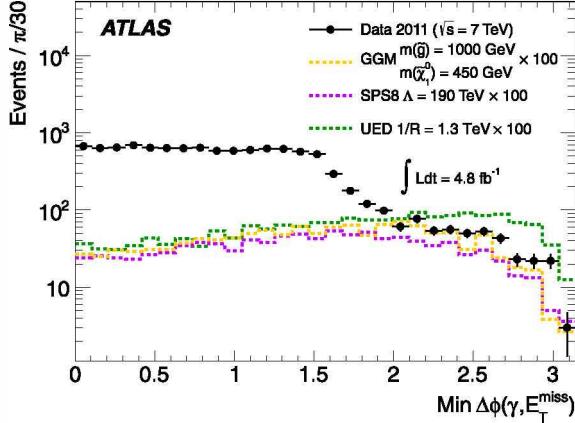


Fig. 4. The minimum $\Delta\phi(\gamma, E_T^{\text{miss}})$ spectrum of $\gamma\gamma$ candidate events in the data (points, statistical uncertainty only) together with the spectra from simulated GGM ($m_{\tilde{g}} = 1000$ GeV, $m_{\tilde{\chi}_1^0} = 450$ GeV), SPS8 ($\Lambda = 190$ TeV), and UED ($1/R = 1.3$ TeV) samples after the diphoton requirement. The signal samples are scaled by a factor of 100 for clarity.

Table 1

Definition of the three SRs (A, B and C) based on the quantities E_T^{miss} , H_T and $\Delta\phi_{\min}(\gamma, E_T^{\text{miss}})$.

	SR A	SR B	SR C
$E_T^{\text{miss}} >$	200 GeV	100 GeV	125 GeV
$H_T >$	600 GeV	1100 GeV	–
$\Delta\phi_{\min}(\gamma, E_T^{\text{miss}}) >$	0.5	–	0.5

Table 2 shows the numbers of events remaining after several stages of the selection. A total of 117, 9 and 7293 candidate events were observed to pass all but the E_T^{miss} requirement of SR A, B and C, respectively. After imposing the final E_T^{miss} requirement, no events remained for SR A and B, while two events remained for SR C.

Fig. 5 shows the E_T^{miss} distribution for SR C, the expected contributions from the SPS8 MC sample with $\Lambda = 190$ TeV, and estimated background contributions from various sources (described below).

Table 2

Samples of selected events at progressive stages of the selection. Where no number is shown the cut was not applied.

Triggered events	11166060		
	10455		
Diphoton selection	A	B	C
$\Delta\phi_{\min}(\gamma, E_T^{\text{miss}})$ requirement	7293	–	7293
H_T requirement	117	9	–
E_T^{miss} requirement	0	0	2

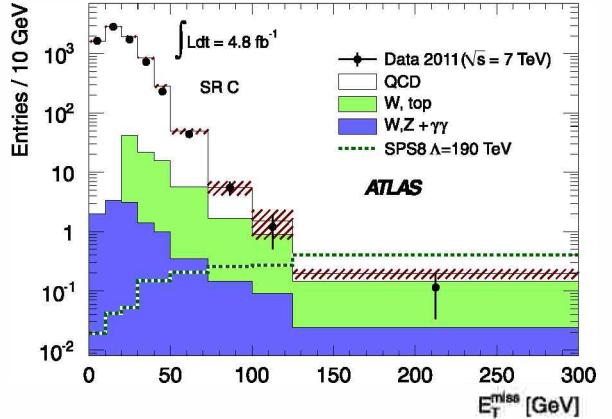


Fig. 5. E_T^{miss} spectra in SR C for the $\gamma\gamma$ candidate events in data (points, statistical uncertainty only) and the estimated QCD background (normalised to the number of $\gamma\gamma$ candidates with $E_T^{\text{miss}} < 20$ GeV), the $W \rightarrow ev + \text{jets}/\gamma$ and $t\bar{t} \rightarrow ev + \text{jets}$ backgrounds as estimated from the electron-photon control sample, and the irreducible background of $Z \rightarrow \nu\bar{\nu} + \gamma\gamma$ and $W \rightarrow \ell\nu + \gamma\gamma$. The hatched region represents the extent of the uncertainty on the total background prediction. Also shown is the expected signal from the SPS8 ($\Lambda = 190$ TeV) sample.

8. Background estimation

Following the procedure described in Ref. [61], the contribution to the large E_T^{miss} diphoton sample from SM sources can be grouped into three primary components. The first of these, referred to as “QCD background”, arises from a mixture of processes that include $\gamma\gamma$ production as well as $\gamma + \text{jet}$ and multijet events with at least one jet mis-reconstructed as a photon. The second background component, referred to as “EW background”, is due to $W + X$ and $t\bar{t}$ events (here “X” can be any number of photons or jets) for which mis-reconstructed photons arise from electrons and jets, and for which final-state neutrinos produce significant E_T^{miss} . The QCD and EW backgrounds were estimated via dedicated control samples of data events. The third background component, referred to as “irreducible”, consists of W and Z bosons produced in association with two real photons, with a subsequent decay into one or more neutrinos.

To estimate the QCD background from $\gamma\gamma$, $\gamma + \text{jet}$, and multijet events, a “QCD control sample” was selected from the diphoton trigger sample by selecting events for which at least one of the photon candidates passes the loose but not the tight photon identification. Events with electrons were vetoed to remove contamination from $W \rightarrow ev$ decays. The H_T and $\Delta\phi_{\min}(\gamma, E_T^{\text{miss}})$ requirements associated with each of the three SRs were then applied, yielding three separate QCD samples, or “templates”. An estimate of the QCD background contamination in each SR was obtained from imposing the E_T^{miss} requirement associated with the given SR upon the corresponding QCD template, after normalising each template to the diphoton data with $E_T^{\text{miss}} < 20$ GeV from the given SR. This yielded a QCD background expectation of $0.85 \pm 0.30(\text{stat})$ events for SR C. No events above the corresponding E_T^{miss} requirement were observed for the A and B control samples, yielding an

Table 3

The expected number of $\gamma\gamma$ events for each of the three signal regions. The uncertainties are statistical, arising from the limited numbers of events in the control samples, and systematic, the details of which are given in the text. For the irreducible background, the statistical uncertainty is due to limited numbers of events in the corresponding MC samples.

	SR A	SR B	SR C
QCD	$0.07 \pm 0.00 \pm 0.07$	$0.27 \pm 0.00 \pm 0.27$	$0.85 \pm 0.30 \pm 0.71$
Electroweak	$0.03 \pm 0.03 \pm 0.01$	$0.09 \pm 0.05 \pm 0.02$	$0.80 \pm 0.16 \pm 0.22$
$W(\rightarrow \ell\nu) + \gamma\gamma$	< 0.01	< 0.01	$0.18 \pm 0.13 \pm 0.18$
$Z(\rightarrow \nu\bar{\nu}) + \gamma\gamma$	< 0.01	< 0.01	$0.27 \pm 0.09 \pm 0.04$
Total	$0.10 \pm 0.03 \pm 0.07$	$0.36 \pm 0.05 \pm 0.27$	$2.11 \pm 0.37 \pm 0.77$
Observed events	0	0	2

estimate of 0 events with a 90% confidence-level (CL) upper limit of less than 1.01 and 1.15 background events for SR A and SR B, respectively.

To improve the constraint on the estimated background for SRs A and B, a complementary method making use of H_T sidebands of the QCD control sample was employed. The H_T requirement applied to the QCD templates of SR A and B was relaxed in three steps: to 400 GeV, 200 GeV and 0 GeV for the SR A control sample, and to 800 GeV, 400 GeV and 200 GeV for the SR B control sample. For each SR, the E_T^{miss} distribution of each of these relaxed control samples was scaled to the diphoton E_T^{miss} distribution for $E_T^{\text{miss}} < 20$ GeV of the given SR, yielding a series of three expected values for the QCD background as a function of the applied H_T requirement. The complementary estimate for the background contribution to the signal region employed a parabolic extrapolation to the actual H_T requirement used for the analysis (600 GeV and 1100 GeV for SRs A and B, respectively); a linear fit yielded a significantly lower background estimate for both SRs. The parabolic extrapolation yielded conservative upper estimates of 0.14 and 0.54 events for SRs A and B, respectively. The overall QCD background estimates for SRs A and B were taken to be 0.07 ± 0.07 (syst) and 0.27 ± 0.27 (syst) events, respectively, half of the value of this upper estimate, with systematic uncertainty assigned to cover the entire range between 0 and the upper estimate. The choice of a parabolic function constrained by three H_T points does not permit an estimation of statistical uncertainty on the extrapolation.

Other sources of systematic uncertainty in the estimated QCD background were considered. Using the E_T^{miss} distribution from a sample of $Z \rightarrow e^+e^-$ events instead of that of the QCD sample yielded estimates of 0, 0 and 0.15 events for SRs A, B and C, respectively. The difference between this estimate and that of the QCD sample was incorporated as a systematic uncertainty of ± 0.71 on the SR C QCD background estimate. Making use of the alternative ranges $5 \text{ GeV} < E_T^{\text{miss}} < 25 \text{ GeV}$ and $10 \text{ GeV} < E_T^{\text{miss}} < 30 \text{ GeV}$ over which the QCD sample was normalised to the $\gamma\gamma$ sample resulted in a further systematic uncertainty of ± 0.03 events on the QCD background estimate for SR C. The resulting QCD background estimates for the three SRs, along with their uncertainties, are compiled in Table 3.

The EW background, from $W + X$ and $t\bar{t}$ events, was estimated via an “electron-photon” control sample composed of events with at least one tight photon and one electron, each with $E_T > 50$ GeV, and scaled by the probability for an electron to be mis-reconstructed as a tight photon, as estimated from a “tag-and-probe” study of the Z boson in the ee and $e\gamma$ sample. The scaling factor varies between 2.5% ($0 < |\eta| < 0.6$) and 7.0% ($1.52 < |\eta| < 1.81$), since it depends on the amount of material in front of the calorimeter. Events with two or more tight photons were vetoed from the control sample to preserve its orthogonality to the signal sample. In case of more than one electron, the one with the highest p_T was used.

After applying corresponding selection requirements on H_T , $\Delta\phi_{\min}(\gamma, E_T^{\text{miss}})$ and E_T^{miss} , a total of 1, 3 and 26 electron-photon events were observed for SRs A, B and C, respectively. After multiplying by the η -dependent electron-to-photon mis-reconstruction probability, the resulting EW background contamination was estimated to be 0.03 ± 0.03 , 0.09 ± 0.05 and 0.80 ± 0.16 events for SRs A, B and C, respectively, where the uncertainties are statistical only.

The systematic uncertainty on the determination of the electron-to-photon mis-reconstruction probability is assessed by performing an independent tag-and-probe analysis with looser electron E_T and identification requirements. Differences with the nominal tag-and-probe analysis are taken as systematic uncertainty on the EW background estimate, resulting in relative systematic uncertainties of $\pm 6.9\%$, $\pm 7.1\%$ and $\pm 10.0\%$ for SRs A, B and C, respectively. MC studies suggest that approximately 25% of the EW background involves no electron-to-photon mis-reconstruction, and thus are not accounted for with the electron-photon control sample. These events, however, typically involve a jet-to-photon mis-reconstruction (for example, an event with one radiated photon and a hadronic τ decay with an energetic leading π^0 mis-reconstructed as a photon), and are thus potentially accounted for in the QCD background estimate. A relative systematic uncertainty of $\pm 25\%$ is conservatively assigned to the EW background estimates for all three SRs to account for this ambiguity. The resulting EW background estimates for the three SRs, along with their uncertainties, are compiled in Table 3.

The contribution of the irreducible background from the $Z(\rightarrow \nu\bar{\nu}) + \gamma\gamma$ and $W(\rightarrow \ell\nu) + \gamma\gamma$ processes was estimated using MC samples. It was found to be negligible for SRs A and B, and estimated to be $0.46 \pm 0.16 \pm 0.19$ events for SR C, where the first uncertainty is due to the limited number of events in the MC sample and the second to the uncertainty on the applied K -factor. These estimates, along with the resulting estimates for the total background from all sources, are reported in Table 3.

The contamination from cosmic-ray muons, estimated using events triggered in empty LHC bunches, was found to be negligible.

9. Signal efficiencies and systematic uncertainties

Signal efficiencies were estimated using MC simulation. GGM signal efficiencies were estimated over an area of the GGM parameter space that ranges from 800 GeV to 1300 GeV for the gluino or squark mass, and from 50 GeV to within 10 GeV of the gluino or squark mass for the neutralino mass. For SR A the efficiency increases smoothly from 1.2% to 25% for $(m_{\tilde{g}}, m_{\tilde{\chi}_1^0}) = (800, 50)$ GeV to $(1300, 1280)$ GeV, but then drops to 20% for the case for which the gluino and neutralino masses are only separated by 10 GeV. For SR B the efficiency increases smoothly from 2.8% to 26% for $(m_{\tilde{g}}, m_{\tilde{\chi}_1^0}) = (800, 790)$ GeV to $(1300, 50)$ GeV. The SPS8 signal efficiency in SR C increases smoothly from 5.9% ($A = 100$ TeV) to

Table 4

Relative systematic uncertainties on the expected signal yield for the GGM model with $m_{\tilde{g}} = 1000$ GeV and $m_{\tilde{\chi}_1^0} = 450$ GeV, the SPS8 model with $\Lambda = 190$ TeV, and the UED model with $1/R = 1.3$ TeV. For the GGM model, when the uncertainty differs for SRs A and B, it is presented as SRA/SRB. No PDF and scale uncertainties are given for the UED case as the cross section is evaluated only to LO.

Source of uncertainty	Uncertainty		
	GGM	SPS8	UED
Integrated luminosity	3.9%	3.9%	3.9%
Trigger	0.5%	0.5%	0.5%
Photon identification	4.4%	4.4%	4.4%
Photon isolation	0.9%	0.2%	0.4%
Pile-up	0.8%	0.5%	0.5%
E_T^{miss} reconstruction	3.9/1.1%	2.8%	1.5%
H_T	0.0/2.1%	—	0.4%
Signal MC statistics	3.0%	2.1%	1.4%
Total signal uncertainty	7.6/7.1%	6.8%	6.3%
PDF and scale	31%	5.5%	—
Total	32%	8.7%	6.3%

21% ($\Lambda = 250$ TeV). For SR A the UED signal efficiency increases smoothly from 28% ($1/R = 1.0$ TeV) to 37% ($1/R = 1.5$ TeV).

The various relative systematic uncertainties on the GGM, SPS8 and UED signal cross sections are summarised in Table 4 for the chosen reference points: $(m_{\tilde{g}}, m_{\tilde{\chi}_1^0}) = (1000, 450)$ GeV for GGM, $\Lambda = 190$ TeV for SPS8, and $1/R = 1.3$ TeV for UED. The uncertainty on the luminosity is $\pm 3.9\%$ [58,59]. The efficiency of the required diphoton trigger was estimated using a single photon trigger according to [62], yielding $99.8^{+0.2}_{-0.8}\%$ for events passing the diphoton selection. To estimate the systematic uncertainty due to the unknown composition of the data sample, the trigger efficiency was also evaluated on MC events using mis-reconstructed photons from filtered multijet samples and photons from signal (GGM, SPS8 and UED) samples. A conservative systematic uncertainty of $\pm 0.5\%$ was derived from the difference between the obtained efficiencies. Uncertainties on the photon selection, the photon energy scale, and the detailed material composition of the detector, as described in Ref. [61], result in an uncertainty of $\pm 4.4\%$ for the GGM, SPS8 and UED signals. The uncertainty due to the photon isolation requirement was estimated by varying the energy leakage and the pile-up corrections independently, resulting in an uncertainty of $\pm 0.9\%$, $\pm 0.2\%$ and $\pm 0.4\%$ for the GGM, SPS8 and UED signals, respectively. The influence of pile-up on the signal efficiency, evaluated by scaling the number of pile-up events in the MC simulation by a factor of 0.9 (chosen to reflect the range of uncertainty inherent in estimating and modelling the effects of pile-up), leads to a systematic uncertainty of $\pm 0.8\%$ (GGM), $\pm 0.5\%$ (SPS8) and $\pm 0.5\%$ (UED). Systematic uncertainties due to the E_T^{miss} reconstruction, estimated by varying the cluster energies and the E_T^{miss} resolution between the measured performance and MC expectations [55], contribute an uncertainty of $\pm 0.1/0.5\%$ to $\pm 5.3/16.1\%$ (GGM, SR A/B), $\pm 1.6\%$ to $\pm 9.7\%$ (SPS8) and $\pm 0.9\%$ to $\pm 2\%$ (UED). Systematic uncertainties due to the H_T reconstruction, estimated by varying the energy scale and resolution of the individual objects entering H_T , are below $\pm 0.3\%$ (GGM, SR A), $\pm 0.1\%$ to $\pm 7.3\%$ (GGM, SR B) and $\pm 0.1\%$ to $\pm 1.1\%$ (UED). The systematic uncertainties from E_T^{miss} and H_T are taken to be fully correlated. Added in quadrature, the total systematic uncertainty on the signal yield varies between $\pm 6\%$ and $\pm 20\%$ (GGM), $\pm 6\%$ and $\pm 15\%$ (SPS8), and $\pm 6\%$ and $\pm 7\%$ (UED).

The PDF and factorisation and renormalisation scale uncertainties on the GGM (SPS8) cross sections were evaluated as described in Section 4, leading to a combined systematic uncertainty between $\pm 23\text{--}39\%$, $\pm 29\text{--}49\%$ and $\pm 4.7\text{--}6.4\%$ for the GGM (gluino), GGM (squark) and SPS8 models, respectively. The different impact

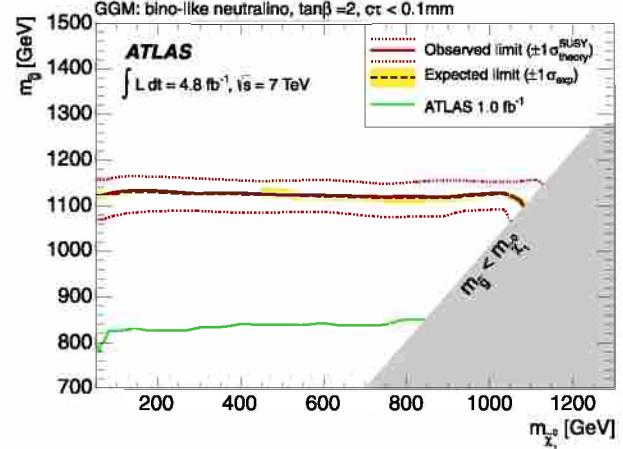


Fig. 6. Expected and observed 95% CL lower limits on the gluino mass as a function of the neutralino mass in the GGM model with a bino-like lightest neutralino NLSP (the grey area indicates the region for which the gluino mass is less than the bino mass, which is not considered here). The other sparticle masses are assumed to be decoupled. Further model parameters are $\tan\beta = 2$ and $c\tau_{\text{NLSP}} < 0.1$ mm. The previous ATLAS limit [1] is also shown.

of the PDF and scale uncertainties on the GGM and SPS8 yields is related to the different production mechanisms in the two models (see Section 2). In the case of UED, the PDF uncertainties were evaluated by using the MSTW2008 LO [63] PDF error sets in the LO cross-section calculation and are about $\pm 4\%$. The scale of α_s in the LO cross section calculation was increased and decreased by a factor of two, leading to a systematic uncertainty of $\pm 4.5\%$ and $\pm 9\%$, respectively. NLO calculations are not yet available, so the LO cross sections were used for the limit calculation without any theoretical uncertainty, and the effect of PDF and scale uncertainties on the final limit is discussed separately.

10. Results

No evidence for physics beyond the SM was observed in any of the SRs. Based on the numbers of observed events in SR A, B and C and the background expectation shown in Table 3, 95% CL upper limits are set on the numbers of events in the different SRs from any scenario of physics beyond the SM using the profile likelihood and CL_s prescriptions [64]. Uncertainties on the background expectation are treated as Gaussian-distributed nuisance parameters in the maximum likelihood fit, resulting in observed upper limits of 3.1, 3.1 and 4.9 events for SRs A, B and C, respectively. In the context of the GGM model, these limits translate into 95% upper limits on the visible cross section for new physics, defined by the product of cross section, branching ratio, acceptance and efficiency for the different SR definitions, of 0.6, 0.6 and 1.0 fb, respectively. As for background uncertainties, uncertainties on the luminosity, acceptance and efficiency are taken into account as Gaussian-distributed nuisance parameters in the maximum likelihood fit. Because the observed numbers of events are close to the expected numbers of background events for all three SRs, expected limits on the numbers of events from and visible cross section for new physics are, to the quoted accuracy, identical to the observed limits.

Limits are also set on the GGM squark and gluino masses as a function of the bino-like neutralino mass, making use of the SR (A or B) that provides the most stringent expected limit for the given neutralino mass. Figs. 6 and 7 show the expected and observed lower limits on the GGM gluino and squark masses, respectively, as a function of the neutralino mass. Three observed-limit contours are shown, corresponding to the nominal assumption for the SUSY production cross section as well as those derived by

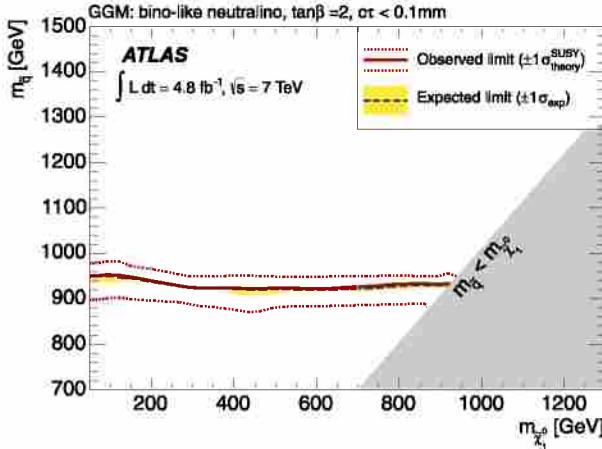


Fig. 7. Expected and observed 95% CL lower limits on the squark mass as a function of the neutralino mass in the GGM model with a bino-like lightest neutralino NLSP (the grey area indicates the region for which the squark mass is less than the bino mass, which is not considered here). The other sparticle masses are assumed to be decoupled. Further model parameters are $\tan\beta = 2$ and $c\tau_{\text{NLSP}} < 0.1 \text{ mm}$.

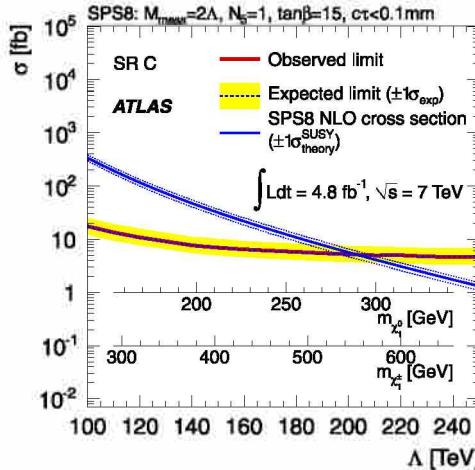


Fig. 8. Expected and observed 95% CL upper limits on the sparticle production cross section in the SPS8 model, and the NLO cross-section prediction, as a function of Λ and the lightest neutralino and chargino masses. Further SPS8 model parameters are $M_{\text{mess}} = 2\Lambda$, $N_5 = 1$, $\tan\beta = 15$ and $c\tau_{\text{NLSP}} < 0.1 \text{ mm}$. Limits are set based on SR C.

reducing and increasing the cross section by one standard deviation of theoretical uncertainty (the combined uncertainty due to the PDFs and renormalisation and factorisation scales). For comparison the lower limits on the GGM gluino mass from ATLAS [1] based on 1 fb^{-1} from 2011 are also shown.

Including all sources of uncertainty other than the theoretical uncertainty, 95% CL upper limits on the cross section of the SPS8 model are derived from the SR C result and displayed in Fig. 8 for the range $\Lambda = 100\text{--}250 \text{ TeV}$ along with the overall production cross section and its theoretical uncertainty. For illustration the cross-section dependence as a function of the lightest neutralino and chargino masses is also shown.

Fig. 9 shows the limit on the cross section times branching ratio for the UED model as a function of the compactification scale $1/R$, derived from the result of SR A. A 95% CL lower limit of $1/R > 1.40 \text{ TeV}$ is set. For illustration the cross-section dependence as a function of the KK quark and KK gluon masses is also shown. Again, neither PDF nor scale uncertainties are included when calculating the limits; including PDF and scale uncertainties, com-

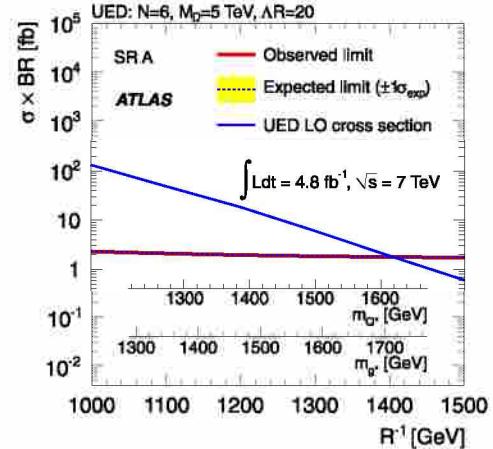


Fig. 9. Expected and observed 95% CL upper limits on the KK particle production cross section times branching ratio to two photons in the UED model, and the LO cross-section prediction times branching ratio, as a function of $1/R$ and the KK quark (Q^*) and KK gluon (g^*) masses. The $\pm 1\sigma$ expected-limit error band overlaps the observed limit contour and is too narrow to be distinguished. No error is shown for the UED cross section since the cross-section calculation is available only to LO (see text for further discussion). The UED model parameters are $N = 6$, $M_D = 5 \text{ TeV}$ and $\Delta R = 20$. Limits are set based on SR A.

puted at LO, in the limit calculation degrades the limit on $1/R$ by a few GeV.

11. Conclusions

A search for events with two photons and substantial E_T^{miss} , performed using 4.8 fb^{-1} of 7 TeV pp collision data recorded with the ATLAS detector at the LHC, is presented. The sensitivity to different new physics models producing this final state was optimised by defining three different signal regions. No significant excess above the expected background is found in any signal region. The results are used to set model-independent 95% CL upper limits on possible contributions from new physics. In addition, under the GGM hypothesis, considering cross sections one standard deviation of theoretical uncertainty below the nominal value, a lower limit on the gluino/squark mass of $1.07 \text{ TeV}/0.87 \text{ TeV}$ is determined for bino masses above 50 GeV . Under similar assumptions, a lower limit of 196 TeV is set on the SUSY-breaking scale Λ of the SPS8 model. Considering nominal values of the leading-order UED cross section, a lower limit of 1.40 TeV is set on the UED compactification scale $1/R$. These results provide the most stringent tests of these models to date.

Acknowledgements

We thank CERN for the very successful operation of the LHC, as well as the support staff from our institutions without whom ATLAS could not be operated efficiently.

We acknowledge the support of ANPCyT, Argentina; YerPhI, Armenia; ARC, Australia; BMWF and FWF, Austria; ANAS, Azerbaijan; SSTD, Belarus; CNPq and FAPESP, Brazil; NSERC, NRC and CFI, Canada; CERN; CONICYT, Chile; CAS, MOST and NSFC, China; COLCIENCIAS, Colombia; MSMT CR, MPO CR and VSC CR, Czech Republic; DNR, DNSRC and Lundbeck Foundation, Denmark; EPLANET and ERC, European Union; IN2P3-CNRS, CEA-DSM/IRFU, France; GNSF, Georgia; BMBF, DFG, HGF, MPG and AvH Foundation, Germany; GSRT, Greece; ISF, MINERVA, GIF, DIP and Benoziyo Center, Israel; INFN, Italy; MEXT and JSPS, Japan; CNRST, Morocco; FOM and NWO, Netherlands; RCN, Norway; MNiSW, Poland; GRICES and FCT, Portugal; MERYS (MECTS), Romania; MES of Russia and

ROSATOM, Russian Federation; JINR; MSTD, Serbia; MSSR, Slovakia; ARRS and MVZT, Slovenia; DST/NRF, South Africa; MICINN, Spain; SRC and Wallenberg Foundation, Sweden; SER, SNSF and Cantons of Bern and Geneva, Switzerland; NSC, Taiwan; TAEK, Turkey; STFC, the Royal Society and Leverhulme Trust, United Kingdom; DOE and NSF, United States of America.

The crucial computing support from all WLCG partners is acknowledged gratefully, in particular from CERN and the ATLAS Tier-1 facilities at TRIUMF (Canada), NDGF (Denmark, Norway, Sweden), CC-IN2P3 (France), KIT/GridKA (Germany), INFN-CNAF (Italy), NL-T1 (Netherlands), PIC (Spain), ASGC (Taiwan), RAL (UK) and BNL (USA) and in the Tier-2 facilities worldwide.

Open access

This article is published Open Access at sciedirect.com. It is distributed under the terms of the Creative Commons Attribution License 3.0, which permits unrestricted use, distribution, and reproduction in any medium, provided the original authors and source are credited.

References

- [1] ATLAS Collaboration, Phys. Lett. B 710 (2012) 519, arXiv:1111.4116 [hep-ex].
- [2] P. Meade, N. Seiberg, D. Shih, Prog. Theor. Phys. Suppl. 177 (2009) 143, arXiv: 0801.3278 [hep-ph].
- [3] M. Buican, P. Meade, N. Seiberg, D. Shih, JHEP 0903 (2009) 016, arXiv: 0812.3668 [hep-ph].
- [4] J.T. Ruderman, D. Shih, JHEP 1208 (2012) 159, arXiv:1103.6083 [hep-ph].
- [5] B.C. Allanach, et al., Eur. Phys. J. C 25 (2002) 113, arXiv:hep-ph/0202233.
- [6] T. Appelquist, H.-C. Cheng, B.A. Dobrescu, Phys. Rev. D 64 (2001) 035002, arXiv:hep-ph/0012100.
- [7] C. Macesanu, C. McMullen, S. Nandi, Phys. Lett. B 546 (2002) 253, arXiv: hep-ph/0207269.
- [8] C. Macesanu, Int. J. Mod. Phys. A 21 (2006) 2259, arXiv:hep-ph/0510418.
- [9] H. Miyazawa, Prog. Theor. Phys. 36 (6) (1966) 1266.
- [10] P. Ramond, Phys. Rev. D 3 (1971) 2415.
- [11] Y.A. Gol'fand, E.P. Likhtman, JETP Lett. 7113 (1971) 323; Y.A. Gol'fand, E.P. Likhtman, Pisma Zh. Eksp. Teor. Fiz. 13 (1971) 452.
- [12] A. Neveu, J.H. Schwarz, Nucl. Phys. B 31 (1971) 86.
- [13] A. Neveu, J.H. Schwarz, Phys. Rev. D 4 (1971) 1109.
- [14] J. Gervais, B. Sakita, Nucl. Phys. B 34 (1971) 632.
- [15] D.V. Volkov, V.P. Akulov, Phys. Lett. B 46 (1973) 109.
- [16] J. Wess, B. Zumino, Phys. Lett. B 49 (1974) 52.
- [17] J. Wess, B. Zumino, Nucl. Phys. B 70 (1974) 39.
- [18] P. Fayet, Phys. Lett. B 64 (1976) 159.
- [19] P. Fayet, Phys. Lett. B 69 (1977) 489.
- [20] G.R. Farrar, P. Fayet, Phys. Lett. B 76 (1978) 575.
- [21] P. Fayet, Phys. Lett. B 84 (1979) 416.
- [22] S. Dimopoulos, H. Georgi, Nucl. Phys. B 193 (1981) 150.
- [23] M. Dine, W. Fischler, Phys. Lett. B 110 (1982) 227.
- [24] L. Alvarez-Gaume, M. Claudson, M.B. Wise, Nucl. Phys. B 207 (1982) 96.
- [25] C.R. Nappi, B.A. Ovrut, Phys. Lett. B 113 (1982) 175.
- [26] M. Dine, A.E. Nelson, Phys. Rev. D 48 (1993) 1277, arXiv:hep-ph/9303230.
- [27] M. Dine, A.E. Nelson, Y. Shirman, Phys. Rev. D 51 (1995) 1362, arXiv:hep-ph/9408384.
- [28] M. Dine, A.E. Nelson, Y. Nir, Y. Shirman, Phys. Rev. D 53 (1996) 2658, arXiv: hep-ph/9507378.
- [29] CMS Collaboration, JHEP 1106 (2011) 093, arXiv:1105.3152 [hep-ex].
- [30] C. Macesanu, C. McMullen, S. Nandi, Phys. Rev. D 66 (2002) 015009, arXiv: hep-ph/0201300.
- [31] A. De Rujula, A. Donini, M. Gavela, S. Rigolin, Phys. Lett. B 482 (2000) 195, arXiv:hep-ph/0001335.
- [32] H.-C. Cheng, K.T. Matchev, M. Schmaltz, Phys. Rev. D 66 (2002) 036005, arXiv:hep-ph/0204342.
- [33] M. Elkacimi, D. Goujdami, H. Przysiezniak, P.Z. Skands, Comput. Phys. Commun. 181 (2010) 122, arXiv:0901.4087 [hep-ph].
- [34] A. Djouadi, J.-L. Kneur, G. Moultaka, Comput. Phys. Commun. 176 (2007) 426, arXiv:hep-ph/0211331.
- [35] M. Mühlleitner, A. Djouadi, Y. Mambrini, Comput. Phys. Commun. 168 (2005) 46, arXiv:hep-ph/0311167.
- [36] F.E. Paige, S.D. Protopopescu, H. Baer, X. Tata, ISAJET 7.69: A Monte Carlo event generator for pp , $\bar{p}p$, and e^+e^- reactions, arXiv:hep-ph/0312045.
- [37] M. Bahr, et al., Eur. Phys. J. C 58 (2008) 639, arXiv:0803.0883 [hep-ph].
- [38] A. Sherstnev, R.S. Thorne, Eur. Phys. J. C 55 (2008) 553, arXiv:0711.2473 [hep-ph].
- [39] W. Beenakker, R. Hopker, M. Spira, P. Zerwas, Nucl. Phys. B 492 (1997) 51, arXiv:hep-ph/9610490.
- [40] A. Kulesza, L. Motyka, Phys. Rev. Lett. 102 (2009) 111802, arXiv:0807.2405 [hep-ph].
- [41] A. Kulesza, L. Motyka, Phys. Rev. D 80 (2009) 095004, arXiv:0905.4749 [hep-ph].
- [42] W. Beenakker, S. Brening, M. Kramer, A. Kulesza, E. Laenen, et al., JHEP 0912 (2009) 041, arXiv:0909.4418 [hep-ph].
- [43] W. Beenakker, S. Brening, M. Kramer, A. Kulesza, E. Laenen, et al., Int. J. Mod. Phys. A 26 (2011) 2637, arXiv:1105.1110 [hep-ph].
- [44] M. Kramer, A. Kulesza, R. van der Leeuw, M. Mangano, S. Padhi, et al., Supersymmetry production cross sections in pp collisions at $\sqrt{s} = 7$ TeV, arXiv:1206.2892 [hep-ph].
- [45] T. Sjostrand, S. Mrenna, P. Skands, JHEP 0605 (2006) 026, arXiv:hep-ph/0603175.
- [46] J. Alwall, et al., JHEP 0709 (2007) 028, arXiv:0706.2334 [hep-ph].
- [47] D. Stump, et al., JHEP 0310 (2003) 046, arXiv:hep-ph/0303013.
- [48] G. Bozzi, F. Campanario, M. Rauch, D. Zeppenfeld, Phys. Rev. D 84 (2011) 074028, arXiv:1107.3149 [hep-ph].
- [49] G. Bozzi, F. Campanario, M. Rauch, D. Zeppenfeld, Phys. Rev. D 83 (2011) 114035, arXiv:1103.4613 [hep-ph].
- [50] GEANT4 Collaboration, S. Agostinelli, et al., Nucl. Instrum. Meth. A 506 (2003) 250.
- [51] ATLAS Collaboration, Eur. Phys. J. C 70 (2010) 823, arXiv:1005.4568 [physics.ins-det].
- [52] ATLAS Collaboration, JINST 3 (2008) S08003.
- [53] ATLAS Collaboration, Phys. Rev. D 83 (2011) 052005, arXiv:1012.4389 [hep-ex].
- [54] ATLAS Collaboration, Eur. Phys. J. C 72 (2012) 1909, arXiv:1110.3174 [hep-ex].
- [55] ATLAS Collaboration, Eur. Phys. J. C 72 (2012) 1844, arXiv:1108.5602 [hep-ex].
- [56] M. Cacciari, G. Salam, G. Soyez, JHEP 0804 (2008) 063, arXiv:0802.1189.
- [57] ATLAS Collaboration, Eur. Phys. J. C, submitted for publication, arXiv:1112.6426 [hep-ex].
- [58] ATLAS Collaboration, Luminosity determination in pp collisions at $\sqrt{s} = 7$ TeV using the ATLAS detector in 2011, ATLAS-CONF-2011-116, <http://cdsweb.cern.ch/record/1376384>.
- [59] ATLAS Collaboration, Eur. Phys. J. C 71 (2011) 1630, arXiv:1101.2185 [hep-ex].
- [60] G. Choudalakis, On hypothesis testing, trials factor, hypertests and the Bump-Hunter, arXiv:1101.0390 [physics.data-an].
- [61] ATLAS Collaboration, Eur. Phys. J. C 71 (2011) 1744, arXiv:1107.0561 [hep-ex].
- [62] ATLAS Collaboration, Eur. Phys. J. C 72 (2012) 1849, arXiv:1110.1530 [hep-ex].
- [63] A. Martin, W. Stirling, R. Thorne, G. Watt, Eur. Phys. J. C 63 (2008) 189, arXiv: 0901.0002 [hep-ph].
- [64] A.L. Read, J. Phys. G 28 (2002) 2693.

ATLAS Collaboration

G. Aad⁴⁸, T. Abajyan²¹, B. Abbott¹¹¹, J. Abdallah¹², S. Abdel Khalek¹¹⁵, A.A. Abdelalim⁴⁹, O. Abdinov¹¹, R. Aben¹⁰⁵, B. Abi¹¹², M. Abolins⁸⁸, O.S. AbouZeid¹⁵⁸, H. Abramowicz¹⁵³, H. Abreu¹³⁶, E. Acerbi^{89a,89b}, B.S. Acharya^{164a,164b}, L. Adamczyk³⁸, D.L. Adams²⁵, T.N. Addy⁵⁶, J. Adelman¹⁷⁶, S. Adomeit⁹⁸, P. Adragna⁷⁵, T. Adye¹²⁹, S. Aefsky²³, J.A. Aguilar-Saavedra^{124b,a}, M. Agustoni¹⁷, M. Aharrouche⁸¹, S.P. Ahlen²², F. Ahles⁴⁸, A. Ahmad¹⁴⁸, M. Ahsan⁴¹, G. Aielli^{133a,133b}, T. Akdogan^{19a}, T.P.A. Åkesson⁷⁹, G. Akimoto¹⁵⁵, A.V. Akimov⁹⁴, M.S. Alam², M.A. Alam⁷⁶, J. Albert¹⁶⁹, S. Albrand⁵⁵, M. Aleksi³⁰, I.N. Aleksandrov⁶⁴, F. Alessandria^{89a}, C. Alexa^{26a}, G. Alexander¹⁵³, G. Alexandre⁴⁹, T. Alexopoulos¹⁰, M. Alhroob^{164a,164c}, M. Aliev¹⁶, G. Alimonti^{89a}, J. Alison¹²⁰, B.M.M. Allbrooke¹⁸, P.P. Allport⁷³,

- S.E. Allwood-Spiers⁵³, J. Almond⁸², A. Aloisio^{102a,102b}, R. Alon¹⁷², A. Alonso⁷⁹, F. Alonso⁷⁰,
 B. Alvarez Gonzalez⁸⁸, M.G. Alviggi^{102a,102b}, K. Amako⁶⁵, C. Amelung²³, V.V. Ammosov^{128,*},
 S.P. Amor Dos Santos^{124a}, A. Amorim^{124a,b}, N. Amram¹⁵³, C. Anastopoulos³⁰, L.S. Ancu¹⁷, N. Andari¹¹⁵,
 T. Andeen³⁵, C.F. Anders^{58b}, G. Anders^{58a}, K.J. Anderson³¹, A. Andreazza^{89a,89b}, V. Andrei^{58a},
 M.-L. Andrieux⁵⁵, X.S. Anduaga⁷⁰, P. Anger⁴⁴, A. Angerami³⁵, F. Anghinolfi³⁰, A. Anisenkov¹⁰⁷,
 N. Anjos^{124a}, A. Annovi⁴⁷, A. Antonaki⁹, M. Antonelli⁴⁷, A. Antonov⁹⁶, J. Antos^{144b}, F. Anulli^{132a},
 M. Aoki¹⁰¹, S. Aoun⁸³, L. Aperio Bella⁵, R. Apolle^{118,c}, G. Arabidze⁸⁸, I. Aracena¹⁴³, Y. Arai⁶⁵,
 A.T.H. Arce⁴⁵, S. Arfaoui¹⁴⁸, J.-F. Arguin¹⁵, E. Arik^{19a,*}, M. Arik^{19a}, A.J. Armbruster⁸⁷, O. Arnaez⁸¹,
 V. Arnal⁸⁰, C. Arnault¹¹⁵, A. Artamonov⁹⁵, G. Artoni^{132a,132b}, D. Arutinov²¹, S. Asai¹⁵⁵,
 R. Asfandiyarov¹⁷³, S. Ask²⁸, B. Åsman^{146a,146b}, L. Asquith⁶, K. Assamagan²⁵, A. Astbury¹⁶⁹,
 M. Atkinson¹⁶⁵, B. Aubert⁵, E. Auge¹¹⁵, K. Augsten¹²⁷, M. Aurousseau^{145a}, G. Avolio¹⁶³,
 R. Avramidou¹⁰, D. Axen¹⁶⁸, G. Azuelos^{93,d}, Y. Azuma¹⁵⁵, M.A. Baak³⁰, G. Baccaglioni^{89a},
 C. Bacci^{134a,134b}, A.M. Bach¹⁵, H. Bachacou¹³⁶, K. Bachas³⁰, M. Backes⁴⁹, M. Backhaus²¹, E. Badescu^{26a},
 P. Bagnaia^{132a,132b}, S. Bahinipati³, Y. Bai^{33a}, D.C. Bailey¹⁵⁸, T. Bain¹⁵⁸, J.T. Baines¹²⁹, O.K. Baker¹⁷⁶,
 M.D. Baker²⁵, S. Baker⁷⁷, E. Banas³⁹, P. Banerjee⁹³, Sw. Banerjee¹⁷³, D. Banfi³⁰, A. Bangert¹⁵⁰,
 V. Bansal¹⁶⁹, H.S. Bansil¹⁸, L. Barak¹⁷², S.P. Baranov⁹⁴, A. Barbaro Galtieri¹⁵, T. Barber⁴⁸,
 E.L. Barberio⁸⁶, D. Barberis^{50a,50b}, M. Barbero²¹, D.Y. Bardin⁶⁴, T. Barillari⁹⁹, M. Barisonzi¹⁷⁵,
 T. Barklow¹⁴³, N. Barlow²⁸, B.M. Barnett¹²⁹, R.M. Barnett¹⁵, A. Baroncelli^{134a}, G. Barone⁴⁹, A.J. Barr¹¹⁸,
 F. Barreiro⁸⁰, J. Barreiro Guimarães da Costa⁵⁷, P. Barrillon¹¹⁵, R. Bartoldus¹⁴³, A.E. Barton⁷¹,
 V. Bartsch¹⁴⁹, A. Basye¹⁶⁵, R.L. Bates⁵³, L. Batkova^{144a}, J.R. Batley²⁸, A. Battaglia¹⁷, M. Battistin³⁰,
 F. Bauer¹³⁶, H.S. Bawa^{143,e}, S. Beale⁹⁸, T. Beau⁷⁸, P.H. Beauchemin¹⁶¹, R. Beccherle^{50a}, P. Bechtle²¹,
 H.P. Beck¹⁷, A.K. Becker¹⁷⁵, S. Becker⁹⁸, M. Beckingham¹³⁸, K.H. Becks¹⁷⁵, A.J. Beddall^{19c}, A. Beddall^{19c},
 S. Bedikian¹⁷⁶, V.A. Bednyakov⁶⁴, C.P. Bee⁸³, L.J. Beemster¹⁰⁵, M. Begel²⁵, S. Behar Harpaz¹⁵²,
 P.K. Behera⁶², M. Beimforde⁹⁹, C. Belanger-Champagne⁸⁵, P.J. Bell⁴⁹, W.H. Bell⁴⁹, G. Bella¹⁵³,
 L. Bellagamba^{20a}, F. Bellina³⁰, M. Bellomo³⁰, A. Belloni⁵⁷, O. Beloborodova^{107,f}, K. Belotskiy⁹⁶,
 O. Beltramello³⁰, O. Benary¹⁵³, D. Benchekroun^{135a}, K. Bendtz^{146a,146b}, N. Benekos¹⁶⁵,
 Y. Benhammou¹⁵³, E. Benhar Noccioli⁴⁹, J.A. Benitez Garcia^{159b}, D.P. Benjamin⁴⁵, M. Benoit¹¹⁵,
 J.R. Bensinger²³, K. Benslama¹³⁰, S. Bentvelsen¹⁰⁵, D. Berge³⁰, E. Bergeaas Kuutmann⁴², N. Berger⁵,
 F. Berghaus¹⁶⁹, E. Berglund¹⁰⁵, J. Beringer¹⁵, P. Bernat⁷⁷, R. Bernhard⁴⁸, C. Bernius²⁵, T. Berry⁷⁶,
 C. Bertella⁸³, A. Bertin^{20a,20b}, F. Bertolucci^{122a,122b}, M.I. Besana^{89a,89b}, G.J. Besjes¹⁰⁴, N. Besson¹³⁶,
 S. Bethke⁹⁹, W. Bhimji⁴⁶, R.M. Bianchi³⁰, M. Bianco^{72a,72b}, O. Biebel⁹⁸, S.P. Bieniek⁷⁷, K. Bierwagen⁵⁴,
 J. Biesiada¹⁵, M. Biglietti^{134a}, H. Bilokon⁴⁷, M. Bindi^{20a,20b}, S. Binet¹¹⁵, A. Bingul^{19c}, C. Bini^{132a,132b},
 C. Biscarat¹⁷⁸, B. Bittner⁹⁹, K.M. Black²², R.E. Blair⁶, J.-B. Blanchard¹³⁶, G. Blanchot³⁰, T. Blazek^{144a},
 I. Bloch⁴², C. Blocker²³, J. Blocki³⁹, A. Blondel⁴⁹, W. Blum⁸¹, U. Blumenschein⁵⁴, G.J. Bobbink¹⁰⁵,
 V.B. Bobrovnikov¹⁰⁷, S.S. Bocchetta⁷⁹, A. Bocci⁴⁵, C.R. Boddy¹¹⁸, M. Boehler⁴⁸, J. Boek¹⁷⁵, N. Boelaert³⁶,
 J.A. Bogaerts³⁰, A. Bogdanchikov¹⁰⁷, A. Bogouch^{90,*}, C. Bohm^{146a}, J. Bohm¹²⁵, V. Boisvert⁷⁶, T. Bold³⁸,
 V. Boldea^{26a}, N.M. Bolnet¹³⁶, M. Bomben⁷⁸, M. Bona⁷⁵, M. Boonekamp¹³⁶, C.N. Booth¹³⁹, S. Bordoni⁷⁸,
 C. Borer¹⁷, A. Borisov¹²⁸, G. Borissov⁷¹, I. Borjanovic^{13a}, M. Borri⁸², S. Borroni⁸⁷, V. Bortolotto^{134a,134b},
 K. Bos¹⁰⁵, D. Boscherini^{20a}, M. Bosman¹², H. Boterenbrood¹⁰⁵, J. Bouchami⁹³, J. Boudreau¹²³,
 E.V. Bouhouva-Thacker⁷¹, D. Boumediene³⁴, C. Bourdarios¹¹⁵, N. Bousson⁸³, A. Boveia³¹, J. Boyd³⁰,
 I.R. Boyko⁶⁴, I. Bozovic-Jelisavcic^{13b}, J. Bracinik¹⁸, J. Bradmiller-feld¹²⁰, P. Branchini^{134a},
 G.W. Brandenburg⁵⁷, A. Brandt⁸, G. Brandt¹¹⁸, O. Brandt⁵⁴, U. Bratzler¹⁵⁶, B. Brau⁸⁴, J.E. Brau¹¹⁴,
 H.M. Braun^{175,*}, S.F. Brazzale^{164a,164c}, B. Brelier¹⁵⁸, J. Bremer³⁰, K. Brendlinger¹²⁰, R. Brenner¹⁶⁶,
 S. Bressler¹⁷², D. Britton⁵³, F.M. Brochu²⁸, I. Brock²¹, R. Brock⁸⁸, F. Broggi^{89a}, C. Bromberg⁸⁸,
 J. Bronner⁹⁹, G. Brooijmans³⁵, T. Brooks⁷⁶, W.K. Brooks^{32b}, G. Brown⁸², H. Brown⁸,
 P.A. Bruckman de Renstrom³⁹, D. Bruncko^{144b}, R. Bruneliere⁴⁸, S. Brunet⁶⁰, A. Bruni^{20a}, G. Bruni^{20a},
 M. Bruschi^{20a}, T. Buanes¹⁴, Q. Buat⁵⁵, F. Bucci⁴⁹, J. Buchanan¹¹⁸, P. Buchholz¹⁴¹, R.M. Buckingham¹¹⁸,
 A.G. Buckley⁴⁶, S.I. Buda^{26a}, I.A. Budagov⁶⁴, B. Budick¹⁰⁸, V. Büscher⁸¹, L. Bugge¹¹⁷, O. Bulekov⁹⁶,
 A.C. Bundock⁷³, M. Bunse⁴³, T. Buran¹¹⁷, H. Burckhart³⁰, S. Burdin⁷³, T. Burgess¹⁴, S. Burke¹²⁹,
 E. Busato³⁴, P. Bussey⁵³, C.P. Buszello¹⁶⁶, B. Butler¹⁴³, J.M. Butler²², C.M. Buttar⁵³, J.M. Butterworth⁷⁷,
 W. Buttlinger²⁸, M. Byszewski³⁰, S. Cabrera Urbán¹⁶⁷, D. Caforio^{20a,20b}, O. Cakir^{4a}, P. Calafiura¹⁵,
 G. Calderini⁷⁸, P. Calfayan⁹⁸, R. Calkins¹⁰⁶, L.P. Caloba^{24a}, R. Caloi^{132a,132b}, D. Calvet³⁴, S. Calvet³⁴,

- R. Camacho Toro ³⁴, P. Camarri ^{133a,133b}, D. Cameron ¹¹⁷, L.M. Caminada ¹⁵, R. Caminal Armadans ¹²,
 S. Campana ³⁰, M. Campanelli ⁷⁷, V. Canale ^{102a,102b}, F. Canelli ^{31,g}, A. Canepa ^{159a}, J. Cantero ⁸⁰,
 R. Cantrill ⁷⁶, L. Capasso ^{102a,102b}, M.D.M. Capeans Garrido ³⁰, I. Caprini ^{26a}, M. Caprini ^{26a}, D. Capriotti ⁹⁹,
 M. Capua ^{37a,37b}, R. Caputo ⁸¹, R. Cardarelli ^{133a}, T. Carli ³⁰, G. Carlino ^{102a}, L. Carminati ^{89a,89b}, B. Caron ⁸⁵,
 S. Caron ¹⁰⁴, E. Carquin ^{32b}, G.D. Carrillo-Montoya ¹⁷³, A.A. Carter ⁷⁵, J.R. Carter ²⁸, J. Carvalho ^{124a,h},
 D. Casadei ¹⁰⁸, M.P. Casado ¹², M. Cascella ^{122a,122b}, C. Caso ^{50a,50b,*}, A.M. Castaneda Hernandez ^{173,i},
 E. Castaneda-Miranda ¹⁷³, V. Castillo Gimenez ¹⁶⁷, N.F. Castro ^{124a}, G. Cataldi ^{72a}, P. Catastini ⁵⁷,
 A. Catinaccio ³⁰, J.R. Catmore ³⁰, A. Cattai ³⁰, G. Cattani ^{133a,133b}, S. Caughron ⁸⁸, V. Cavaliere ¹⁶⁵,
 P. Cavalleri ⁷⁸, D. Cavalli ^{89a}, M. Cavalli-Sforza ¹², V. Cavasinni ^{122a,122b}, F. Ceradini ^{134a,134b},
 A.S. Cerqueira ^{24b}, A. Cerri ³⁰, L. Cerrito ⁷⁵, F. Cerutti ⁴⁷, S.A. Cetin ^{19b}, A. Chafaq ^{135a}, D. Chakraborty ¹⁰⁶,
 I. Chalupkova ¹²⁶, K. Chan ³, P. Chang ¹⁶⁵, B. Chapleau ⁸⁵, J.D. Chapman ²⁸, J.W. Chapman ⁸⁷, E. Chareyre ⁷⁸,
 D.G. Charlton ¹⁸, V. Chavda ⁸², C.A. Chavez Barajas ³⁰, S. Cheatham ⁸⁵, S. Chekanov ⁶, S.V. Chekulaev ^{159a},
 G.A. Chelkov ⁶⁴, M.A. Chelstowska ¹⁰⁴, C. Chen ⁶³, H. Chen ²⁵, S. Chen ^{33c}, X. Chen ¹⁷³, Y. Chen ³⁵,
 A. Cheplakov ⁶⁴, R. Cherkaoui El Moursli ^{135e}, V. Chernyatin ²⁵, E. Cheu ⁷, S.L. Cheung ¹⁵⁸, L. Chevalier ¹³⁶,
 G. Chiefari ^{102a,102b}, L. Chikovani ^{51a,*}, J.T. Childers ³⁰, A. Chilingarov ⁷¹, G. Chiodini ^{72a}, A.S. Chisholm ¹⁸,
 R.T. Chislett ⁷⁷, A. Chitan ^{26a}, M.V. Chizhov ⁶⁴, G. Choudalakis ³¹, S. Chouridou ¹³⁷, I.A. Christidi ⁷⁷,
 A. Christov ⁴⁸, D. Chromek-Burckhart ³⁰, M.L. Chu ¹⁵¹, J. Chudoba ¹²⁵, G. Ciapetti ^{132a,132b}, A.K. Ciftci ^{4a},
 R. Ciftci ^{4a}, D. Cinca ³⁴, V. Cindro ⁷⁴, C. Ciocca ^{20a,20b}, A. Ciocio ¹⁵, M. Cirilli ⁸⁷, P. Cirkovic ^{13b},
 Z.H. Citron ¹⁷², M. Citterio ^{89a}, M. Ciubancan ^{26a}, A. Clark ⁴⁹, P.J. Clark ⁴⁶, R.N. Clarke ¹⁵, W. Cleland ¹²³,
 J.C. Clemens ⁸³, B. Clement ⁵⁵, C. Clement ^{146a,146b}, Y. Coadou ⁸³, M. Cobal ^{164a,164c}, A. Coccaro ¹³⁸,
 J. Cochran ⁶³, J.G. Cogan ¹⁴³, J. Coggesshall ¹⁶⁵, E. Cogneras ¹⁷⁸, J. Colas ⁵, S. Cole ¹⁰⁶, A.P. Colijn ¹⁰⁵,
 N.J. Collins ¹⁸, C. Collins-Tooth ⁵³, J. Collot ⁵⁵, T. Colombo ^{119a,119b}, G. Colon ⁸⁴, P. Conde Muiño ^{124a},
 E. Coniavitis ¹¹⁸, M.C. Conidi ¹², S.M. Consonni ^{89a,89b}, V. Consorti ⁴⁸, S. Constantinescu ^{26a},
 C. Conta ^{119a,119b}, G. Conti ⁵⁷, F. Conventi ^{102a,j}, M. Cooke ¹⁵, B.D. Cooper ⁷⁷, A.M. Cooper-Sarkar ¹¹⁸,
 K. Copic ¹⁵, T. Cornelissen ¹⁷⁵, M. Corradi ^{20a}, F. Corriveau ^{85,k}, A. Cortes-Gonzalez ¹⁶⁵, G. Cortiana ⁹⁹,
 G. Costa ^{89a}, M.J. Costa ¹⁶⁷, D. Costanzo ¹³⁹, D. Côté ³⁰, L. Courneyea ¹⁶⁹, G. Cowan ⁷⁶, C. Cowden ²⁸,
 B.E. Cox ⁸², K. Cranmer ¹⁰⁸, F. Crescioli ^{122a,122b}, M. Cristinziani ²¹, G. Crosetti ^{37a,37b}, S. Crépé-Renaudin ⁵⁵,
 C.-M. Cuciuc ^{26a}, C. Cuenda Almenar ¹⁷⁶, T. Cuhadar Donszelmann ¹³⁹, M. Curatolo ⁴⁷, C.J. Curtis ¹⁸,
 C. Cuthbert ¹⁵⁰, P. Cwetanski ⁶⁰, H. Czirr ¹⁴¹, P. Czodrowski ⁴⁴, Z. Czyczula ¹⁷⁶, S. D'Auria ⁵³,
 M. D'Onofrio ⁷³, A. D'Orazio ^{132a,132b}, M.J. Da Cunha Sargedas De Sousa ^{124a}, C. Da Via ⁸²,
 W. Dabrowski ³⁸, A. Dafinca ¹¹⁸, T. Dai ⁸⁷, C. Dallapiccola ⁸⁴, M. Dam ³⁶, M. Dameri ^{50a,50b},
 D.S. Damiani ¹³⁷, H.O. Danielsson ³⁰, V. Dao ⁴⁹, G. Darbo ^{50a}, G.L. Darlea ^{26b}, J.A. Dassoulas ⁴², W. Davey ²¹,
 T. Davidek ¹²⁶, N. Davidson ⁸⁶, R. Davidson ⁷¹, E. Davies ^{118,c}, M. Davies ⁹³, O. Davignon ⁷⁸, A.R. Davison ⁷⁷,
 Y. Davygora ^{58a}, E. Dawe ¹⁴², I. Dawson ¹³⁹, R.K. Daya-Ishmukhametova ²³, K. De ⁸, R. de Asmundis ^{102a},
 S. De Castro ^{20a,20b}, S. De Cecco ⁷⁸, J. de Graat ⁹⁸, N. De Groot ¹⁰⁴, P. de Jong ¹⁰⁵, C. De La Taille ¹¹⁵,
 H. De la Torre ⁸⁰, F. De Lorenzi ⁶³, L. de Mora ⁷¹, L. De Nooij ¹⁰⁵, D. De Pedis ^{132a}, A. De Salvo ^{132a},
 U. De Sanctis ^{164a,164c}, A. De Santo ¹⁴⁹, J.B. De Vivie De Regie ¹¹⁵, G. De Zorzi ^{132a,132b}, W.J. Dearnaley ⁷¹,
 R. Debbe ²⁵, C. Debenedetti ⁴⁶, B. Dechenaux ⁵⁵, D.V. Dedovich ⁶⁴, J. Degenhardt ¹²⁰, C. Del Papa ^{164a,164c},
 J. Del Peso ⁸⁰, T. Del Prete ^{122a,122b}, T. Delemontex ⁵⁵, M. Deliyergiyev ⁷⁴, A. Dell'Acqua ³⁰, L. Dell'Asta ²²,
 M. Della Pietra ^{102a,j}, D. della Volpe ^{102a,102b}, M. Delmastro ⁵, P.A. Delsart ⁵⁵, C. Deluca ¹⁰⁵, S. Demers ¹⁷⁶,
 M. Demichev ⁶⁴, B. Demirkoz ^{12,l}, J. Deng ¹⁶³, S.P. Denisov ¹²⁸, D. Derendarz ³⁹, J.E. Derkaoui ^{135d},
 F. Derue ⁷⁸, P. Dervan ⁷³, K. Desch ²¹, E. Devetak ¹⁴⁸, P.O. Deviveiros ¹⁰⁵, A. Dewhurst ¹²⁹, B. DeWilde ¹⁴⁸,
 S. Dhaliwal ¹⁵⁸, R. Dhullipudi ^{25,m}, A. Di Ciaccio ^{133a,133b}, L. Di Ciaccio ⁵, A. Di Girolamo ³⁰,
 B. Di Girolamo ³⁰, S. Di Luise ^{134a,134b}, A. Di Mattia ¹⁷³, B. Di Micco ³⁰, R. Di Nardo ⁴⁷,
 A. Di Simone ^{133a,133b}, R. Di Sipio ^{20a,20b}, M.A. Diaz ^{32a}, E.B. Diehl ⁸⁷, J. Dietrich ⁴², T.A. Dietzsches ^{58a},
 S. Diglio ⁸⁶, K. Dindar Yagci ⁴⁰, J. Dingfelder ²¹, F. Dinut ^{26a}, C. Dionisi ^{132a,132b}, P. Dita ^{26a}, S. Dita ^{26a},
 F. Dittus ³⁰, F. Djama ⁸³, T. Djobava ^{51b}, M.A.B. do Vale ^{24c}, A. Do Valle Wemans ^{124a,n}, T.K.O. Doan ⁵,
 M. Dobbs ⁸⁵, R. Dobinson ^{30,*}, D. Dobos ³⁰, E. Dobson ^{30,o}, J. Dodd ³⁵, C. Doglioni ⁴⁹, T. Doherty ⁵³,
 Y. Doi ^{65,*}, J. Dolejsi ¹²⁶, I. Dolenc ⁷⁴, Z. Dolezal ¹²⁶, B.A. Dolgoshein ^{96,*}, T. Dohmae ¹⁵⁵, M. Donadelli ^{24d},
 J. Donini ³⁴, J. Dopke ³⁰, A. Doria ^{102a}, A. Dos Anjos ¹⁷³, A. Dotti ^{122a,122b}, M.T. Dova ⁷⁰, A.D. Doxiadis ¹⁰⁵,
 A.T. Doyle ⁵³, N. Dressnandt ¹²⁰, M. Dris ¹⁰, J. Dubbert ⁹⁹, S. Dube ¹⁵, E. Duchovni ¹⁷², G. Duckeck ⁹⁸,
 D. Duda ¹⁷⁵, A. Dudarev ³⁰, F. Dudziak ⁶³, M. Dührssen ³⁰, I.P. Duerdorff ⁸², L. Duflot ¹¹⁵, M.-A. Dufour ⁸⁵,

- L. Duguid ⁷⁶, M. Dunford ³⁰, H. Duran Yildiz ^{4a}, R. Duxfield ¹³⁹, M. Dwuznik ³⁸, F. Dydak ³⁰, M. Düren ⁵², W.L. Ebenstein ⁴⁵, J. Ebke ⁹⁸, S. Eckweiler ⁸¹, K. Edmonds ⁸¹, W. Edson ², C.A. Edwards ⁷⁶, N.C. Edwards ⁵³, W. Ehrenfeld ⁴², T. Eifert ¹⁴³, G. Eigen ¹⁴, K. Einsweiler ¹⁵, E. Eisenhandler ⁷⁵, T. Ekelof ¹⁶⁶, M. El Kacimi ^{135c}, M. Ellert ¹⁶⁶, S. Elles ⁵, F. Ellinghaus ⁸¹, K. Ellis ⁷⁵, N. Ellis ³⁰, J. Elmsheuser ⁹⁸, M. Elsing ³⁰, D. Emeliyanov ¹²⁹, R. Engelmann ¹⁴⁸, A. Engl ⁹⁸, B. Epp ⁶¹, J. Erdmann ⁵⁴, A. Ereditato ¹⁷, D. Eriksson ^{146a}, J. Ernst ², M. Ernst ²⁵, J. Ernwein ¹³⁶, D. Errede ¹⁶⁵, S. Errede ¹⁶⁵, E. Ertel ⁸¹, M. Escalier ¹¹⁵, H. Esch ⁴³, C. Escobar ¹²³, X. Espinal Curull ¹², B. Esposito ⁴⁷, F. Etienne ⁸³, A.I. Etienvre ¹³⁶, E. Etzion ¹⁵³, D. Evangelakou ⁵⁴, H. Evans ⁶⁰, L. Fabbri ^{20a,20b}, C. Fabre ³⁰, R.M. Fakhrutdinov ¹²⁸, S. Falciano ^{132a}, Y. Fang ¹⁷³, M. Fanti ^{89a,89b}, A. Farbin ⁸, A. Farilla ^{134a}, J. Farley ¹⁴⁸, T. Farooque ¹⁵⁸, S. Farrell ¹⁶³, S.M. Farrington ¹⁷⁰, P. Farthouat ³⁰, F. Fassi ¹⁶⁷, P. Fassnacht ³⁰, D. Fassouliotis ⁹, B. Fatholahzadeh ¹⁵⁸, A. Favareto ^{89a,89b}, L. Fayard ¹¹⁵, S. Fazio ^{37a,37b}, R. Febbraro ³⁴, P. Federic ^{144a}, O.L. Fedin ¹²¹, W. Fedorko ⁸⁸, M. Fehling-Kaschek ⁴⁸, L. Feligioni ⁸³, D. Fellmann ⁶, C. Feng ^{33d}, E.J. Feng ⁶, A.B. Fenyuk ¹²⁸, J. Ferencei ^{144b}, W. Fernando ⁶, S. Ferrag ⁵³, J. Ferrando ⁵³, V. Ferrara ⁴², A. Ferrari ¹⁶⁶, P. Ferrari ¹⁰⁵, R. Ferrari ^{119a}, D.E. Ferreira de Lima ⁵³, A. Ferrer ¹⁶⁷, D. Ferrere ⁴⁹, C. Ferretti ⁸⁷, A. Ferretto Parodi ^{50a,50b}, M. Fiascaris ³¹, F. Fiedler ⁸¹, A. Filipčič ⁷⁴, F. Filthaut ¹⁰⁴, M. Fincke-Keeler ¹⁶⁹, M.C.N. Fiolhais ^{124a,h}, L. Fiorini ¹⁶⁷, A. Firar ⁴⁰, G. Fischer ⁴², M.J. Fisher ¹⁰⁹, M. Flechl ⁴⁸, I. Fleck ¹⁴¹, J. Fleckner ⁸¹, P. Fleischmann ¹⁷⁴, S. Fleischmann ¹⁷⁵, T. Flick ¹⁷⁵, A. Floderus ⁷⁹, L.R. Flores Castillo ¹⁷³, M.J. Flowerdew ⁹⁹, T. Fonseca Martin ¹⁷, A. Formica ¹³⁶, A. Forti ⁸², D. Fortin ^{159a}, D. Fournier ¹¹⁵, A.J. Fowler ⁴⁵, H. Fox ⁷¹, P. Francavilla ¹², M. Franchini ^{20a,20b}, S. Franchino ^{119a,119b}, D. Francis ³⁰, T. Frank ¹⁷², S. Franz ³⁰, M. Fraternali ^{119a,119b}, S. Fratina ¹²⁰, S.T. French ²⁸, C. Friedrich ⁴², F. Friedrich ⁴⁴, R. Froeschl ³⁰, D. Froidevaux ³⁰, J.A. Frost ²⁸, C. Fukunaga ¹⁵⁶, E. Fullana Torregrosa ³⁰, B.G. Fulson ¹⁴³, J. Fuster ¹⁶⁷, C. Gabaldon ³⁰, O. Gabizon ¹⁷², T. Gadfort ²⁵, S. Gadomski ⁴⁹, G. Gagliardi ^{50a,50b}, P. Gagnon ⁶⁰, C. Galea ⁹⁸, B. Galhardo ^{124a}, E.J. Gallas ¹¹⁸, V. Gallo ¹⁷, B.J. Gallop ¹²⁹, P. Gallus ¹²⁵, K.K. Gan ¹⁰⁹, Y.S. Gao ^{143,e}, A. Gaponenko ¹⁵, F. Garberson ¹⁷⁶, M. Garcia-Sciveres ¹⁵, C. García ¹⁶⁷, J.E. García Navarro ¹⁶⁷, R.W. Gardner ³¹, N. Garelli ³⁰, H. Garitaonandia ¹⁰⁵, V. Garonne ³⁰, C. Gatti ⁴⁷, G. Gaudio ^{119a}, B. Gaur ¹⁴¹, L. Gauthier ¹³⁶, P. Gauzzi ^{132a,132b}, I.L. Gavrilenco ⁹⁴, C. Gay ¹⁶⁸, G. Gaycken ²¹, E.N. Gazis ¹⁰, P. Ge ^{33d}, Z. Gecse ¹⁶⁸, C.N.P. Gee ¹²⁹, D.A.A. Geerts ¹⁰⁵, Ch. Geich-Gimbel ²¹, K. Gellerstedt ^{146a,146b}, C. Gemme ^{50a}, A. Gemmell ⁵³, M.H. Genest ⁵⁵, S. Gentile ^{132a,132b}, M. George ⁵⁴, S. George ⁷⁶, P. Gerlach ¹⁷⁵, A. Gershon ¹⁵³, C. Geweniger ^{58a}, H. Ghazlane ^{135b}, N. Ghodbane ³⁴, B. Giacobbe ^{20a}, S. Giagu ^{132a,132b}, V. Giakoumopoulou ⁹, V. Giangiobbe ¹², F. Gianotti ³⁰, B. Gibbard ²⁵, A. Gibson ¹⁵⁸, S.M. Gibson ³⁰, M. Gilchriese ¹⁵, D. Gillberg ²⁹, A.R. Gillman ¹²⁹, D.M. Gingrich ^{3,d}, J. Ginzburg ¹⁵³, N. Giokaris ⁹, M.P. Giordani ^{164c}, R. Giordano ^{102a,102b}, F.M. Giorgi ¹⁶, P. Giovannini ⁹⁹, P.F. Giraud ¹³⁶, D. Giugni ^{89a}, M. Giunta ⁹³, P. Giusti ^{20a}, B.K. Gjelsten ¹¹⁷, L.K. Gladilin ⁹⁷, C. Glasman ⁸⁰, J. Glatzer ⁴⁸, A. Glazov ⁴², K.W. Glitza ¹⁷⁵, G.L. Glonti ⁶⁴, J.R. Goddard ⁷⁵, J. Godfrey ¹⁴², J. Godlewski ³⁰, M. Goebel ⁴², T. Göpfert ⁴⁴, C. Goeringer ⁸¹, C. Gössling ⁴³, S. Goldfarb ⁸⁷, T. Golling ¹⁷⁶, A. Gomes ^{124a,b}, L.S. Gomez Fajardo ⁴², R. Gonçalo ⁷⁶, J. Goncalves Pinto Firmino Da Costa ⁴², L. Gonella ²¹, S. Gonzalez ¹⁷³, S. González de la Hoz ¹⁶⁷, G. Gonzalez Parra ¹², M.L. Gonzalez Silva ²⁷, S. Gonzalez-Sevilla ⁴⁹, J.J. Goodson ¹⁴⁸, L. Goossens ³⁰, P.A. Gorbounov ⁹⁵, H.A. Gordon ²⁵, I. Gorelov ¹⁰³, G. Gorfine ¹⁷⁵, B. Gorini ³⁰, E. Gorini ^{72a,72b}, A. Gorišek ⁷⁴, E. Gornicki ³⁹, B. Gosdzik ⁴², A.T. Goshaw ⁶, M. Gosselink ¹⁰⁵, M.I. Gostkin ⁶⁴, I. Gough Eschrich ¹⁶³, M. Gouighri ^{135a}, D. Goujdami ^{135c}, M.P. Goulette ⁴⁹, A.G. Goussiou ¹³⁸, C. Goy ⁵, S. Gozpinar ²³, I. Grabowska-Bold ³⁸, P. Grafström ^{20a,20b}, K.-J. Grahn ⁴², F. Grancagnolo ^{72a}, S. Grancagnolo ¹⁶, V. Grassi ¹⁴⁸, V. Gratchev ¹²¹, N. Grau ³⁵, H.M. Gray ³⁰, J.A. Gray ¹⁴⁸, E. Graziani ^{134a}, O.G. Grebenyuk ¹²¹, T. Greenshaw ⁷³, Z.D. Greenwood ^{25,m}, K. Gregersen ³⁶, I.M. Gregor ⁴², P. Grenier ¹⁴³, J. Griffiths ⁸, N. Grigalashvili ⁶⁴, A.A. Grillo ¹³⁷, S. Grinstein ¹², Ph. Gris ³⁴, Y.V. Grishkevich ⁹⁷, J.-F. Grivaz ¹¹⁵, E. Gross ¹⁷², J. Grosse-Knetter ⁵⁴, J. Groth-Jensen ¹⁷², K. Grybel ¹⁴¹, D. Guest ¹⁷⁶, C. Guicheney ³⁴, S. Guindon ⁵⁴, U. Gul ⁵³, H. Guler ^{85,p}, J. Gunther ¹²⁵, B. Guo ¹⁵⁸, J. Guo ³⁵, P. Gutierrez ¹¹¹, N. Guttman ¹⁵³, O. Gutzwiller ¹⁷³, C. Guyot ¹³⁶, C. Gwenlan ¹¹⁸, C.B. Gwilliam ⁷³, A. Haas ¹⁴³, S. Haas ³⁰, C. Haber ¹⁵, H.K. Hadavand ⁴⁰, D.R. Hadley ¹⁸, P. Haefner ²¹, F. Hahn ³⁰, S. Haider ³⁰, Z. Hajduk ³⁹, H. Hakobyan ¹⁷⁷, D. Hall ¹¹⁸, J. Haller ⁵⁴, K. Hamacher ¹⁷⁵, P. Hamal ¹¹³, K. Hamano ⁸⁶, M. Hamer ⁵⁴, A. Hamilton ^{145b,q}, S. Hamilton ¹⁶¹, L. Han ^{33b}, K. Hanagaki ¹¹⁶, K. Hanawa ¹⁶⁰, M. Hance ¹⁵, C. Handel ⁸¹, P. Hanke ^{58a}, J.R. Hansen ³⁶, J.B. Hansen ³⁶, J.D. Hansen ³⁶, P.H. Hansen ³⁶, P. Hansson ¹⁴³, K. Hara ¹⁶⁰, G.A. Hare ¹³⁷, T. Harenberg ¹⁷⁵, S. Harkusha ⁹⁰, D. Harper ⁸⁷, R.D. Harrington ⁴⁶,

- O.M. Harris ¹³⁸, J. Hartert ⁴⁸, F. Hartjes ¹⁰⁵, T. Haruyama ⁶⁵, A. Harvey ⁵⁶, S. Hasegawa ¹⁰¹, Y. Hasegawa ¹⁴⁰, S. Hassani ¹³⁶, S. Haug ¹⁷, M. Hauschild ³⁰, R. Hauser ⁸⁸, M. Havranek ²¹, C.M. Hawkes ¹⁸, R.J. Hawkings ³⁰, A.D. Hawkins ⁷⁹, D. Hawkins ¹⁶³, T. Hayakawa ⁶⁶, T. Hayashi ¹⁶⁰, D. Hayden ⁷⁶, C.P. Hays ¹¹⁸, H.S. Hayward ⁷³, S.J. Haywood ¹²⁹, M. He ^{33d}, S.J. Head ¹⁸, V. Hedberg ⁷⁹, L. Heelan ⁸, S. Heim ⁸⁸, B. Heinemann ¹⁵, S. Heisterkamp ³⁶, L. Helary ²², C. Heller ⁹⁸, M. Heller ³⁰, S. Hellman ^{146a,146b}, D. Hellmich ²¹, C. Helsens ¹², R.C.W. Henderson ⁷¹, M. Henke ^{58a}, A. Henrichs ⁵⁴, A.M. Henriques Correia ³⁰, S. Henrot-Versille ¹¹⁵, C. Hensel ⁵⁴, T. Henß ¹⁷⁵, C.M. Hernandez ⁸, Y. Hernández Jiménez ¹⁶⁷, R. Herrberg ¹⁶, G. Herten ⁴⁸, R. Hertenberger ⁹⁸, L. Hervas ³⁰, G.G. Hesketh ⁷⁷, N.P. Hessey ¹⁰⁵, E. Higón-Rodriguez ¹⁶⁷, J.C. Hill ²⁸, K.H. Hiller ⁴², S. Hillert ²¹, S.J. Hillier ¹⁸, I. Hinchliffe ¹⁵, E. Hines ¹²⁰, M. Hirose ¹¹⁶, F. Hirsch ⁴³, D. Hirschbuehl ¹⁷⁵, J. Hobbs ¹⁴⁸, N. Hod ¹⁵³, M.C. Hodgkinson ¹³⁹, P. Hodgson ¹³⁹, A. Hoecker ³⁰, M.R. Hoeferkamp ¹⁰³, J. Hoffman ⁴⁰, D. Hoffmann ⁸³, M. Hohlfeld ⁸¹, M. Holder ¹⁴¹, S.O. Holmgren ^{146a}, T. Holy ¹²⁷, J.L. Holzbauer ⁸⁸, T.M. Hong ¹²⁰, L. Hooft van Huysduynen ¹⁰⁸, S. Horner ⁴⁸, J.-Y. Hostachy ⁵⁵, S. Hou ¹⁵¹, A. Hoummada ^{135a}, J. Howard ¹¹⁸, J. Howarth ⁸², I. Hristova ¹⁶, J. Hrvnac ¹¹⁵, T. Hryna'ova ⁵, P.J. Hsu ⁸¹, S.-C. Hsu ¹⁵, D. Hu ³⁵, Z. Hubacek ¹²⁷, F. Hubaut ⁸³, F. Huegging ²¹, A. Huettmann ⁴², T.B. Huffman ¹¹⁸, E.W. Hughes ³⁵, G. Hughes ⁷¹, M. Huhtinen ³⁰, M. Hurwitz ¹⁵, U. Husemann ⁴², N. Huseynov ^{64,r}, J. Huston ⁸⁸, J. Huth ⁵⁷, G. Iacobucci ⁴⁹, G. Iakovidis ¹⁰, M. Ibbotson ⁸², I. Ibragimov ¹⁴¹, L. Iconomidou-Fayard ¹¹⁵, J. Idarraga ¹¹⁵, P. Iengo ^{102a}, O. Igonkina ¹⁰⁵, Y. Ikegami ⁶⁵, M. Ikeno ⁶⁵, D. Iliadis ¹⁵⁴, N. Ilic ¹⁵⁸, T. Ince ²¹, J. Inigo-Golfin ³⁰, P. Ioannou ⁹, M. Iodice ^{134a}, K. Iordanidou ⁹, V. Ippolito ^{132a,132b}, A. Irles Quiles ¹⁶⁷, C. Isaksson ¹⁶⁶, M. Ishino ⁶⁷, M. Ishitsuka ¹⁵⁷, R. Ishmukhametov ⁴⁰, C. Issever ¹¹⁸, S. Istin ^{19a}, A.V. Ivashin ¹²⁸, W. Iwanski ³⁹, H. Iwasaki ⁶⁵, J.M. Izen ⁴¹, V. Izzo ^{102a}, B. Jackson ¹²⁰, J.N. Jackson ⁷³, P. Jackson ¹, M.R. Jaekel ³⁰, V. Jain ⁶⁰, K. Jakobs ⁴⁸, S. Jakobsen ³⁶, T. Jakoubek ¹²⁵, J. Jakubek ¹²⁷, D.K. Jana ¹¹¹, E. Jansen ⁷⁷, H. Jansen ³⁰, A. Jantsch ⁹⁹, M. Janus ⁴⁸, G. Jarlskog ⁷⁹, L. Jeanty ⁵⁷, I. Jen-La Plante ³¹, D. Jennens ⁸⁶, P. Jenni ³⁰, A.E. Loevschall-Jensen ³⁶, P. Jež ³⁶, S. Jézéquel ⁵, M.K. Jha ^{20a}, H. Ji ¹⁷³, W. Ji ⁸¹, J. Jia ¹⁴⁸, Y. Jiang ^{33b}, M. Jimenez Belenguer ⁴², S. Jin ^{33a}, O. Jinnouchi ¹⁵⁷, M.D. Joergensen ³⁶, D. Joffe ⁴⁰, M. Johansen ^{146a,146b}, K.E. Johansson ^{146a}, P. Johansson ¹³⁹, S. Johnert ⁴², K.A. Johns ⁷, K. Jon-And ^{146a,146b}, G. Jones ¹⁷⁰, R.W.L. Jones ⁷¹, T.J. Jones ⁷³, C. Joram ³⁰, P.M. Jorge ^{124a}, K.D. Joshi ⁸², J. Jovicevic ¹⁴⁷, T. Jovin ^{13b}, X. Ju ¹⁷³, C.A. Jung ⁴³, R.M. Jungst ³⁰, V. Juranek ¹²⁵, P. Jussel ⁶¹, A. Juste Rozas ¹², S. Kabana ¹⁷, M. Kaci ¹⁶⁷, A. Kaczmarśka ³⁹, P. Kadlecik ³⁶, M. Kado ¹¹⁵, H. Kagan ¹⁰⁹, M. Kagan ⁵⁷, E. Kajomovitz ¹⁵², S. Kalinin ¹⁷⁵, L.V. Kalinovskaya ⁶⁴, S. Kama ⁴⁰, N. Kanaya ¹⁵⁵, M. Kaneda ³⁰, S. Kaneti ²⁸, T. Kanno ¹⁵⁷, V.A. Kantserov ⁹⁶, J. Kanzaki ⁶⁵, B. Kaplan ¹⁰⁸, A. Kapliy ³¹, J. Kaplon ³⁰, D. Kar ⁵³, M. Karagounis ²¹, K. Karakostas ¹⁰, M. Karnevskiy ⁴², V. Kartvelishvili ⁷¹, A.N. Karyukhin ¹²⁸, L. Kashif ¹⁷³, G. Kasieczka ^{58b}, R.D. Kass ¹⁰⁹, A. Kastanas ¹⁴, M. Kataoka ⁵, Y. Kataoka ¹⁵⁵, E. Katsoufis ¹⁰, J. Katzy ⁴², V. Kaushik ⁷, K. Kawagoe ⁶⁹, T. Kawamoto ¹⁵⁵, G. Kawamura ⁸¹, M.S. Kayl ¹⁰⁵, S. Kazama ¹⁵⁵, V.A. Kazanin ¹⁰⁷, M.Y. Kazarinov ⁶⁴, R. Keeler ¹⁶⁹, P.T. Keener ¹²⁰, R. Kehoe ⁴⁰, M. Keil ⁵⁴, G.D. Kekelidze ⁶⁴, J.S. Keller ¹³⁸, M. Kenyon ⁵³, O. Kepka ¹²⁵, N. Kerschen ³⁰, B.P. Kerševan ⁷⁴, S. Kersten ¹⁷⁵, K. Kessoku ¹⁵⁵, J. Keung ¹⁵⁸, F. Khalil-zada ¹¹, H. Khandanyan ^{146a,146b}, A. Khanov ¹¹², D. Kharchenko ⁶⁴, A. Khodinov ⁹⁶, A. Khomich ^{58a}, T.J. Khoo ²⁸, G. Khoriauli ²¹, A. Khoroshilov ¹⁷⁵, V. Khovanskiy ⁹⁵, E. Khramov ⁶⁴, J. Khubua ^{51b}, H. Kim ^{146a,146b}, S.H. Kim ¹⁶⁰, N. Kimura ¹⁷¹, O. Kind ¹⁶, B.T. King ⁷³, M. King ⁶⁶, R.S.B. King ¹¹⁸, J. Kirk ¹²⁹, A.E. Kiryunin ⁹⁹, T. Kishimoto ⁶⁶, D. Kisielewska ³⁸, T. Kitamura ⁶⁶, T. Kittelmann ¹²³, K. Kiuchi ¹⁶⁰, E. Kladiva ^{144b}, M. Klein ⁷³, U. Klein ⁷³, K. Kleinknecht ⁸¹, M. Klemetti ⁸⁵, A. Klier ¹⁷², P. Klimek ^{146a,146b}, A. Klimentov ²⁵, R. Klingenberg ⁴³, J.A. Klinger ⁸², E.B. Klinkby ³⁶, T. Klioutchnikova ³⁰, P.F. Klok ¹⁰⁴, S. Kloous ¹⁰⁵, E.-E. Kluge ^{58a}, T. Kluge ⁷³, P. Kluit ¹⁰⁵, S. Kluth ⁹⁹, N.S. Knecht ¹⁵⁸, E. Kneringer ⁶¹, E.B.F.G. Knoops ⁸³, A. Knue ⁵⁴, B.R. Ko ⁴⁵, T. Kobayashi ¹⁵⁵, M. Kobel ⁴⁴, M. Kocian ¹⁴³, P. Kodys ¹²⁶, K. Köneke ³⁰, A.C. König ¹⁰⁴, S. Koenig ⁸¹, L. Köpke ⁸¹, F. Koetsveld ¹⁰⁴, P. Koevesarki ²¹, T. Koffas ²⁹, E. Koffeman ¹⁰⁵, L.A. Kogan ¹¹⁸, S. Kohlmann ¹⁷⁵, F. Kohn ⁵⁴, Z. Kohout ¹²⁷, T. Kohriki ⁶⁵, T. Koi ¹⁴³, G.M. Kolachev ^{107,*}, H. Kolanoski ¹⁶, V. Kolesnikov ⁶⁴, I. Koletsou ^{89a}, J. Koll ⁸⁸, M. Kollefrath ⁴⁸, A.A. Komar ⁹⁴, Y. Komori ¹⁵⁵, T. Kondo ⁶⁵, T. Kono ^{42,s}, A.I. Kononov ⁴⁸, R. Konoplich ^{108,t}, N. Konstantinidis ⁷⁷, S. Koperny ³⁸, K. Korcyl ³⁹, K. Kordas ¹⁵⁴, A. Korn ¹¹⁸, A. Korol ¹⁰⁷, I. Korolkov ¹², E.V. Korolkova ¹³⁹, V.A. Korotkov ¹²⁸, O. Kortner ⁹⁹, S. Kortner ⁹⁹, V.V. Kostyukhin ²¹, S. Kotov ⁹⁹, V.M. Kotov ⁶⁴, A. Kotwal ⁴⁵, C. Kourkoumelis ⁹, V. Kouskoura ¹⁵⁴, A. Koutsman ^{159a}, R. Kowalewski ¹⁶⁹, T.Z. Kowalski ³⁸, W. Kozanecki ¹³⁶, A.S. Kozhin ¹²⁸, V. Kral ¹²⁷, V.A. Kramarenko ⁹⁷, G. Kramberger ⁷⁴, M.W. Krasny ⁷⁸, A. Krasznahorkay ¹⁰⁸, J.K. Kraus ²¹,

- S. Kreiss ¹⁰⁸, F. Krejci ¹²⁷, J. Kretzschmar ⁷³, N. Krieger ⁵⁴, P. Krieger ¹⁵⁸, K. Kroeninger ⁵⁴, H. Kroha ⁹⁹,
 J. Kroll ¹²⁰, J. Kroseberg ²¹, J. Krstic ^{13a}, U. Kruchonak ⁶⁴, H. Krüger ²¹, T. Krucker ¹⁷, N. Krumnack ⁶³,
 Z.V. Krumshteyn ⁶⁴, T. Kubota ⁸⁶, S. Kuday ^{4a}, S. Kuehn ⁴⁸, A. Kugel ^{58c}, T. Kuhl ⁴², D. Kuhn ⁶¹, V. Kukhtin ⁶⁴,
 Y. Kulchitsky ⁹⁰, S. Kuleshov ^{32b}, C. Kummer ⁹⁸, M. Kuna ⁷⁸, J. Kunkle ¹²⁰, A. Kupco ¹²⁵, H. Kurashige ⁶⁶,
 M. Kurata ¹⁶⁰, Y.A. Kurochkin ⁹⁰, V. Kus ¹²⁵, E.S. Kuwertz ¹⁴⁷, M. Kuze ¹⁵⁷, J. Kvita ¹⁴², R. Kwee ¹⁶,
 A. La Rosa ⁴⁹, L. La Rotonda ^{37a,37b}, L. Labarga ⁸⁰, J. Labbe ⁵, S. Lablak ^{135a}, C. Lacasta ¹⁶⁷, F. Lacava ^{132a,132b},
 H. Lacker ¹⁶, D. Lacour ⁷⁸, V.R. Lacuesta ¹⁶⁷, E. Ladygin ⁶⁴, R. Lafaye ⁵, B. Laforge ⁷⁸, T. Lagouri ⁸⁰, S. Lai ⁴⁸,
 E. Laisne ⁵⁵, M. Lamanna ³⁰, L. Lambourne ⁷⁷, C.L. Lampen ⁷, W. Lampl ⁷, E. Lancon ¹³⁶, U. Landgraf ⁴⁸,
 M.P.J. Landon ⁷⁵, J.L. Lane ⁸², V.S. Lang ^{58a}, C. Lange ⁴², A.J. Lankford ¹⁶³, F. Lanni ²⁵, K. Lantzsch ¹⁷⁵,
 S. Laplace ⁷⁸, C. Lapoire ²¹, J.F. Laporte ¹³⁶, T. Lari ¹¹⁸, A. Larner ¹¹⁸, M. Lassnig ³⁰, P. Laurelli ⁴⁷,
 V. Lavorini ^{37a,37b}, W. Lavrijsen ¹⁵, P. Laycock ⁷³, O. Le Dortz ⁷⁸, E. Le Guiriec ⁸³, C. Le Maner ¹⁵⁸,
 E. Le Menedeu ¹², T. LeCompte ⁶, F. Ledroit-Guillon ⁵⁵, H. Lee ¹⁰⁵, J.S.H. Lee ¹¹⁶, S.C. Lee ¹⁵¹, L. Lee ¹⁷⁶,
 M. Lefebvre ¹⁶⁹, M. Legendre ¹³⁶, F. Legger ⁹⁸, C. Leggett ¹⁵, M. Lehmann Miotto ³⁰,
 X. Lei ⁷, M.A.L. Leite ^{24d}, R. Leitner ¹²⁶, D. Lellouch ¹⁷², B. Lemmer ⁵⁴, V. Lendermann ^{58a}, K.J.C. Leney ^{145b},
 T. Lenz ¹⁰⁵, G. Lenzen ¹⁷⁵, B. Lenzi ³⁰, K. Leonhardt ⁴⁴, S. Leontsinis ¹⁰, F. Lepold ^{58a}, C. Leroy ⁹³,
 J.-R. Lessard ¹⁶⁹, C.G. Lester ²⁸, C.M. Lester ¹²⁰, J. Levêque ⁵, D. Levin ⁸⁷, L.J. Levinson ¹⁷², A. Lewis ¹¹⁸,
 G.H. Lewis ¹⁰⁸, A.M. Leyko ²¹, M. Leyton ¹⁶, B. Li ⁸³, H. Li ^{173,u}, S. Li ^{33b,v}, X. Li ⁸⁷, Z. Liang ^{118,w}, H. Liao ³⁴,
 B. Liberti ^{133a}, P. Lichard ³⁰, M. Lichtnecker ⁹⁸, K. Lie ¹⁶⁵, W. Liebig ¹⁴, C. Limbach ²¹, A. Limosani ⁸⁶,
 M. Limper ⁶², S.C. Lin ^{151,x}, F. Linde ¹⁰⁵, J.T. Linnemann ⁸⁸, E. Lipeles ¹²⁰, A. Lipniacka ¹⁴, T.M. Liss ¹⁶⁵,
 D. Lissauer ²⁵, A. Lister ⁴⁹, A.M. Litke ¹³⁷, C. Liu ²⁹, D. Liu ¹⁵¹, H. Liu ⁸⁷, J.B. Liu ⁸⁷, L. Liu ⁸⁷, M. Liu ^{33b},
 Y. Liu ^{33b}, M. Livan ^{119a,119b}, S.S.A. Livermore ¹¹⁸, A. Lleres ⁵⁵, J. Llorente Merino ⁸⁰, S.L. Lloyd ⁷⁵,
 E. Lobodzinska ⁴², P. Loch ⁷, W.S. Lockman ¹³⁷, T. Loddenkoetter ²¹, F.K. Loebinger ⁸², A. Loginov ¹⁷⁶,
 C.W. Loh ¹⁶⁸, T. Lohse ¹⁶, K. Lohwasser ⁴⁸, M. Lokajicek ¹²⁵, V.P. Lombardo ⁵, R.E. Long ⁷¹, L. Lopes ^{124a},
 D. Lopez Mateos ⁵⁷, J. Lorenz ⁹⁸, N. Lorenzo Martinez ¹¹⁵, M. Losada ¹⁶², P. Loscutoff ¹⁵,
 F. Lo Sterzo ^{132a,132b}, M.J. Losty ^{159a,*}, X. Lou ⁴¹, A. Lounis ¹¹⁵, K.F. Loureiro ¹⁶², J. Love ⁶, P.A. Love ⁷¹,
 A.J. Lowe ^{143,e}, F. Lu ^{33a}, H.J. Lubatti ¹³⁸, C. Luci ^{132a,132b}, A. Lucotte ⁵⁵, A. Ludwig ⁴⁴, D. Ludwig ⁴²,
 I. Ludwig ⁴⁸, J. Ludwig ⁴⁸, F. Luehring ⁶⁰, G. Luijckx ¹⁰⁵, W. Lukas ⁶¹, D. Lumb ⁴⁸, L. Luminari ^{132a},
 E. Lund ¹¹⁷, B. Lund-Jensen ¹⁴⁷, B. Lundberg ⁷⁹, J. Lundberg ^{146a,146b}, O. Lundberg ^{146a,146b}, J. Lundquist ³⁶,
 M. Lungwitz ⁸¹, D. Lynn ²⁵, E. Lytken ⁷⁹, H. Ma ²⁵, LL. Ma ¹⁷³, G. Maccarrone ⁴⁷, A. Macchiolo ⁹⁹,
 B. Maček ⁷⁴, J. Machado Miguens ^{124a}, R. Mackeprang ³⁶, R.J. Madaras ¹⁵, H.J. Maddocks ⁷¹, W.F. Mader ⁴⁴,
 R. Maenner ^{58c}, T. Maeno ²⁵, P. Mättig ¹⁷⁵, S. Mättig ⁸¹, L. Magnoni ¹⁶³, E. Magradze ⁵⁴, K. Mahboubi ⁴⁸,
 S. Mahmoud ⁷³, G. Mahout ¹⁸, C. Maiani ¹³⁶, C. Maidantchik ^{24a}, A. Maio ^{124a,b}, S. Majewski ²⁵,
 Y. Makida ⁶⁵, N. Makovec ¹¹⁵, P. Mal ¹³⁶, B. Malaescu ³⁰, Pa. Malecki ³⁹, P. Malecki ³⁹, V.P. Maleev ¹²¹,
 F. Malek ⁵⁵, U. Mallik ⁶², D. Malon ⁶, C. Malone ¹⁴³, S. Maltezos ¹⁰, V. Malyshev ¹⁰⁷, S. Malyukov ³⁰,
 R. Mameghani ⁹⁸, J. Mamuzic ^{13b}, A. Manabe ⁶⁵, L. Mandelli ^{89a}, I. Mandić ⁷⁴, R. Mandrysch ¹⁶,
 J. Maneira ^{124a}, A. Manfredini ⁹⁹, P.S. Mangeard ⁸⁸, L. Manhaes de Andrade Filho ^{24b},
 J.A. Manjarres Ramos ¹³⁶, A. Mann ⁵⁴, P.M. Manning ¹³⁷, A. Manousakis-Katsikakis ⁹, B. Mansoulie ¹³⁶,
 A. Mapelli ³⁰, L. Mapelli ³⁰, L. March ⁸⁰, J.F. Marchand ²⁹, F. Marchese ^{133a,133b}, G. Marchiori ⁷⁸,
 M. Marcisovsky ¹²⁵, C.P. Marino ¹⁶⁹, F. Marroquim ^{24a}, Z. Marshall ³⁰, F.K. Martens ¹⁵⁸, L.F. Marti ¹⁷,
 S. Marti-Garcia ¹⁶⁷, B. Martin ³⁰, B. Martin ⁸⁸, J.P. Martin ⁹³, T.A. Martin ¹⁸, V.J. Martin ⁴⁶,
 B. Martin dit Latour ⁴⁹, S. Martin-Haugh ¹⁴⁹, M. Martinez ¹², V. Martinez Outschoorn ⁵⁷,
 A.C. Martyniuk ¹⁶⁹, M. Marx ⁸², F. Marzano ^{132a}, A. Marzin ¹¹¹, L. Masetti ⁸¹, T. Mashimo ¹⁵⁵,
 R. Mashinistov ⁹⁴, J. Masik ⁸², A.L. Maslenikov ¹⁰⁷, I. Massa ^{20a,20b}, G. Massaro ¹⁰⁵, N. Massol ⁵,
 P. Mastrandrea ¹⁴⁸, A. Mastroberardino ^{37a,37b}, T. Masubuchi ¹⁵⁵, P. Matricon ¹¹⁵, H. Matsunaga ¹⁵⁵,
 T. Matsushita ⁶⁶, C. Mattravers ^{118,c}, J. Maurer ⁸³, S.J. Maxfield ⁷³, A. Mayne ¹³⁹, R. Mazini ¹⁵¹, M. Mazur ²¹,
 L. Mazzaferro ^{133a,133b}, M. Mazzanti ^{89a}, J. Mc Donald ⁸⁵, S.P. Mc Kee ⁸⁷, A. McCarn ¹⁶⁵, R.L. McCarthy ¹⁴⁸,
 T.G. McCarthy ²⁹, N.A. McCubbin ¹²⁹, K.W. McFarlane ^{56,*}, J.A. McFayden ¹³⁹, G. Mchedlidze ^{51b},
 T. McLaughlan ¹⁸, S.J. McMahon ¹²⁹, R.A. McPherson ^{169,k}, A. Meade ⁸⁴, J. Mechlich ¹⁰⁵, M. Mechtel ¹⁷⁵,
 M. Medinnis ⁴², R. Meera-Lebbai ¹¹¹, T. Meguro ¹¹⁶, R. Mehdiyev ⁹³, S. Mehlhase ³⁶, A. Mehta ⁷³,
 K. Meier ^{58a}, B. Meirose ⁷⁹, C. Melachrinos ³¹, B.R. Mellado Garcia ¹⁷³, F. Meloni ^{89a,89b},
 L. Mendoza Navas ¹⁶², Z. Meng ^{151,u}, A. Mengarelli ^{20a,20b}, S. Menke ⁹⁹, E. Meoni ¹⁶¹, K.M. Mercurio ⁵⁷,
 P. Mermod ⁴⁹, L. Merola ^{102a,102b}, C. Meroni ^{89a}, F.S. Merritt ³¹, H. Merritt ¹⁰⁹, A. Messina ^{30,y},

- J. Metcalfe²⁵, A.S. Mete¹⁶³, C. Meyer⁸¹, C. Meyer³¹, J.-P. Meyer¹³⁶, J. Meyer¹⁷⁴, J. Meyer⁵⁴, T.C. Meyer³⁰, J. Miao^{33d}, S. Michal³⁰, L. Micu^{26a}, R.P. Middleton¹²⁹, S. Migas⁷³, L. Mijović¹³⁶, G. Mikenberg¹⁷², M. Mikestikova¹²⁵, M. Mikuž⁷⁴, D.W. Miller³¹, R.J. Miller⁸⁸, W.J. Mills¹⁶⁸, C. Mills⁵⁷, A. Milov¹⁷², D.A. Milstead^{146a,146b}, D. Milstein¹⁷², A.A. Minaenko¹²⁸, M. Miñano Moya¹⁶⁷, I.A. Minashvili⁶⁴, A.I. Mincer¹⁰⁸, B. Mindur³⁸, M. Mineev⁶⁴, Y. Ming¹⁷³, L.M. Mir¹², G. Mirabelli^{132a}, J. Mitrevski¹³⁷, V.A. Mitsou¹⁶⁷, S. Mitsui⁶⁵, P.S. Miyagawa¹³⁹, J.U. Mjörnmark⁷⁹, T. Moa^{146a,146b}, V. Moeller²⁸, K. Möning⁴², N. Möser²¹, S. Mohapatra¹⁴⁸, W. Mohr⁴⁸, R. Moles-Valls¹⁶⁷, A. Molfetas³⁰, J. Monk⁷⁷, E. Monnier⁸³, J. Montejo Berlingen¹², F. Monticelli⁷⁰, S. Monzani^{20a,20b}, R.W. Moore³, G.F. Moorhead⁸⁶, C. Mora Herrera⁴⁹, A. Moraes⁵³, N. Morange¹³⁶, J. Morel⁵⁴, G. Morello^{37a,37b}, D. Moreno⁸¹, M. Moreno Llácer¹⁶⁷, P. Morettini^{50a}, M. Morgenstern⁴⁴, M. Morii⁵⁷, A.K. Morley³⁰, G. Mornacchi³⁰, J.D. Morris⁷⁵, L. Morvaj¹⁰¹, H.G. Moser⁹⁹, M. Mosidze^{51b}, J. Moss¹⁰⁹, R. Mount¹⁴³, E. Mountricha^{10,z}, S.V. Mouraviev^{94,*}, E.J.W. Moyse⁸⁴, F. Mueller^{58a}, J. Mueller¹²³, K. Mueller²¹, T.A. Müller⁹⁸, T. Mueller⁸¹, D. Muenstermann³⁰, Y. Munwes¹⁵³, W.J. Murray¹²⁹, I. Mussche¹⁰⁵, E. Musto^{102a,102b}, A.G. Myagkov¹²⁸, M. Myska¹²⁵, J. Nadal¹², K. Nagai¹⁶⁰, R. Nagai¹⁵⁷, K. Nagano⁶⁵, A. Nagarkar¹⁰⁹, Y. Nagasaka⁵⁹, M. Nagel⁹⁹, A.M. Nairz³⁰, Y. Nakahama³⁰, K. Nakamura¹⁵⁵, T. Nakamura¹⁵⁵, I. Nakano¹¹⁰, G. Nanava²¹, A. Napier¹⁶¹, R. Narayan^{58b}, M. Nash^{77,c}, T. Nattermann²¹, T. Naumann⁴², G. Navarro¹⁶², H.A. Neal⁸⁷, P.Yu. Nechaeva⁹⁴, T.J. Neep⁸², A. Negri^{119a,119b}, G. Negri³⁰, M. Negrini^{20a}, S. Nektarijevic⁴⁹, A. Nelson¹⁶³, T.K. Nelson¹⁴³, S. Nemecek¹²⁵, P. Nemethy¹⁰⁸, A.A. Nepomuceno^{24a}, M. Nessi^{30,aa}, M.S. Neubauer¹⁶⁵, M. Neumann¹⁷⁵, A. Neusiedl⁸¹, R.M. Neves¹⁰⁸, P. Nevski²⁵, F.M. Newcomer¹²⁰, P.R. Newman¹⁸, V. Nguyen Thi Hong¹³⁶, R.B. Nickerson¹¹⁸, R. Nicolaïdou¹³⁶, B. Nicquevert³⁰, F. Niedercorn¹¹⁵, J. Nielsen¹³⁷, N. Nikiforou³⁵, A. Nikiforov¹⁶, V. Nikolaenko¹²⁸, I. Nikolic-Audit⁷⁸, K. Nikolics⁴⁹, K. Nikolopoulos¹⁸, H. Nilsen⁴⁸, P. Nilsson⁸, Y. Ninomiya¹⁵⁵, A. Nisati^{132a}, R. Nisius⁹⁹, T. Nobe¹⁵⁷, L. Nodulman⁶, M. Nomachi¹¹⁶, I. Nomidis¹⁵⁴, S. Norberg¹¹¹, M. Nordberg³⁰, P.R. Norton¹²⁹, J. Novakova¹²⁶, M. Nozaki⁶⁵, L. Nozka¹¹³, I.M. Nugent^{159a}, A.-E. Nuncio-Quiroz²¹, G. Nunes Hanninger⁸⁶, T. Nunnemann⁹⁸, E. Nurse⁷⁷, B.J. O'Brien⁴⁶, S.W. O'Neale^{18,*}, D.C. O'Neil¹⁴², V. O'Shea⁵³, L.B. Oakes⁹⁸, F.G. Oakham^{29,d}, H. Oberlack⁹⁹, J. Ocariz⁷⁸, A. Ochi⁶⁶, S. Oda⁶⁹, S. Odaka⁶⁵, J. Odier⁸³, H. Ogren⁶⁰, A. Oh⁸², S.H. Oh⁴⁵, C.C. Ohm³⁰, T. Ohshima¹⁰¹, H. Okawa²⁵, Y. Okumura³¹, T. Okuyama¹⁵⁵, A. Olariu^{26a}, A.G. Olchevski⁶⁴, S.A. Olivares Pino^{32a}, M. Oliveira^{124a,h}, D. Oliveira Damazio²⁵, E. Oliver Garcia¹⁶⁷, D. Olivito¹²⁰, A. Olszewski³⁹, J. Olszowska³⁹, A. Onofre^{124a,ab}, P.U.E. Onyisi³¹, C.J. Oram^{159a}, M.J. Oreglia³¹, Y. Oren¹⁵³, D. Orestano^{134a,134b}, N. Orlando^{72a,72b}, I. Orlov¹⁰⁷, C. Oropeza Barrera⁵³, R.S. Orr¹⁵⁸, B. Osculati^{50a,50b}, R. Ospanov¹²⁰, C. Osuna¹², G. Otero y Garzon²⁷, J.P. Ottersbach¹⁰⁵, M. Ouchrif^{135d}, E.A. Ouellette¹⁶⁹, F. Ould-Saada¹¹⁷, A. Ouraou¹³⁶, Q. Ouyang^{33a}, A. Ovcharova¹⁵, M. Owen⁸², S. Owen¹³⁹, V.E. Ozcan^{19a}, N. Ozturk⁸, A. Pacheco Pages¹², C. Padilla Aranda¹², S. Pagan Griso¹⁵, E. Paganis¹³⁹, C. Pahl⁹⁹, F. Paige²⁵, P. Pais⁸⁴, K. Pajchel¹¹⁷, G. Palacino^{159b}, C.P. Paleari⁷, S. Palestini³⁰, D. Pallin³⁴, A. Palma^{124a}, J.D. Palmer¹⁸, Y.B. Pan¹⁷³, E. Panagiotopoulou¹⁰, P. Pani¹⁰⁵, N. Panikashvili⁸⁷, S. Panitkin²⁵, D. Pantea^{26a}, A. Papadelis^{146a}, Th.D. Papadopoulou¹⁰, A. Paramonov⁶, D. Paredes Hernandez³⁴, W. Park^{25,ac}, M.A. Parker²⁸, F. Parodi^{50a,50b}, J.A. Parsons³⁵, U. Parzefall⁴⁸, S. Pashapour⁵⁴, E. Pasqualucci^{132a}, S. Passaggio^{50a}, A. Passeri^{134a}, F. Pastore^{134a,134b,*}, Fr. Pastore⁷⁶, G. Pásztor^{49,ad}, S. Pataraia¹⁷⁵, N. Patel¹⁵⁰, J.R. Pater⁸², S. Patricelli^{102a,102b}, T. Pauly³⁰, M. Pecsy^{144a}, S. Pedraza Lopez¹⁶⁷, M.I. Pedraza Morales¹⁷³, S.V. Peleganchuk¹⁰⁷, D. Pelikan¹⁶⁶, H. Peng^{33b}, B. Penning³¹, A. Penson³⁵, J. Penwell⁶⁰, M. Perantoni^{24a}, K. Perez^{35,ae}, T. Perez Cavalcanti⁴², E. Perez Codina^{159a}, M.T. Pérez García-Estañ¹⁶⁷, V. Perez Reale³⁵, L. Perini^{89a,89b}, H. Pernegger³⁰, R. Perrino^{72a}, P. Perrodo⁵, V.D. Peshekhonov⁶⁴, K. Peters³⁰, B.A. Petersen³⁰, J. Petersen³⁰, T.C. Petersen³⁶, E. Petit⁵, A. Petridis¹⁵⁴, C. Petridou¹⁵⁴, E. Petrolo^{132a}, F. Petrucci^{134a,134b}, D. Petschull⁴², M. Petteni¹⁴², R. Pezoa^{32b}, A. Phan⁸⁶, P.W. Phillips¹²⁹, G. Piacquadio³⁰, A. Picazio⁴⁹, E. Piccaro⁷⁵, M. Piccinini^{20a,20b}, S.M. Piec⁴², R. Piegala²⁷, D.T. Pignotti¹⁰⁹, J.E. Pilcher³¹, A.D. Pilkington⁸², J. Pina^{124a,b}, M. Pinamonti^{164a,164c}, A. Pinder¹¹⁸, J.L. Pinfold³, B. Pinto^{124a}, C. Pizio^{89a,89b}, M. Plamondon¹⁶⁹, M.-A. Pleier²⁵, E. Plotnikova⁶⁴, A. Poblaguev²⁵, S. Poddar^{58a}, F. Podlyski³⁴, L. Poggioli¹¹⁵, D. Pohl²¹, M. Pohl⁴⁹, G. Polesello^{119a}, A. Policicchio^{37a,37b}, A. Polini^{20a}, J. Poll⁷⁵, V. Polychronakos²⁵, D. Pomeroy²³, K. Pommès³⁰, L. Pontecorvo^{132a}, B.G. Pope⁸⁸, G.A. Popeneciu^{26a}, D.S. Popovic^{13a}, A. Poppleton³⁰, X. Portell Bueso³⁰, G.E. Pospelov⁹⁹, S. Pospisil¹²⁷,

- I.N. Potrap⁹⁹, C.J. Potter¹⁴⁹, C.T. Potter¹¹⁴, G. Poulard³⁰, J. Poveda⁶⁰, V. Pozdnyakov⁶⁴, R. Prabhu⁷⁷, P. Pralavorio⁸³, A. Pranko¹⁵, S. Prasad³⁰, R. Pravahan²⁵, S. Prell⁶³, K. Pretzl¹⁷, D. Price⁶⁰, J. Price⁷³, L.E. Price⁶, D. Prieur¹²³, M. Primavera^{72a}, K. Prokofiev¹⁰⁸, F. Prokoshin^{32b}, S. Protopopescu²⁵, J. Proudfoot⁶, X. Prudent⁴⁴, M. Przybycien³⁸, H. Przysiezniak⁵, S. Psoroulas²¹, E. Ptacek¹¹⁴, E. Pueschel⁸⁴, J. Purdham⁸⁷, M. Purohit^{25,ac}, P. Puzo¹¹⁵, Y. Pylypchenko⁶², J. Qian⁸⁷, A. Quadt⁵⁴, D.R. Quarrie¹⁵, W.B. Quayle¹⁷³, F. Quinonez^{32a}, M. Raas¹⁰⁴, V. Radeka²⁵, V. Radescu⁴², P. Radloff¹¹⁴, T. Rador^{19a}, F. Ragusa^{89a,89b}, G. Rahal¹⁷⁸, A.M. Rahimi¹⁰⁹, D. Rahm²⁵, S. Rajagopalan²⁵, M. Rammensee⁴⁸, M. Rammes¹⁴¹, A.S. Randle-Conde⁴⁰, K. Randrianarivony²⁹, F. Rauscher⁹⁸, T.C. Rave⁴⁸, M. Raymond³⁰, A.L. Read¹¹⁷, D.M. Rebuzzi^{119a,119b}, A. Redelbach¹⁷⁴, G. Redlinger²⁵, R. Reece¹²⁰, K. Reeves⁴¹, E. Reinherz-Aronis¹⁵³, A. Reinsch¹¹⁴, I. Reisinger⁴³, C. Rembser³⁰, Z.L. Ren¹⁵¹, A. Renaud¹¹⁵, M. Rescigno^{132a}, S. Resconi^{89a}, B. Resende¹³⁶, P. Reznicek⁹⁸, R. Rezvani¹⁵⁸, R. Richter⁹⁹, E. Richter-Was^{5,af}, M. Ridel⁷⁸, M. Rijpstra¹⁰⁵, M. Rijssenbeek¹⁴⁸, A. Rimoldi^{119a,119b}, L. Rinaldi^{20a}, R.R. Rios⁴⁰, I. Riu¹², G. Rivoltella^{89a,89b}, F. Rizatdinova¹¹², E. Rizvi⁷⁵, S.H. Robertson^{85,k}, A. Robichaud-Veronneau¹¹⁸, D. Robinson²⁸, J.E.M. Robinson⁸², A. Robson⁵³, J.G. Rocha de Lima¹⁰⁶, C. Roda^{122a,122b}, D. Roda Dos Santos³⁰, A. Roe⁵⁴, S. Roe³⁰, O. Røhne¹¹⁷, S. Rolli¹⁶¹, A. Romanikou⁹⁶, M. Romano^{20a,20b}, G. Romeo²⁷, E. Romero Adam¹⁶⁷, N. Rompotis¹³⁸, L. Roos⁷⁸, E. Ros¹⁶⁷, S. Rosati^{132a}, K. Rosbach⁴⁹, A. Rose¹⁴⁹, M. Rose⁷⁶, G.A. Rosenbaum¹⁵⁸, E.I. Rosenberg⁶³, P.L. Rosendahl¹⁴, O. Rosenthal¹⁴¹, L. Rosselet⁴⁹, V. Rossetti¹², E. Rossi^{132a,132b}, L.P. Rossi^{50a}, M. Rotaru^{26a}, I. Roth¹⁷², J. Rothberg¹³⁸, D. Rousseau¹¹⁵, C.R. Royon¹³⁶, A. Rozanov⁸³, Y. Rozen¹⁵², X. Ruan^{33a,ag}, F. Rubbo¹², I. Rubinskiy⁴², N. Ruckstuhl¹⁰⁵, V.I. Rud⁹⁷, C. Rudolph⁴⁴, G. Rudolph⁶¹, F. Rühr⁷, A. Ruiz-Martinez⁶³, L. Rumyantsev⁶⁴, Z. Rurikova⁴⁸, N.A. Rusakovich⁶⁴, J.P. Rutherford⁷, C. Ruwiedel^{15,*}, P. Ruzicka¹²⁵, Y.F. Ryabov¹²¹, M. Rybar¹²⁶, G. Rybkin¹¹⁵, N.C. Ryder¹¹⁸, A.F. Saavedra¹⁵⁰, I. Sadeh¹⁵³, H.F-W. Sadrozinski¹³⁷, R. Sadykov⁶⁴, F. Safai Tehrani^{132a}, H. Sakamoto¹⁵⁵, G. Salamanna⁷⁵, A. Salamon^{133a}, M. Saleem¹¹¹, D. Salek³⁰, D. Salihagic⁹⁹, A. Salnikov¹⁴³, J. Salt¹⁶⁷, B.M. Salvachua Ferrando⁶, D. Salvatore^{37a,37b}, F. Salvatore¹⁴⁹, A. Salvucci¹⁰⁴, A. Salzburger³⁰, D. Sampsonidis¹⁵⁴, B.H. Samset¹¹⁷, A. Sanchez^{102a,102b}, V. Sanchez Martinez¹⁶⁷, H. Sandaker¹⁴, H.G. Sander⁸¹, M.P. Sanders⁹⁸, M. Sandhoff¹⁷⁵, T. Sandoval²⁸, C. Sandoval¹⁶², R. Sandstroem⁹⁹, D.P.C. Sankey¹²⁹, A. Sansoni⁴⁷, C. Santamarina Rios⁸⁵, C. Santoni³⁴, R. Santonico^{133a,133b}, H. Santos^{124a}, J.G. Saraiva^{124a}, T. Sarangi¹⁷³, E. Sarkisyan-Grinbaum⁸, F. Sarri^{122a,122b}, G. Sartisohn¹⁷⁵, O. Sasaki⁶⁵, Y. Sasaki¹⁵⁵, N. Sasa⁶⁷, I. Satsounkevitch⁹⁰, G. Sauvage^{5,*}, E. Sauvan⁵, J.B. Sauvan¹¹⁵, P. Savard^{158,d}, V. Savinov¹²³, D.O. Savu³⁰, L. Sawyer^{25,m}, D.H. Saxon⁵³, J. Saxon¹²⁰, C. Sbarra^{20a}, A. Sbrizzi^{20a,20b}, D.A. Scannicchio¹⁶³, M. Scarcella¹⁵⁰, J. Schaarschmidt¹¹⁵, P. Schacht⁹⁹, D. Schaefer¹²⁰, U. Schäfer⁸¹, S. Schaepe²¹, S. Schaetzl^{58b}, A.C. Schaffer¹¹⁵, D. Schaile⁹⁸, R.D. Schamberger¹⁴⁸, A.G. Schamov¹⁰⁷, V. Scharf^{58a}, V.A. Schegelsky¹²¹, D. Scheirich⁸⁷, M. Schernau¹⁶³, M.I. Scherzer³⁵, C. Schiavi^{50a,50b}, J. Schieck⁹⁸, M. Schioppa^{37a,37b}, S. Schlenker³⁰, E. Schmidt⁴⁸, K. Schmieden²¹, C. Schmitt⁸¹, S. Schmitt^{58b}, M. Schmitz²¹, B. Schneider¹⁷, U. Schnoor⁴⁴, A. Schoening^{58b}, A.L.S. Schorlemmer⁵⁴, M. Schott³⁰, D. Schouten^{159a}, J. Schovancova¹²⁵, M. Schram⁸⁵, C. Schroeder⁸¹, N. Schroer^{58c}, M.J. Schultens²¹, J. Schultes¹⁷⁵, H.-C. Schultz-Coulon^{58a}, H. Schulz¹⁶, M. Schumacher⁴⁸, B.A. Schumm¹³⁷, Ph. Schune¹³⁶, C. Schwanenberger⁸², A. Schwartzman¹⁴³, Ph. Schwegler⁹⁹, Ph. Schwemling⁷⁸, R. Schwienhorst⁸⁸, R. Schwierz⁴⁴, J. Schwindling¹³⁶, T. Schwindt²¹, M. Schwoerer⁵, G. Sciolla²³, W.G. Scott¹²⁹, J. Searcy¹¹⁴, G. Sedov⁴², E. Sedykh¹²¹, S.C. Seidel¹⁰³, A. Seiden¹³⁷, F. Seifert⁴⁴, J.M. Seixas^{24a}, G. Sekhniaidze^{102a}, S.J. Sekula⁴⁰, K.E. Selbach⁴⁶, D.M. Seliverstov¹²¹, B. Sellden^{146a}, G. Sellers⁷³, M. Seman^{144b}, N. Semprini-Cesari^{20a,20b}, C. Serfon⁹⁸, L. Serin¹¹⁵, L. Serkin⁵⁴, R. Seuster⁹⁹, H. Severini¹¹¹, A. Sfyrla³⁰, E. Shabalina⁵⁴, M. Shamim¹¹⁴, L.Y. Shan^{33a}, J.T. Shank²², Q.T. Shao⁸⁶, M. Shapiro¹⁵, P.B. Shatalov⁹⁵, K. Shaw^{164a,164c}, D. Sherman¹⁷⁶, P. Sherwood⁷⁷, A. Shibata¹⁰⁸, S. Shimizu¹⁰¹, M. Shimojima¹⁰⁰, T. Shin⁵⁶, M. Shiyakova⁶⁴, A. Shmeleva⁹⁴, M.J. Shochet³¹, D. Short¹¹⁸, S. Shrestha⁶³, E. Shulga⁹⁶, M.A. Shupe⁷, P. Sicho¹²⁵, A. Sidoti^{132a}, F. Siegert⁴⁸, Dj. Sijacki^{13a}, O. Silbert¹⁷², J. Silva^{124a}, Y. Silver¹⁵³, D. Silverstein¹⁴³, S.B. Silverstein^{146a}, V. Simak¹²⁷, O. Simard¹³⁶, Lj. Simic^{13a}, S. Simion¹¹⁵, E. Simioni⁸¹, B. Simmons⁷⁷, R. Simonello^{89a,89b}, M. Simonyan³⁶, P. Sinervo¹⁵⁸, N.B. Sinev¹¹⁴, V. Sipica¹⁴¹, G. Siragusa¹⁷⁴, A. Sircar²⁵, A.N. Sisakyan^{64,*}, S.Yu. Sivoklokov⁹⁷, J. Sjölin^{146a,146b}, T.B. Sjursen¹⁴, L.A. Skinnari¹⁵, H.P. Skottowe⁵⁷, K. Skovpen¹⁰⁷, P. Skubic¹¹¹, M. Slater¹⁸, T. Slavicek¹²⁷, K. Sliwa¹⁶¹, V. Smakhtin¹⁷², B.H. Smart⁴⁶, L. Smestad¹¹⁷,

- S.Yu. Smirnov⁹⁶, Y. Smirnov⁹⁶, L.N. Smirnova⁹⁷, O. Smirnova⁷⁹, B.C. Smith⁵⁷, D. Smith¹⁴³,
 K.M. Smith⁵³, M. Smizanska⁷¹, K. Smolek¹²⁷, A.A. Snesarev⁹⁴, S.W. Snow⁸², J. Snow¹¹¹, S. Snyder²⁵,
 R. Sobie^{169,k}, J. Sodomka¹²⁷, A. Soffer¹⁵³, C.A. Solans¹⁶⁷, M. Solar¹²⁷, J. Solc¹²⁷, E.Yu. Soldatov⁹⁶,
 U. Soldevila¹⁶⁷, E. Solfaroli Camillocci^{132a,132b}, A.A. Solodkov¹²⁸, O.V. Solovyev¹²⁸, V. Solovyev¹²¹,
 N. Soni¹, V. Sopko¹²⁷, B. Sopko¹²⁷, M. Sosebee⁸, R. Soualah^{164a,164c}, A. Soukharev¹⁰⁷
 S. Spagnolo^{72a,72b}, F. Spanò⁷⁶, R. Spighi^{20a}, G. Spigo³⁰, R. Spiwoks³⁰, M. Spousta^{126,ah}, T. Spreitzer¹⁵⁸,
 B. Spurlock⁸, R.D. St. Denis⁵³, J. Stahlman¹²⁰, R. Stamen^{58a}, E. Stanecka³⁹, R.W. Stanek⁶,
 C. Stanescu^{134a}, M. Stanescu-Bellu⁴², M.M. Stanitzki⁴², S. Stapnes¹¹⁷, E.A. Starchenko¹²⁸, J. Stark⁵⁵,
 P. Staroba¹²⁵, P. Starovoitov⁴², R. Staszewski³⁹, A. Staude⁹⁸, P. Stavina^{144a,*}, G. Steele⁵³, P. Steinbach⁴⁴,
 P. Steinberg²⁵, I. Stekl¹²⁷, B. Stelzer¹⁴², H.J. Stelzer⁸⁸, O. Stelzer-Chilton^{159a}, H. Stenzel⁵², S. Stern⁹⁹,
 G.A. Stewart³⁰, J.A. Stillings²¹, M.C. Stockton⁸⁵, K. Stoerig⁴⁸, G. Stoicea^{26a}, S. Stonjek⁹⁹, P. Strachota¹²⁶,
 A.R. Stradling⁸, A. Straessner⁴⁴, J. Strandberg¹⁴⁷, S. Strandberg^{146a,146b}, A. Strandlie¹¹⁷, M. Strang¹⁰⁹,
 E. Strauss¹⁴³, M. Strauss¹¹¹, P. Strizenec^{144b}, R. Ströhmer¹⁷⁴, D.M. Strom¹¹⁴, J.A. Strong^{76,*},
 R. Stroynowski⁴⁰, J. Strube¹²⁹, B. Stugu¹⁴, I. Stumer^{25,*}, J. Stupak¹⁴⁸, P. Sturm¹⁷⁵, N.A. Styles⁴²,
 D.A. Soh^{151,w}, D. Su¹⁴³, HS. Subramania³, A. Succurro¹², Y. Sugaya¹¹⁶, C. Suhr¹⁰⁶, M. Suk¹²⁶,
 V.V. Sulin⁹⁴, S. Sultansoy^{4d}, T. Sumida⁶⁷, X. Sun⁵⁵, J.E. Sundermann⁴⁸, K. Suruliz¹³⁹, G. Susinno^{37a,37b},
 M.R. Sutton¹⁴⁹, Y. Suzuki⁶⁵, Y. Suzuki⁶⁶, M. Svatos¹²⁵, S. Swedish¹⁶⁸, I. Sykora^{144a}, T. Sykora¹²⁶,
 J. Sánchez¹⁶⁷, D. Ta¹⁰⁵, K. Tackmann⁴², A. Taffard¹⁶³, R. Tafirout^{159a}, N. Taiblum¹⁵³, Y. Takahashi¹⁰¹,
 H. Takai²⁵, R. Takashima⁶⁸, H. Takeda⁶⁶, T. Takeshita¹⁴⁰, Y. Takubo⁶⁵, M. Talby⁸³, A. Talyshov^{107,f},
 M.C. Tamsett²⁵, K.G. Tan⁸⁶, J. Tanaka¹⁵⁵, R. Tanaka¹¹⁵, S. Tanaka¹³¹, S. Tanaka⁶⁵, A.J. Tanasijczuk¹⁴²,
 K. Tani⁶⁶, N. Tannoury⁸³, S. Tapprogge⁸¹, D. Tardif¹⁵⁸, S. Tarem¹⁵², F. Tarrade²⁹, G.F. Tartarelli^{89a},
 P. Tas¹²⁶, M. Tasevsky¹²⁵, E. Tassi^{37a,37b}, M. Tatarkhanov¹⁵, Y. Tayalati^{135d}, C. Taylor⁷⁷, F.E. Taylor⁹²,
 G.N. Taylor⁸⁶, W. Taylor^{159b}, M. Teinturier¹¹⁵, F.A. Teischinger³⁰, M. Teixeira Dias Castanheira⁷⁵,
 P. Teixeira-Dias⁷⁶, K.K. Temming⁴⁸, H. Ten Kate³⁰, P.K. Teng¹⁵¹, S. Terada⁶⁵, K. Terashi¹⁵⁵, J. Terron⁸⁰,
 M. Testa⁴⁷, R.J. Teuscher^{158,k}, J. Therhaag²¹, T. Theveneaux-Pelzer⁷⁸, S. Thoma⁴⁸, J.P. Thomas¹⁸,
 E.N. Thompson³⁵, P.D. Thompson¹⁸, P.D. Thompson¹⁵⁸, A.S. Thompson⁵³, L.A. Thomsen³⁶,
 E. Thomson¹²⁰, M. Thomson²⁸, W.M. Thong⁸⁶, R.P. Thun⁸⁷, F. Tian³⁵, M.J. Tibbetts¹⁵, T. Tic¹²⁵,
 V.O. Tikhomirov⁹⁴, Y.A. Tikhonov^{107,f}, S. Timoshenko⁹⁶, P. Tipton¹⁷⁶, S. Tisserant⁸³, T. Todorov⁵,
 S. Todorova-Nova¹⁶¹, B. Toggerson¹⁶³, J. Tojo⁶⁹, S. Tokár^{144a}, K. Tokushuku⁶⁵, K. Tollefson⁸⁸,
 M. Tomoto¹⁰¹, L. Tompkins³¹, K. Toms¹⁰³, A. Tonoyan¹⁴, C. Topfel¹⁷, N.D. Topilin⁶⁴, I. Torchiani³⁰,
 E. Torrence¹¹⁴, H. Torres⁷⁸, E. Torró Pastor¹⁶⁷, J. Toth^{83,ad}, F. Touchard⁸³, D.R. Tovey¹³⁹, T. Trefzger¹⁷⁴,
 L. Tremblet³⁰, A. Tricoli³⁰, I.M. Trigger^{159a}, S. Trincaz-Duvold⁷⁸, M.F. Tripiana⁷⁰, N. Triplett²⁵,
 W. Trischuk¹⁵⁸, B. Trocmé⁵⁵, C. Troncon^{89a}, M. Trottier-McDonald¹⁴², M. Trzebinski³⁹, A. Trzupek³⁹,
 C. Tsarouchas³⁰, J.-C.-L. Tseng¹¹⁸, M. Tsiakiris¹⁰⁵, P.V. Tsiareshka⁹⁰, D. Tsionou^{5,ai}, G. Tsipolitis¹⁰,
 S. Tsiskaridze¹², V. Tsiskaridze⁴⁸, E.G. Tskhadadze^{51a}, I.I. Tsukerman⁹⁵, V. Tsulaia¹⁵, J.-W. Tsung²¹,
 S. Tsuno⁶⁵, D. Tsybychev¹⁴⁸, A. Tua¹³⁹, A. Tudorache^{26a}, V. Tudorache^{26a}, J.M. Tuggle³¹, M. Turala³⁹,
 D. Turecek¹²⁷, I. Turk Cakir^{4e}, E. Turlay¹⁰⁵, R. Turra^{89a,89b}, P.M. Tuts³⁵, A. Tykhonov⁷⁴,
 M. Tylmad^{146a,146b}, M. Tyndel¹²⁹, G. Tzanakos⁹, K. Uchida²¹, I. Ueda¹⁵⁵, R. Ueno²⁹, M. Ugland¹⁴,
 M. Uhlenbrock²¹, M. Uhrmacher⁵⁴, F. Ukegawa¹⁶⁰, G. Unal³⁰, A. Undrus²⁵, G. Unel¹⁶³, Y. Unno⁶⁵,
 D. Urbaniec³⁵, G. Usai⁸, M. Uslenghi^{119a,119b}, L. Vacavant⁸³, V. Vacek¹²⁷, B. Vachon⁸⁵, S. Vahsen¹⁵,
 J. Valenta¹²⁵, S. Valentini^{20a,20b}, A. Valero¹⁶⁷, S. Valkar¹²⁶, E. Valladolid Gallego¹⁶⁷, S. Vallecorsa¹⁵²,
 J.A. Valls Ferrer¹⁶⁷, R. Van Berg¹²⁰, P.C. Van Der Deijl¹⁰⁵, R. van der Geer¹⁰⁵, H. van der Graaf¹⁰⁵,
 R. Van Der Leeuw¹⁰⁵, E. van der Poel¹⁰⁵, D. van der Ster³⁰, N. van Eldik³⁰, P. van Gemmeren⁶,
 I. van Vulpen¹⁰⁵, M. Vanadia⁹⁹, W. Vandelli³⁰, A. Vaniachine⁶, P. Vankov⁴², F. Vannucci⁷⁸, R. Vari^{132a},
 T. Varol⁸⁴, D. Varouchas¹⁵, A. Vartapetian⁸, K.E. Varvell¹⁵⁰, V.I. Vassilakopoulos⁵⁶, F. Vazeille³⁴,
 T. Vazquez Schroeder⁵⁴, G. Vegni^{89a,89b}, J.J. Veillet¹¹⁵, F. Veloso^{124a}, R. Veness³⁰, S. Veneziano^{132a},
 A. Ventura^{72a,72b}, D. Ventura⁸⁴, M. Venturi⁴⁸, N. Venturi¹⁵⁸, V. Vercesi^{119a}, M. Verducci¹³⁸,
 W. Verkerke¹⁰⁵, J.C. Vermeulen¹⁰⁵, A. Vest⁴⁴, M.C. Vetterli^{142,d}, I. Vichou¹⁶⁵, T. Vickey^{145b,aj},
 O.E. Vickey Boeriu^{145b}, G.H.A. Viehhauser¹¹⁸, S. Viel¹⁶⁸, M. Villa^{20a,20b}, M. Villaplana Perez¹⁶⁷,
 E. Vilucchi⁴⁷, M.G. Vincter²⁹, E. Vinek³⁰, V.B. Vinogradov⁶⁴, M. Virchaux^{136,*}, J. Virzi¹⁵, O. Vitells¹⁷²,
 M. Viti⁴², I. Vivarelli⁴⁸, F. Vives Vaque³, S. Vlachos¹⁰, D. Vladoiu⁹⁸, M. Vlasak¹²⁷, A. Vogel²¹,
 P. Vokac¹²⁷, G. Volpi⁴⁷, M. Volpi⁸⁶, G. Volpini^{89a}, H. von der Schmitt⁹⁹, H. von Radziewski⁴⁸,

E. von Toerne²¹, V. Vorobel¹²⁶, V. Vorwerk¹², M. Vos¹⁶⁷, R. Voss³⁰, T.T. Voss¹⁷⁵, J.H. Vossebeld⁷³, N. Vranjes¹³⁶, M. Vranjes Milosavljevic¹⁰⁵, V. Vrba¹²⁵, M. Vreeswijk¹⁰⁵, T. Vu Anh⁴⁸, R. Vuillermet³⁰, I. Vukotic³¹, W. Wagner¹⁷⁵, P. Wagner¹²⁰, H. Wahlen¹⁷⁵, S. Wahrmund⁴⁴, J. Wakabayashi¹⁰¹, S. Walch⁸⁷, J. Walder⁷¹, R. Walker⁹⁸, W. Walkowiak¹⁴¹, R. Wall¹⁷⁶, P. Waller⁷³, B. Walsh¹⁷⁶, C. Wang⁴⁵, H. Wang¹⁷³, H. Wang^{33b,ak}, J. Wang¹⁵¹, J. Wang⁵⁵, R. Wang¹⁰³, S.M. Wang¹⁵¹, T. Wang²¹, A. Warburton⁸⁵, C.P. Ward²⁸, M. Warsinsky⁴⁸, A. Washbrook⁴⁶, C. Wasicki⁴², I. Watanabe⁶⁶, P.M. Watkins¹⁸, A.T. Watson¹⁸, I.J. Watson¹⁵⁰, M.F. Watson¹⁸, G. Watts¹³⁸, S. Watts⁸², A.T. Waugh¹⁵⁰, B.M. Waugh⁷⁷, M.S. Weber¹⁷, P. Weber⁵⁴, A.R. Weidberg¹¹⁸, P. Weigell⁹⁹, J. Weingarten⁵⁴, C. Weiser⁴⁸, P.S. Wells³⁰, T. Wenaus²⁵, D. Wendland¹⁶, Z. Weng^{151,w}, T. Wengler³⁰, S. Wenig³⁰, N. Wermes²¹, M. Werner⁴⁸, P. Werner³⁰, M. Werth¹⁶³, M. Wessels^{58a}, J. Wetter¹⁶¹, C. Weydert⁵⁵, K. Whalen²⁹, S.J. Wheeler-Ellis¹⁶³, A. White⁸, M.J. White⁸⁶, S. White^{122a,122b}, S.R. Whitehead¹¹⁸, D. Whiteson¹⁶³, D. Whittington⁶⁰, F. Wicek¹¹⁵, D. Wicke¹⁷⁵, F.J. Wickens¹²⁹, W. Wiedenmann¹⁷³, M. Wielers¹²⁹, P. Wienemann²¹, C. Wiglesworth⁷⁵, L.A.M. Wiik-Fuchs⁴⁸, P.A. Wijeratne⁷⁷, A. Wildauer⁹⁹, M.A. Wildt^{42,s}, I. Wilhelm¹²⁶, H.G. Wilkens³⁰, J.Z. Will⁹⁸, E. Williams³⁵, H.H. Williams¹²⁰, W. Willis³⁵, S. Willocq⁸⁴, J.A. Wilson¹⁸, M.G. Wilson¹⁴³, A. Wilson⁸⁷, I. Wingerter-Seez⁵, S. Winkelmann⁴⁸, F. Winklmeier³⁰, M. Wittgen¹⁴³, S.J. Wollstadt⁸¹, M.W. Wolter³⁹, H. Wolters^{124a,h}, W.C. Wong⁴¹, G. Wooden⁸⁷, B.K. Wosiek³⁹, J. Wotschack³⁰, M.J. Woudstra⁸², K.W. Wozniak³⁹, K. Wraight⁵³, M. Wright⁵³, B. Wrona⁷³, S.L. Wu¹⁷³, X. Wu⁴⁹, Y. Wu^{33b,al}, E. Wulf³⁵, B.M. Wynne⁴⁶, S. Xella³⁶, M. Xiao¹³⁶, S. Xie⁴⁸, C. Xu^{33b,z}, D. Xu¹³⁹, B. Yabsley¹⁵⁰, S. Yacoob^{145a,am}, M. Yamada⁶⁵, H. Yamaguchi¹⁵⁵, A. Yamamoto⁶⁵, K. Yamamoto⁶³, S. Yamamoto¹⁵⁵, T. Yamamura¹⁵⁵, T. Yamanaka¹⁵⁵, J. Yamaoka⁴⁵, T. Yamazaki¹⁵⁵, Y. Yamazaki⁶⁶, Z. Yan²², H. Yang⁸⁷, U.K. Yang⁸², Y. Yang⁶⁰, Z. Yang^{146a,146b}, S. Yanush⁹¹, L. Yao^{33a}, Y. Yao¹⁵, Y. Yasu⁶⁵, G.V. Ybeles Smit¹³⁰, J. Ye⁴⁰, S. Ye²⁵, M. Yilmaz^{4c}, R. Yoosoofmiya¹²³, K. Yorita¹⁷¹, R. Yoshida⁶, C. Young¹⁴³, C.J. Young¹¹⁸, S. Youssef²², D. Yu²⁵, J. Yu⁸, J. Yu¹¹², L. Yuan⁶⁶, A. Yurkewicz¹⁰⁶, B. Zabinski³⁹, R. Zaidan⁶², A.M. Zaitsev¹²⁸, Z. Zajacova³⁰, L. Zanello^{132a,132b}, D. Zanzi⁹⁹, A. Zaytsev²⁵, C. Zeitnitz¹⁷⁵, M. Zeman¹²⁵, A. Zemla³⁹, C. Zendler²¹, O. Zenin¹²⁸, T. Ženiš^{144a}, Z. Zinonos^{122a,122b}, S. Zenz¹⁵, D. Zerwas¹¹⁵, G. Zevi della Porta⁵⁷, Z. Zhan^{33d}, D. Zhang^{33b,ak}, H. Zhang⁸⁸, J. Zhang⁶, X. Zhang^{33d}, Z. Zhang¹¹⁵, L. Zhao¹⁰⁸, T. Zhao¹³⁸, Z. Zhao^{33b}, A. Zhemchugov⁶⁴, J. Zhong¹¹⁸, B. Zhou⁸⁷, N. Zhou¹⁶³, Y. Zhou¹⁵¹, C.G. Zhu^{33d}, H. Zhu⁴², J. Zhu⁸⁷, Y. Zhu^{33b}, X. Zhuang⁹⁸, V. Zhuravlov⁹⁹, D. Zieminska⁶⁰, N.I. Zimin⁶⁴, R. Zimmermann²¹, S. Zimmermann²¹, S. Zimmermann⁴⁸, M. Ziolkowski¹⁴¹, R. Zitoun⁵, L. Živković³⁵, V.V. Zmouchko^{128,*}, G. Zobernig¹⁷³, A. Zoccoli^{20a,20b}, M. zur Nedden¹⁶, V. Zutshi¹⁰⁶, L. Zwalski³⁰

¹ School of Chemistry and Physics, University of Adelaide, Adelaide, Australia² Physics Department, SUNY Albany, Albany, NY, United States³ Department of Physics, University of Alberta, Edmonton, AB, Canada⁴ (a) Department of Physics, Ankara University, Ankara; (b) Department of Physics, Dumlupinar University, Kutahya; (c) Department of Physics, Gazi University, Ankara; (d) Division of Physics, TOBB University of Economics and Technology, Ankara; (e) Turkish Atomic Energy Authority, Ankara, Turkey⁵ LAPP, CNRS/IN2P3 and Université de Savoie, Annecy-le-Vieux, France⁶ High Energy Physics Division, Argonne National Laboratory, Argonne, IL, United States⁷ Department of Physics, University of Arizona, Tucson, AZ, United States⁸ Department of Physics, The University of Texas at Arlington, Arlington, TX, United States⁹ Physics Department, University of Athens, Athens, Greece¹⁰ Physics Department, National Technical University of Athens, Zografou, Greece¹¹ Institute of Physics, Azerbaijan Academy of Sciences, Baku, Azerbaijan¹² Institut de Física d'Altes Energies and Departament de Física de la Universitat Autònoma de Barcelona and ICREA, Barcelona, Spain¹³ (a) Institute of Physics, University of Belgrade, Belgrade; (b) Vinca Institute of Nuclear Sciences, University of Belgrade, Belgrade, Serbia¹⁴ Department for Physics and Technology, University of Bergen, Bergen, Norway¹⁵ Physics Division, Lawrence Berkeley National Laboratory and University of California, Berkeley, CA, United States¹⁶ Department of Physics, Humboldt University, Berlin, Germany¹⁷ Albert Einstein Center for Fundamental Physics and Laboratory for High Energy Physics, University of Bern, Bern, Switzerland¹⁸ School of Physics and Astronomy, University of Birmingham, Birmingham, United Kingdom¹⁹ (a) Department of Physics, Bogazici University, Istanbul; (b) Division of Physics, Dogus University, Istanbul; (c) Department of Physics Engineering, Gaziantep University, Gaziantep;²⁰ (a) Department of Physics, Istanbul Technical University, Istanbul, Turkey²⁰ (a) INFN Sezione di Bologna; (b) Dipartimento di Fisica, Università di Bologna, Bologna, Italy²¹ Physikalisches Institut, University of Bonn, Bonn, Germany²² Department of Physics, Boston University, Boston, MA, United States²³ Department of Physics, Brandeis University, Waltham, MA, United States²⁴ (a) Universidade Federal do Rio De Janeiro COPPE/EE/IF, Rio de Janeiro; (b) Federal University of Juiz de Fora (UFJF), Juiz de Fora; (c) Federal University of São João del Rei (UFSJ), São João del Rei; (d) Instituto de Física, Universidade de São Paulo, São Paulo, Brazil²⁵ Physics Department, Brookhaven National Laboratory, Upton, NY, United States²⁶ (a) National Institute of Physics and Nuclear Engineering, Bucharest; (b) University Politehnica Bucharest, Bucharest; (c) West University in Timisoara, Timisoara, Romania²⁷ Departamento de Física, Universidad de Buenos Aires, Buenos Aires, Argentina²⁸ Cavendish Laboratory, University of Cambridge, Cambridge, United Kingdom

- ²⁹ Department of Physics, Carleton University, Ottawa, ON, Canada
³⁰ CERN, Geneva, Switzerland
³¹ Enrico Fermi Institute, University of Chicago, Chicago, IL, United States
³² ^(a) Departamento de Física, Pontificia Universidad Católica de Chile, Santiago; ^(b) Departamento de Física, Universidad Técnica Federico Santa María, Valparaíso, Chile
³³ ^(a) Institute of High Energy Physics, Chinese Academy of Sciences, Beijing; ^(b) Department of Modern Physics, University of Science and Technology of China, Anhui;
^(c) Department of Physics, Nanjing University, Jiangsu; ^(d) School of Physics, Shandong University, Shandong, China
³⁴ Laboratoire de Physique Corpusculaire, Clermont Université and Université Blaise Pascal and CNRS/IN2P3, Clermont-Ferrand, France
³⁵ Nevis Laboratory, Columbia University, Irvington, NY, United States
³⁶ Niels Bohr Institute, University of Copenhagen, Copenhagen, Denmark
³⁷ ^(a) INFN Gruppo Collegato di Cosenza; ^(b) Dipartimento di Fisica, Università della Calabria, Arcavate di Rende, Italy
³⁸ AGH University of Science and Technology, Faculty of Physics and Applied Computer Science, Krakow, Poland
³⁹ The Henryk Niewodniczanski Institute of Nuclear Physics, Polish Academy of Sciences, Krakow, Poland
⁴⁰ Physics Department, Southern Methodist University, Dallas, TX, United States
⁴¹ Physics Department, University of Texas at Dallas, Richardson, TX, United States
⁴² DESY, Hamburg and Zeuthen, Germany
⁴³ Institut für Experimentelle Physik IV, Technische Universität Dortmund, Dortmund, Germany
⁴⁴ Institut für Kern- und Teilchenphysik, Technical University Dresden, Dresden, Germany
⁴⁵ Department of Physics, Duke University, Durham, NC, United States
⁴⁶ SUPA – School of Physics and Astronomy, University of Edinburgh, Edinburgh, United Kingdom
⁴⁷ INFN Laboratori Nazionali di Frascati, Frascati, Italy
⁴⁸ Fakultät für Mathematik und Physik, Albert-Ludwigs-Universität, Freiburg, Germany
⁴⁹ Section de Physique, Université de Genève, Geneva, Switzerland
⁵⁰ ^(a) INFN Sezione di Genova; ^(b) Dipartimento di Fisica, Università di Genova, Genova, Italy
⁵¹ ^(a) E. Andronikashvili Institute of Physics, Tbilisi State University, Tbilisi; ^(b) High Energy Physics Institute, Tbilisi State University, Tbilisi, Georgia
⁵² II. Physikalisches Institut, Justus-Liebig-Universität Giessen, Giessen, Germany
⁵³ SUPA – School of Physics and Astronomy, University of Glasgow, Glasgow, United Kingdom
⁵⁴ II. Physikalisches Institut, Georg-August-Universität, Göttingen, Germany
⁵⁵ Laboratoire de Physique Subatomique et de Cosmologie, Université Joseph Fourier and CNRS/IN2P3 and Institut National Polytechnique de Grenoble, Grenoble, France
⁵⁶ Department of Physics, Hampton University, Hampton, VA, United States
⁵⁷ Laboratory for Particle Physics and Cosmology, Harvard University, Cambridge, MA, United States
⁵⁸ ^(a) Kirchhoff-Institut für Physik, Ruprecht-Karls-Universität Heidelberg, Heidelberg; ^(b) Physikalisches Institut, Ruprecht-Karls-Universität Heidelberg, Heidelberg;
^(c) ZITI Institut für technische Informatik, Ruprecht-Karls-Universität Heidelberg, Mannheim, Germany
⁵⁹ Faculty of Applied Information Science, Hiroshima Institute of Technology, Hiroshima, Japan
⁶⁰ Department of Physics, Indiana University, Bloomington, IN, United States
⁶¹ Institut für Astro- und Teilchenphysik, Leopold-Franzens-Universität, Innsbruck, Austria
⁶² University of Iowa, Iowa City, IA, United States
⁶³ Department of Physics and Astronomy, Iowa State University, Ames, IA, United States
⁶⁴ Joint Institute for Nuclear Research, JINR Dubna, Dubna, Russia
⁶⁵ KEK, High Energy Accelerator Research Organization, Tsukuba, Japan
⁶⁶ Graduate School of Science, Kobe University, Kobe, Japan
⁶⁷ Faculty of Science, Kyoto University, Kyoto, Japan
⁶⁸ Kyoto University of Education, Kyoto, Japan
⁶⁹ Department of Physics, Kyushu University, Fukuoka, Japan
⁷⁰ Instituto de Física La Plata, Universidad Nacional de La Plata and CONICET, La Plata, Argentina
⁷¹ Physics Department, Lancaster University, Lancaster, United Kingdom
⁷² ^(a) INFN Sezione di Lecce; ^(b) Dipartimento di Matematica e Fisica, Università del Salento, Lecce, Italy
⁷³ Oliver Lodge Laboratory, University of Liverpool, Liverpool, United Kingdom
⁷⁴ Department of Physics, Jožef Stefan Institute and University of Ljubljana, Ljubljana, Slovenia
⁷⁵ School of Physics and Astronomy, Queen Mary University of London, London, United Kingdom
⁷⁶ Department of Physics, Royal Holloway University of London, Surrey, United Kingdom
⁷⁷ Department of Physics and Astronomy, University College London, London, United Kingdom
⁷⁸ Laboratoire de Physique Nucléaire et de Hautes Energies, UPMC and Université Paris-Diderot and CNRS/IN2P3, Paris, France
⁷⁹ Fysiska institutionen, Lunds universitet, Lund, Sweden
⁸⁰ Departamento de Física Teórica C-15, Universidad Autónoma de Madrid, Madrid, Spain
⁸¹ Institut für Physik, Universität Mainz, Mainz, Germany
⁸² School of Physics and Astronomy, University of Manchester, Manchester, United Kingdom
⁸³ CPPM, Aix-Marseille Université and CNRS/IN2P3, Marseille, France
⁸⁴ Department of Physics, University of Massachusetts, Amherst, MA, United States
⁸⁵ Department of Physics, McGill University, Montreal, QC, Canada
⁸⁶ School of Physics, University of Melbourne, Victoria, Australia
⁸⁷ Department of Physics, The University of Michigan, Ann Arbor, MI, United States
⁸⁸ Department of Physics and Astronomy, Michigan State University, East Lansing, MI, United States
⁸⁹ ^(a) INFN Sezione di Milano; ^(b) Dipartimento di Fisica, Università di Milano, Milano, Italy
⁹⁰ B.I. Stepanov Institute of Physics, National Academy of Sciences of Belarus, Minsk, Belarus
⁹¹ National Scientific and Educational Centre for Particle and High Energy Physics, Minsk, Belarus
⁹² Department of Physics, Massachusetts Institute of Technology, Cambridge, MA, United States
⁹³ Group of Particle Physics, University of Montreal, Montreal, QC, Canada
⁹⁴ P.N. Lebedev Institute of Physics, Academy of Sciences, Moscow, Russia
⁹⁵ Institute for Theoretical and Experimental Physics (ITEP), Moscow, Russia
⁹⁶ Moscow Engineering and Physics Institute (MEPhI), Moscow, Russia
⁹⁷ Skobeltsyn Institute of Nuclear Physics, Lomonosov Moscow State University, Moscow, Russia
⁹⁸ Fakultät für Physik, Ludwig-Maximilians-Universität München, München, Germany
⁹⁹ Max-Planck-Institut für Physik (Werner-Heisenberg-Institut), München, Germany
¹⁰⁰ Nagasaki Institute of Applied Science, Nagasaki, Japan
¹⁰¹ Graduate School of Science and Kobayashi-Maskawa Institute, Nagoya University, Nagoya, Japan
¹⁰² ^(a) INFN Sezione di Napoli; ^(b) Dipartimento di Scienze Fisiche, Università di Napoli, Napoli, Italy
¹⁰³ Department of Physics and Astronomy, University of New Mexico, Albuquerque, NM, United States
¹⁰⁴ Institute for Mathematics, Astrophysics and Particle Physics, Radboud University Nijmegen/Nikhef, Nijmegen, Netherlands
¹⁰⁵ Nikhef National Institute for Subatomic Physics and University of Amsterdam, Amsterdam, Netherlands

- ¹⁰⁶ Department of Physics, Northern Illinois University, DeKalb, IL, United States
¹⁰⁷ Budker Institute of Nuclear Physics, SB RAS, Novosibirsk, Russia
¹⁰⁸ Department of Physics, New York University, New York, NY, United States
¹⁰⁹ Ohio State University, Columbus, OH, United States
¹¹⁰ Faculty of Science, Okayama University, Okayama, Japan
¹¹¹ Homer L. Dodge Department of Physics and Astronomy, University of Oklahoma, Norman, OK, United States
¹¹² Department of Physics, Oklahoma State University, Stillwater, OK, United States
¹¹³ Palacký University, RCPMT, Olomouc, Czech Republic
¹¹⁴ Center for High Energy Physics, University of Oregon, Eugene, OR, United States
¹¹⁵ LAI, Université Paris-Sud and CNRS/IN2P3, Orsay, France
¹¹⁶ Graduate School of Science, Osaka University, Osaka, Japan
¹¹⁷ Department of Physics, University of Oslo, Oslo, Norway
¹¹⁸ Department of Physics, Oxford University, Oxford, United Kingdom
¹¹⁹ ^(a)INFN Sezione di Pavia; ^(b) Dipartimento di Fisica, Università di Pavia, Pavia, Italy
¹²⁰ Department of Physics, University of Pennsylvania, Philadelphia, PA, United States
¹²¹ Petersburg Nuclear Physics Institute, Gatchina, Russia
¹²² ^(a)INFN Sezione di Pisa; ^(b) Dipartimento di Fisica E. Fermi, Università di Pisa, Pisa, Italy
¹²³ Department of Physics and Astronomy, University of Pittsburgh, Pittsburgh, PA, United States
¹²⁴ ^(a)Laboratorio de Instrumentacao e Fisica Experimental de Particulas - LIP, Lisboa, Portugal; ^(b) Departamento de Fisica Teorica y del Cosmos and CAFPE, Universidad de Granada, Granada, Spain
¹²⁵ Institute of Physics, Academy of Sciences of the Czech Republic, Praha, Czech Republic
¹²⁶ Faculty of Mathematics and Physics, Charles University in Prague, Praha, Czech Republic
¹²⁷ Czech Technical University in Prague, Praha, Czech Republic
¹²⁸ State Research Center Institute for High Energy Physics, Protvino, Russia
¹²⁹ Particle Physics Department, Rutherford Appleton Laboratory, Didcot, United Kingdom
¹³⁰ Physics Department, University of Regina, Regina, SK, Canada
¹³¹ Ritsumeikan University, Kusatsu, Shiga, Japan
¹³² ^(a)INFN Sezione di Roma I; ^(b) Dipartimento di Fisica, Università La Sapienza, Roma, Italy
¹³³ ^(a)INFN Sezione di Roma Tor Vergata; ^(b) Dipartimento di Fisica, Università di Roma Tor Vergata, Roma, Italy
¹³⁴ ^(a)INFN Sezione di Roma Tre; ^(b) Dipartimento di Fisica, Università Roma Tre, Roma, Italy
¹³⁵ ^(a)Faculté des Sciences Ain Chock, Réseau Universitaire de Physique des Hautes Energies – Université Hassan II, Casablanca; ^(b) Centre National de l'Energie des Sciences Techniques Nucléaires, Rabat; ^(c) Faculté des Sciences Semlalia, Université Cadi Ayyad, LPHEA, Marrakech; ^(d) Faculté des Sciences, Université Mohamed Premier and LPTPM, Oujda; ^(e) Faculté des Sciences, Université Mohammed V, Agdal, Rabat, Morocco
¹³⁶ DSM/IRFU (Institut de Recherches sur les Lois Fondamentales de l'Univers), CEA Saclay (Commissariat à l'Energie Atomique), Gif-sur-Yvette, France
¹³⁷ Santa Cruz Institute for Particle Physics, University of California Santa Cruz, Santa Cruz, CA, United States
¹³⁸ Department of Physics, University of Washington, Seattle, WA, United States
¹³⁹ Department of Physics and Astronomy, University of Sheffield, Sheffield, United Kingdom
¹⁴⁰ Department of Physics, Shinshu University, Nagano, Japan
¹⁴¹ Fachbereich Physik, Universität Siegen, Siegen, Germany
¹⁴² Department of Physics, Simon Fraser University, Burnaby, BC, Canada
¹⁴³ SLAC National Accelerator Laboratory, Stanford, CA, United States
¹⁴⁴ ^(a)Faculty of Mathematics, Physics & Informatics, Comenius University, Bratislava; ^(b) Department of Subnuclear Physics, Institute of Experimental Physics of the Slovak Academy of Sciences, Košice, Slovak Republic
¹⁴⁵ ^(a)Department of Physics, University of Johannesburg, Johannesburg; ^(b) School of Physics, University of the Witwatersrand, Johannesburg, South Africa
¹⁴⁶ ^(a)Department of Physics, Stockholm University; ^(b) The Oskar Klein Centre, Stockholm, Sweden
¹⁴⁷ Physics Department, Royal Institute of Technology, Stockholm, Sweden
¹⁴⁸ Departments of Physics & Astronomy and Chemistry, Stony Brook University, Stony Brook, NY, United States
¹⁴⁹ Department of Physics and Astronomy, University of Sussex, Brighton, United Kingdom
¹⁵⁰ School of Physics, University of Sydney, Sydney, Australia
¹⁵¹ Institute of Physics, Academia Sinica, Taipei, Taiwan
¹⁵² Department of Physics, Technion: Israel Institute of Technology, Haifa, Israel
¹⁵³ Raymond and Beverly Sackler School of Physics and Astronomy, Tel Aviv University, Tel Aviv, Israel
¹⁵⁴ Department of Physics, Aristotle University of Thessaloniki, Thessaloniki, Greece
¹⁵⁵ International Center for Elementary Particle Physics and Department of Physics, The University of Tokyo, Tokyo, Japan
¹⁵⁶ Graduate School of Science and Technology, Tokyo Metropolitan University, Tokyo, Japan
¹⁵⁷ Department of Physics, Tokyo Institute of Technology, Tokyo, Japan
¹⁵⁸ Department of Physics, University of Toronto, Toronto, ON, Canada
¹⁵⁹ ^(a)TRIUMF, Vancouver, BC; ^(b) Department of Physics and Astronomy, York University, Toronto, ON, Canada
¹⁶⁰ Faculty of Pure and Applied Sciences, University of Tsukuba, Tsukuba, Japan
¹⁶¹ Department of Physics and Astronomy, Tufts University, Medford, MA, United States
¹⁶² Centro de Investigaciones, Universidad Antonio Narino, Bogota, Colombia
¹⁶³ Department of Physics and Astronomy, University of California Irvine, Irvine, CA, United States
¹⁶⁴ ^(a)INFN Gruppo Collegato di Udine; ^(b) ICTP, Trieste; ^(c) Dipartimento di Chimica, Fisica e Ambiente, Università di Udine, Udine, Italy
¹⁶⁵ Department of Physics, University of Illinois, Urbana, IL, United States
¹⁶⁶ Department of Physics and Astronomy, University of Uppsala, Uppsala, Sweden
¹⁶⁷ Instituto de Física Corpuscular (IFIC) and Departamento de Física Atómica, Molecular y Nuclear and Departamento de Ingeniería Electrónica and Instituto de Microelectrónica de Barcelona (IMB-CNM), University of Valencia and CSIC, Valencia, Spain
¹⁶⁸ Department of Physics, University of British Columbia, Vancouver, BC, Canada
¹⁶⁹ Department of Physics and Astronomy, University of Victoria, Victoria, BC, Canada
¹⁷⁰ Department of Physics, University of Warwick, Coventry, United Kingdom
¹⁷¹ Waseda University, Tokyo, Japan
¹⁷² Department of Particle Physics, The Weizmann Institute of Science, Rehovot, Israel
¹⁷³ Department of Physics, University of Wisconsin, Madison, WI, United States
¹⁷⁴ Fakultät für Physik und Astronomie, Julius-Maximilians-Universität, Würzburg, Germany
¹⁷⁵ Fachbereich C Physik, Bergische Universität Wuppertal, Wuppertal, Germany
¹⁷⁶ Department of Physics, Yale University, New Haven, CT, United States
¹⁷⁷ Yerevan Physics Institute, Yerevan, Armenia
¹⁷⁸ Centre de Calcul de l'Institut National de Physique Nucléaire et de Physique des Particules (IN2P3), Villeurbanne, France

- ^a Also at Laboratorio de Instrumentacao e Fisica Experimental de Particulas – LIP, Lisboa, Portugal.
- ^b Also at Faculdade de Ciencias and CFNUL, Universidade de Lisboa, Lisboa, Portugal.
- ^c Also at Particle Physics Department, Rutherford Appleton Laboratory, Didcot, United Kingdom.
- ^d Also at TRIUMF, Vancouver, BC, Canada.
- ^e Also at Department of Physics, California State University, Fresno, CA, United States.
- ^f Also at Novosibirsk State University, Novosibirsk, Russia.
- ^g Also at Fermilab, Batavia, IL, United States.
- ^h Also at Department of Physics, University of Coimbra, Coimbra, Portugal.
- ⁱ Also at Department of Physics, UASLP, San Luis Potosi, Mexico.
- ^j Also at Università di Napoli Parthenope, Napoli, Italy.
- ^k Also at Institute of Particle Physics (IPP), Canada.
- ^l Also at Department of Physics, Middle East Technical University, Ankara, Turkey.
- ^m Also at Louisiana Tech University, Ruston, LA, United States.
- ⁿ Also at Dep. Fisica and CEFITEC of Faculdade de Ciencias e Tecnologia, Universidade Nova de Lisboa, Caparica, Portugal.
- ^o Also at Department of Physics and Astronomy, University College London, London, United Kingdom.
- ^p Also at Group of Particle Physics, University of Montreal, Montreal, QC, Canada.
- ^q Also at Department of Physics, University of Cape Town, Cape Town, South Africa.
- ^r Also at Institute of Physics, Azerbaijan Academy of Sciences, Baku, Azerbaijan.
- ^s Also at Institut für Experimentalphysik, Universität Hamburg, Hamburg, Germany.
- ^t Also at Manhattan College, New York, NY, United States.
- ^u Also at School of Physics, Shandong University, Shandong, China.
- ^v Also at CPPM, Aix-Marseille Université and CNRS/IN2P3, Marseille, France.
- ^w Also at School of Physics and Engineering, Sun Yat-sen University, Guangzhou, China.
- ^x Also at Academia Sinica Grid Computing, Institute of Physics, Academia Sinica, Taipei, Taiwan.
- ^y Also at Dipartimento di Fisica, Università La Sapienza, Roma, Italy.
- ^z Also at DSM/IRFU (Institut de Recherches sur les Lois Fondamentales de l'Univers), CEA Saclay (Commissariat à l'Energie Atomique), Gif-sur-Yvette, France.
- ^{aa} Also at Section de Physique, Université de Genève, Geneva, Switzerland.
- ^{ab} Also at Departamento de Fisica, Universidade de Minho, Braga, Portugal.
- ^{ac} Also at Department of Physics and Astronomy, University of South Carolina, Columbia, SC, United States.
- ^{ad} Also at Institute for Particle and Nuclear Physics, Wigner Research Centre for Physics, Budapest, Hungary.
- ^{ae} Also at California Institute of Technology, Pasadena, CA, United States.
- ^{af} Also at Institute of Physics, Jagiellonian University, Krakow, Poland.
- ^{ag} Also at LAL, Université Paris-Sud and CNRS/IN2P3, Orsay, France.
- ^{ah} Also at Nevis Laboratory, Columbia University, Irvington, NY, United States.
- ^{ai} Also at Department of Physics and Astronomy, University of Sheffield, Sheffield, United Kingdom.
- ^{aj} Also at Department of Physics, Oxford University, Oxford, United Kingdom.
- ^{ak} Also at Institute of Physics, Academia Sinica, Taipei, Taiwan.
- ^{al} Also at Department of Physics, The University of Michigan, Ann Arbor, MI, United States.
- ^{am} Also at Discipline of Physics, University of KwaZulu-Natal, Durban, South Africa.
- * Deceased.