



# Search for the $X_b$ and other hidden-beauty states in the $\pi^+\pi^-\gamma(1S)$ channel at ATLAS



ATLAS Collaboration\*

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## ABSTRACT

This Letter presents a search for a hidden-beauty counterpart of the  $X(3872)$  in the mass ranges of 10.05–10.31 GeV and 10.40–11.00 GeV, in the channel  $X_b \rightarrow \pi^+\pi^-\gamma(1S)(\rightarrow \mu^+\mu^-)$ , using  $16.2 \text{ fb}^{-1}$  of  $\sqrt{s} = 8 \text{ TeV}$   $pp$  collision data collected by the ATLAS detector at the LHC. No evidence for new narrow states is found, and upper limits are set on the product of the  $X_b$  cross section and branching fraction, relative to those of the  $\gamma(2S)$ , at the 95% confidence level using the  $\text{CL}_s$  approach. These limits range from 0.8% to 4.0%, depending on mass. For masses above 10.1 GeV, the expected upper limits from this analysis are the most restrictive to date. Searches for production of the  $\gamma(1^3D_J)$ ,  $\gamma(10860)$ , and  $\gamma(11020)$  states also reveal no significant signals.

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## 1. Introduction

The  $X(3872)$  is the first and the best-studied of the new hidden-charm states seen in the last decade. Observed by Belle in decays  $B^\pm \rightarrow K^\pm X(\rightarrow \pi^+\pi^- J/\psi)$  [1], it was quickly confirmed by BaBar [2], CDF [3], and DØ [4]. In particular, CDF and DØ found that the  $X(3872)$  is produced directly in  $p\bar{p}$  collisions; recently CMS has measured the product of the  $pp$  production cross section and the  $\pi^+\pi^- J/\psi$  branching fraction to be  $(6.56 \pm 0.29 \pm 0.65)\%$  of the value for the  $\psi(2S)$  [5]. The mass, narrow width,  $J^{PC} = 1^{++}$  quantum number assignment [6–10], and decay characteristics of the  $X(3872)$  make it unlikely to be a conventional quarkonium state, and there is an extensive literature discussing its structure. Weakly bound  $D^0\bar{D}^{*0}$  molecular models (for example, Refs. [11, 12]) have been popular due to the proximity of the  $X(3872)$  to the  $D^0\bar{D}^{*0}$  threshold; various  $[qc][\bar{q}\bar{c}]$  tetraquark (for example, Refs. [13,14]) and other models have also been proposed.

Heavy-quark symmetry suggests the existence of a hidden-beauty partner – a so-called  $X_b$  state – which should be produced in  $pp$  collisions [15]. The molecular model of Swanson [12,16] predicts an  $X_b$  mass of 10561 MeV, while tetraquark predictions vary: for example, Ref. [14] predicts masses of 10492, 10593, or 10682 MeV, depending on the flavour of the light quarks.

The decay  $X_b \rightarrow \pi^+\pi^-\gamma(1S)(\rightarrow \mu^+\mu^-)$ , analogous to the decay mode in which the  $X(3872)$  was discovered, provides a straightforward way to reconstruct an  $X_b$ . Any resulting measurement or upper limit on the  $X_b$  production cross section then de-

pends on the branching fraction for  $X_b \rightarrow \pi^+\pi^-\gamma(1S)$ , which is unknown.

The  $\pi^+\pi^-\gamma(1S)$  channel also provides the opportunity to measure the production of the  $\gamma(1^3D_J)$  states. These have not been observed at the Tevatron; their production cross sections in  $pp$  collisions are also unknown, but an early colour-octet calculation [17] gives values comparable to that of the  $\gamma(2S)$ . The  $\gamma(1^3D_2)$  has been observed in radiative transitions by CLEO [18] and BaBar [19].

The production of  $\gamma(10860)$  and some other hidden-beauty states may also be studied using  $\pi^+\pi^-\gamma(1S)$ . The  $\gamma(10860)$  decay to  $\pi^+\pi^-\gamma(1S)$  has a surprisingly large partial width [20–22]: current world average results are  $(0.29 \pm 0.15) \text{ MeV}$  for the  $\gamma(10860)$ , to be compared with  $(0.89 \pm 0.08) \text{ keV}$  for the  $\gamma(3S)$ , and  $(1.7 \pm 0.2) \text{ keV}$  for the  $\gamma(4S)$  [23]. Belle has also presented evidence of exotic substructure in this decay [24]. The natural widths of the  $\gamma(10860)$ ,  $\gamma(11020)$ , and other states above the open-beauty threshold are larger than the detector resolution, and must be explicitly considered in any search.

In 2013, CMS reported [25] the results of their search for  $X_b \rightarrow \pi^+\pi^-\gamma(1S)(\rightarrow \mu^+\mu^-)$ , finding no evidence for narrow states in the 10.06–10.31 GeV and 10.40–10.99 GeV mass ranges. They set upper limits on the product of cross section and branching fraction at values between 0.9% and 5.4% of the  $\gamma(2S)$  rate.

This Letter presents a search for the  $X_b$  and other hidden-beauty states at ATLAS, using a  $16.2 \text{ fb}^{-1}$   $pp$  collision data sample collected at  $\sqrt{s} = 8 \text{ TeV}$  during the 2012 run of the LHC. The analysis is performed simultaneously across eight kinematic bins of varying sensitivity; the  $\gamma(2S)$  and  $\gamma(3S) \rightarrow \pi^+\pi^-\gamma(1S)$  peaks are used to validate the measurement technique. Results are

\* E-mail address: atlas.publications@cern.ch.

presented in terms of the product of production cross section and  $\pi^+\pi^-\gamma(1S)$  branching fraction, relative to that for the  $\gamma(2S)$ .

## 2. The ATLAS detector

The ATLAS detector [26] is composed of an inner tracking system, calorimeters, and a muon spectrometer. The inner detector (ID) surrounds the  $pp$  interaction point and consists of silicon pixel and microstrip detectors, and a transition radiation tracker, all immersed in a 2 T axial magnetic field. The ID spans the pseudorapidity<sup>1</sup> range  $|\eta| < 2.5$  and is enclosed by a system of electromagnetic and hadronic calorimeters. Surrounding the calorimeters is the muon spectrometer (MS) consisting of three large air-core superconducting magnets (each with eight coils) providing a toroidal field, a system of precision tracking chambers, and fast detectors for triggering. Monitored drift tubes and cathode-strip chambers provide precision measurements in the bending plane of muons within the pseudorapidity range  $|\eta| < 2.7$ . Resistive plate and thin gap chambers are used to make fast event data-recording decisions in the ranges  $|\eta| < 1.05$  and  $1.05 < |\eta| < 2.4$  respectively, and also provide position measurements in the non-bending plane and improve pattern recognition and track reconstruction. MS momentum measurements are based on track segments formed in at least two of the three precision chamber planes.

The ATLAS detector employs a three-level trigger [27] to reduce the 20 MHz proton bunch collision rate to the few-hundred hertz transfer rate to mass storage. This analysis is based on a Level-1 muon trigger that searches for hit coincidences between muon trigger detector layers inside pre-programmed geometrical windows that bound the path of muon candidates above a given  $p_T^\mu$  threshold, and provide a rough estimate of their position, for  $|\eta^\mu| < 2.4$ . There are two subsequent software-based trigger stages, in which muon candidates incorporate, with increasing precision, information from both the MS and the ID, reaching position and momentum resolution close to that provided by the offline reconstruction.

## 3. Reconstruction and event selection

Events are selected using a trigger requiring two muons of opposite charge, each with  $p_T^\mu > 4$  GeV, successfully fitted to a common vertex. The  $\mu^+\mu^-$  mass range accepted by the trigger, 8–12 GeV, includes the  $\gamma(1S)$ ,  $\gamma(2S)$ , and  $\gamma(3S)$  signal peaks.

In the offline reconstruction for this analysis, muon reconstruction relies on a statistical combination of an MS track and an ID track. The selected muons are restricted to  $|\eta^\mu| < 2.3$ , ensuring high-quality tracking and a reduction of fake muon candidates. This restriction also removes regions of strongly varying efficiency and acceptance.

The reconstruction of  $\pi^+\pi^-\gamma(1S)$  candidates begins with pairs of oppositely charged muon candidates that satisfy the same kinematic conditions used by the trigger, and have  $\geq 2$  pixel and  $\geq 6$  silicon microstrip detector hits. Each pair is subjected to a common vertex fit, and a loose chi-square selection is imposed to exclude very poor candidates. Any dimuon with an invariant mass within

350 MeV of the  $\gamma(1S)$  mass [23],  $m_{1S}$ , is retained and considered an  $\gamma(1S) \rightarrow \mu^+\mu^-$  candidate. To confirm that this pair is the same as that used in the trigger, the reconstructed and trigger-level muons are required to match with  $\Delta R < 0.01$ .

In the remainder of the event, dipion candidates are formed from oppositely charged pions with  $|\eta^\pi| < 2.5$ , each required to have  $\geq 1$  pixel hits,  $\geq 6$  silicon microstrip hits, and  $p_T^\pi > 400$  MeV; no other requirements (such as lepton vetoes) are imposed. The  $\gamma(1S)$  candidate and the dipion system are combined by performing a four-track common-vertex fit, with the  $\mu^+\mu^-$  mass constrained to  $m_{1S}$ , and tracks assigned  $\mu$  or  $\pi$  masses as appropriate. This significantly improves the mass resolution: for example, in the  $\gamma(2S)$  simulation, the RMS improves from 142 MeV to 9.7 MeV. The  $\pi^+\pi^-\gamma(1S)$  vertex fit is required to have a chi-square less than 20; this is 95% efficient for  $\gamma(2S)$  decays but reduces background by a factor of  $\sim 10$ . All remaining  $\pi^+\pi^-\gamma(1S)$  candidates with invariant masses up to 11.2 GeV are retained.

For a state decaying to  $\pi^+\pi^-\gamma(1S)$ , the acceptance  $\mathcal{A}$  is defined as the fraction of decays where both muons have  $p_T^\mu > 4$  GeV, both pions have  $p_T^\pi > 400$  MeV, and all four particles are within  $|\eta| < 2.5$ . States with  $p_T < 5$  GeV or rapidity  $|y| > 2.4$  have very low acceptance, so candidates in these regions (which also suffer from high background) are excluded.

## 4. Data and simulation samples

The techniques adopted in this analysis were developed using  $\sqrt{s} = 7$  TeV data collected in 2011 and simulation samples generated under the same running conditions, with particular attention to  $\gamma(2S) \rightarrow \pi^+\pi^-\gamma(1S)$  mass and  $(|y|, p_T)$  distributions. Backgrounds due to inclusive  $\gamma(1S)$  production and combinatorial  $\mu^+\mu^-$  sources were studied using  $\mu^+\mu^-$  sideband and  $\mu^\pm\mu^\pm$  same-sign samples, and found to be featureless above 9.8 GeV.

The results presented here are based on data from the 2012  $\sqrt{s} = 8$  TeV  $pp$  run at the LHC; standard data-quality criteria are used to ensure efficient detector performance. The trigger employed in the event selection was subject to an instantaneous luminosity-dependent prescale factor,<sup>2</sup> and provided an integrated luminosity of  $\mathcal{L} = 16.2 \text{ fb}^{-1}$ . The resulting data sample includes over 10 million  $\mu^+\mu^-$  combinations: a fit to the sample finds  $(6.00 \pm 0.01) \times 10^6$  from  $\gamma(1S)$  decays,  $(0.200 \pm 0.002) \times 10^6$  from  $\gamma(2S)$  decays, and the remainder from combinatorial background. Given the reconstruction and event selection choices described in the previous section, each dimuon gives rise (on average) to  $19.5 \pi^+\pi^-\gamma(1S)$  candidates with invariant mass below 11.2 GeV. A procedure to select at most one candidate per event was considered. However, after the adoption of the binning approach described in Section 5, candidate selection was found to worsen the expected sensitivity to a hypothetical  $X_b$  signal. Therefore, all  $\pi^+\pi^-\gamma(1S)$  candidates are retained for analysis.

Simulated samples are used to optimise the selections, model signal decays, develop fitting models, and calculate efficiencies. Individual samples are used for the  $\gamma(2S)$ ,  $\gamma(3S)$ ,  $\gamma(1^3D_J)$  triplet,  $\gamma(10860)$ , and for two hypothetical  $X_b$  masses, 10233 MeV and 10561 MeV. Production is modelled with the AU2 [28] tune of PYTHIA 8.170 [29] and the CTEQ6L1 [30] parton distribution functions. In the decay to  $\pi^+\pi^-\gamma(1S)$ , the three-body phase space is uniformly sampled. Isotropic spin alignment of the parent is assumed: for  $\gamma(2S)$  and  $\gamma(3S)$  this is supported by a recent CMS measurement [31]. Passage of particles through the detector is

<sup>2</sup> An un-prescaled trigger with a  $p_T^\mu = 6$  GeV threshold on the muons was also considered, but found to result in a lower sensitivity using the optimisation procedure described in Section 5.

<sup>1</sup> ATLAS uses a right-handed coordinate system with its origin at the nominal interaction point (IP) in the centre of the detector, the  $x$ -axis pointing to the centre of the LHC ring, and the  $z$ -axis along the beam pipe; the  $y$ -axis points upward. Cylindrical coordinates  $(r, \phi)$  are used in the transverse plane;  $\phi$  is the azimuthal angle around the beam pipe. Pseudorapidity and transverse momentum are defined in terms of the polar angle  $\theta$  as  $\eta = -\ln(\tan\theta/2)$  and  $p_T = p \sin\theta$ . The  $(\eta, \phi)$  distance between two particles is defined as  $\Delta R = \sqrt{(\Delta\eta)^2 + (\Delta\phi)^2}$ . For a particle with momentum  $\vec{p} = (p_x, p_y, p_z)$  and energy  $E$ , the rapidity is defined as  $y = 0.5 \ln([E + p_z]/[E - p_z])$ .

simulated with Atfast II [32], supplementing GEANT4 [33,34] with a parameterised calorimeter response.

Simulated kinematic distributions of the final-state pions and muons are sensitive to mismodelling of the  $p_T$  and  $y$  distributions of the parent state. ATLAS has measured doubly differential production cross sections for  $\Upsilon(2S)$  and  $\Upsilon(3S)$  at 7 TeV in  $p_T$  bins up to 70 GeV for  $|y| < 1.2$  and  $1.2 < |y| < 2.25$  [35]. These results can be extended to  $|y| < 2.4$  assuming flat rapidity dependence, and up to  $p_T = 100$  GeV using CMS measurements [36]. The resulting  $\Upsilon(2S)$  cross section is compared to the production kinematics of a 7 TeV simulation using the 2011 ATLAS tune [37] of PYTHIA 6.4 [38]. The ratio of the two in  $(|y|, p_T)$  bins defines *production weights*, which are applied to the 8 TeV  $\Upsilon(2S)$  simulated sample, assuming that PYTHIA correctly models the increase in cross section with  $\sqrt{s}$ . The same procedure is used for the  $\Upsilon(3S)$ ; for other masses, linear extrapolation of the  $\Upsilon(2S)$  and  $\Upsilon(3S)$  weights is used. The simulated  $\Upsilon(2S)$  and  $\Upsilon(3S) \rightarrow \pi^+\pi^-\Upsilon(1S)$  samples are further reweighted to match dipion mass distributions observed by CLEO [39,40], to allow comparison of simulation and data.

## 5. Fitting strategy

Due to the low pion momentum in the  $\pi^+\pi^-\Upsilon(1S)$  rest frame, and the reconstruction threshold of  $p_T^\pi = 400$  MeV in the laboratory frame, true  $\pi^+\pi^-\Upsilon(1S)$  decays are preferentially reconstructed if the parent has large  $p_T$  or small  $\theta^*$  (defined as the angle, in the parent rest frame, between the dipion momentum and the lab-frame parent momentum). In background candidates, the dipion and dimuon systems are unrelated, and the dipions typically have low  $p_T^{\pi\pi}$  in the lab; boosting to the  $\pi^+\pi^-\mu^+\mu^-$  frame yields large values of  $\theta^*$ , with a broad distribution around  $\cos\theta^* = 0$ . In the  $(p_T, \cos\theta^*)$  plane, then, the ratio of signal to background candidates is largest in the upper-right region, and smallest in the lower-left.

The mass resolution and background shape differ at central and forward rapidities, so the analysis is performed in bins of  $|y| < 1.2$  (barrel) and  $1.2 < |y| < 2.4$  (endcap). Several possible ways to exploit the  $(p_T, \cos\theta^*)$  discrimination were considered, including further binning, a diagonal cut in  $(p_T, \cos\theta^*)$ , and a requirement on the  $\Delta R$  between the  $\Upsilon(1S)$  and each pion (as used by CMS [25]). The choice of method was based on optimising the expected significance for a weak signal at a mass of 10561 MeV. In the final approach eight analysis bins of varying sensitivity are used, formed from combinations of high and low  $|y|$ , high and low  $p_T$ , and high and low  $\cos\theta^*$ . The optimal bin boundaries for  $p_T$  and  $\cos\theta^*$  were determined to be 20 GeV and 0, respectively.

The  $\pi^+\pi^-\Upsilon(1S)$  mass distribution for the most sensitive bin ( $|y| < 1.2$ ,  $p_T > 20$  GeV,  $\cos\theta^* > 0$ ) is shown in Fig. 1. The  $\Upsilon(2S)$  and  $\Upsilon(3S)$  peaks are clearly visible, but no other peaks are apparent. The background in this bin decreases with mass above the  $\Upsilon(3S)$ , whereas the fraction of signal events falling in this bin is constant above the  $\Upsilon(3S)$  in the simulation. Thus higher sensitivity is expected at larger masses.

Based on the simulation samples, the fraction of signal in the barrel region  $|y| < 1.2$  is independent of mass with an average value of  $S_{|y|} = 0.606 \pm 0.004$ . Within the barrel, the fraction of signal with  $p_T < 20$  GeV develops smoothly with mass and can be characterised by an analytic turn-on curve  $S_{p_T}^b(m) = a/(1 + e^{-b(m-c)})$ . In the endcap, the dependence is described by  $S_{p_T}^{ec}(m)$ , which has the same functional form as  $S_{p_T}^b(m)$  with different values for the parameters. Similarly, within each  $(|y|, p_T)$  bin the fraction of the signal with  $\cos\theta^* < 0$  is modelled with a quadratic function,  $S_{\cos\theta^*}^{(i)}(m) = a + bm + cm^2$ , where  $i = 1-4$  labels the bin. These  $S$  functions, seven in total, are referred to below

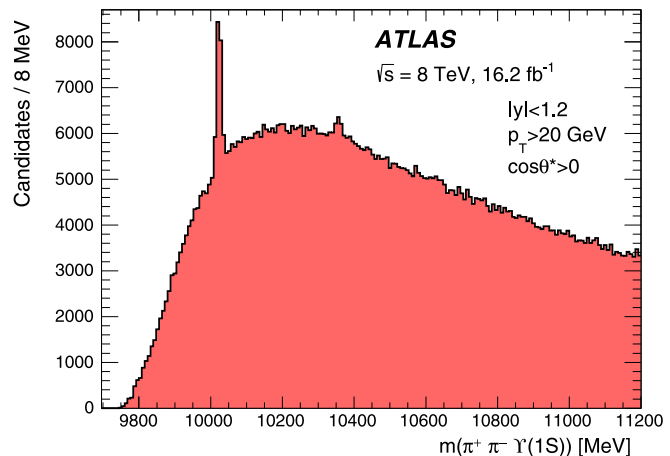


Fig. 1. The  $\pi^+\pi^-\Upsilon(1S)$  invariant mass distribution in the kinematic bin most sensitive to an  $X_b$  signal:  $|y| < 1.2$ ,  $p_T > 20$  GeV, and  $\cos\theta^* > 0$ . The only apparent peaks are at the masses of the  $\Upsilon(2S)$  (10023 MeV) and  $\Upsilon(3S)$  (10355 MeV).

as *splitting functions*. At any specified mass, the signal yield fraction in any particular  $(|y|, p_T, \cos\theta^*)$  bin can be calculated from an appropriately chosen product of three of these and their complements,  $(1 - S)$ . For example, the fraction in the bin ( $|y| < 1.2$ ,  $p_T > 20$  GeV,  $\cos\theta^* < 0$ ) is given by  $S_{|y|} \cdot (1 - S_{p_T}^b(m)) \cdot S_{\cos\theta^*}^{(3)}(m)$ .

The shape of the signal peaks reflects the detector resolution, which differs between the barrel and endcap, and varies as a function of mass; in a given rapidity bin at a given mass, a single function can be used across the whole  $(p_T, \cos\theta^*)$  range. In each rapidity bin, the signal is fitted using two Gaussians with a common mean, a narrow component fraction  $f$ , and a ratio  $r$  of broad to narrow widths. These parameters are found to be independent of mass, and are fixed to the average values across the simulation samples. The remaining parameter, the width of the narrow component,  $\sigma$ , depends linearly on the mass of the parent state. Together with the splitting functions defined above, this allows the signal shape and the fraction of the signal falling in each of the analysis bins to be determined for any  $X_b$  mass.

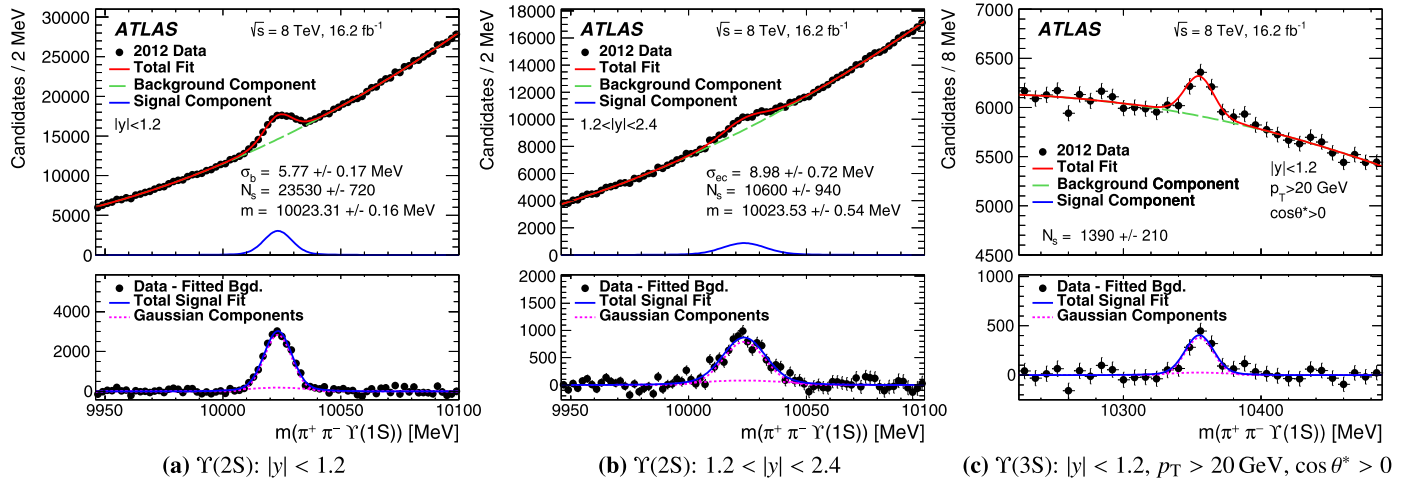
Searches for the production of the  $\Upsilon(10860)$  and  $\Upsilon(11020)$ , which have natural widths larger than the experimental resolution, are also performed. Each of these states is modelled as a Breit–Wigner convolved with the mass-dependent signal shape described in the previous paragraph, representing the detector resolution. The fractions of the  $\Upsilon(10860)$  signal falling in the eight analysis bins are extracted from the simulation, while those for the  $\Upsilon(11020)$  are determined by extrapolation.

All fits presented here are binned, extended maximum-likelihood fits in a local region around the mass of interest. The bin width is 2 MeV in all fits. The background is described by a linear combination of Chebychev polynomials up to second-order, with independent parameters in each analysis bin, unless otherwise specified.

## 6. Results for the $\Upsilon(2S)$ and $\Upsilon(3S)$

Fits to the  $\pi^+\pi^-\Upsilon(1S)$  spectrum near the  $\Upsilon(2S)$  are first performed separately in barrel and endcap bins, across the full  $(p_T, \cos\theta^*)$  range, with signal mass and width parameters free in the fit (Figs. 2a and 2b). In both cases, the mass is consistent with the world average for the  $\Upsilon(2S)$ , and the  $\sigma$  parameters are within uncertainties of the values fitted to the  $\Upsilon(2S)$  simulation.

Signal shape parameters are then fixed to simulated values to reduce uncertainties, and a separate fit is performed on each analysis bin. The fraction of signal decays falling in the barrel is measured to be  $0.67 \pm 0.04$ , compared to  $0.606 \pm 0.004$  in the simula-



**Fig. 2.** The measured  $\pi^+\pi^-\Upsilon(1S)$  invariant mass (data points), together with fits (red solid curves) to the  $\Upsilon(2S)$  peak in the (a) barrel and (b) endcap, with signal mass ( $m$ ) and width parameters ( $\sigma_b$  and  $\sigma_{ec}$ , respectively) free; and (c) a fit to the  $\Upsilon(3S)$  peak in the most sensitive bin, with zero suppressed on the vertical scale.  $N_s$  is the fitted signal yield in each bin. The background component (green, long-dashed) and the signal component (peaked blue curve) are also shown separately. In each case, the lower panel shows background-subtracted data, with the total signal function (blue solid) and its two Gaussian components (pink, short-dashed curves) overlaid. (For interpretation of the references to color in this figure legend, the reader is referred to the web version of this article.)

tion; similar results were seen in 7 TeV data. The ratio  $0.67/0.606$  is used to rescale the barrel fraction in the simulation.

The total fitted yield  $N_{2S} = 34300 \pm 800$  is consistent with

$$N_2^{\text{expected}} = (\sigma\mathcal{B})_{2S} \cdot \mathcal{L} \cdot \mathcal{A} \cdot \epsilon = 33300 \pm 2500, \quad (1)$$

where the product of the cross section and branching fraction,  $(\sigma\mathcal{B})_{2S}$ , is estimated from the extended cross-section measurement (see Section 4) using world-average values for the branching fractions  $\mathcal{B}(\Upsilon(1S) \rightarrow \mu^+\mu^-)$  and  $\mathcal{B}(\Upsilon(2S) \rightarrow \pi^+\pi^-\Upsilon(1S))$  [23]. The shape of the doubly differential cross section is also used to calculate the acceptance,  $\mathcal{A} = (1.442 \pm 0.004)\%$ , assuming the CLEO dipion mass spectrum [39] and isotropic signal decays. The reconstruction efficiency for decays within the acceptance,  $\epsilon = 0.283 \pm 0.002$ , is taken from the  $\Upsilon(2S)$  simulation. To test kinematic distributions in the  $\Upsilon(2S)$  simulation, signal fractions in the eight analysis bins are checked against their expected values, and are found to be consistent within statistical uncertainties.

A simultaneous fit to the analysis bins is also performed at the  $\Upsilon(3S)$  mass, with reduced  $\chi^2 = 1.0$  and statistical significance (from the likelihood-ratio test statistic)  $z = 8.7$ . An individual fit to the most sensitive bin (shown in Fig. 2(c), with larger binning to emphasise the peak) has significance  $z = 6.5$ . The total  $\Upsilon(3S)$  yield,  $11600 \pm 1300$ , agrees with the prediction estimated analogously to Eq. (1),  $11400 \pm 1500$ .

## 7. Results for the $X_b$ search

### 7.1. Hypothesis tests

A hypothesis test for the presence of an  $X_b$  peak is performed every 10 MeV from 10 GeV to 11 GeV, assuming a narrow state<sup>3</sup> that has a differential cross section with a  $(|y|, p_T)$  distribution similar to that of the  $\Upsilon(2S)$  or  $\Upsilon(3S)$ , decaying according to three-body phase space. The signal shape and bin splittings are treated as described in Section 5. At each mass, a simultaneous fit to the analysis bins is performed in a range  $m \pm 8\sigma_{ec}$ , where  $\sigma_{ec}(m)$  is the width of the narrow signal component in the endcap: the window varies from  $\pm 72$  MeV at 10 GeV to  $\pm 224$  MeV at 10.9 GeV;

near 11 GeV,  $m - 8\sigma_{ec} < m < 11.2$  GeV is used. When fitting near the  $\Upsilon(2S)$  and  $\Upsilon(3S)$  regions, the lineshapes for these signal components are added to the background model, with normalisations governed by Gaussian constraints to the values obtained from the fits given in Section 6. In the immediate vicinity of the peaks,  $m_{2S,3S} \pm 4\sigma_b$ , no search is performed; the width of the narrow signal component in the barrel,  $\sigma_b$ , is 5.66 MeV (9.37 MeV) at the  $\Upsilon(2S)$  ( $\Upsilon(3S)$ ) mass. This reduces the analysis range to 10.05–10.31 and 10.40–11.00 GeV.

At each mass, the  $p$ -value is extracted using the asymptotic formula [41] for the  $q_0$  statistic, a modification of the standard likelihood ratio (see Fig. 3). No evidence for new states with local significance  $z \geq 3$  is found.

The expected number of  $X_b$  events can be written as

$$N = N_{2S} \cdot R \cdot \frac{\mathcal{A}}{\mathcal{A}_{2S}} \cdot \frac{\epsilon}{\epsilon_{2S}}, \quad (2)$$

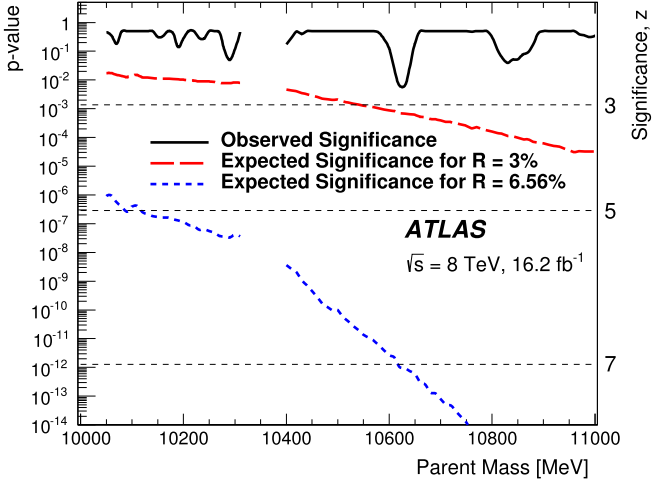
where  $R \equiv (\sigma\mathcal{B})/(\sigma\mathcal{B})_{2S}$ , the production rate relative to that of the  $\Upsilon(2S)$ . The efficiency,  $\epsilon$ , as a function of mass is determined by fitting a function  $a + b/(1 + e^{-c(m-d)})$  to the values from the simulated samples. The acceptance increases with the mass of the parent state due to the increased energy available to the pions and muons. Additionally, the measured production spectra of the  $\Upsilon(1S)$ ,  $\Upsilon(2S)$ , and  $\Upsilon(3S)$  states [35] are progressively harder in  $p_T$  with increasing mass. The acceptance for a hypothetical  $X_b$  state of arbitrary mass is estimated here by linear extrapolation of calculations performed at the  $\Upsilon(2S)$  and  $\Upsilon(3S)$ , using the measured production spectra of these states. The ratio  $(\mathcal{A}\epsilon)/(\mathcal{A}\epsilon)_{2S}$  rises from 1.1 at  $m = 10$  GeV to 7.5 at  $m = 11$  GeV.

Using Eq. (2) and the formalism from Ref. [41], the expected significance for a relative production rate  $R = 6.56\%$  (the value of the analogous quantity for the  $X(3872)$  [5]) is calculated as a function of mass, shown as the dashed blue line in Fig. 3; it exceeds  $5\sigma$  for  $m \gtrsim 10.12$  GeV. Expectations for a weaker signal with  $R = 3\%$  are also shown (long-dashed red line). Given the null result, upper limits are calculated on  $R$  after modifying the fit to include systematic uncertainties.

### 7.2. Systematic uncertainties

The upper limit calculation depends indirectly on signal and background fitting parameters, including the fraction of the signal

<sup>3</sup> Here, a narrow state refers to one whose natural width is much smaller than the experimental resolution.



**Fig. 3.** The solid curve shows the observed local  $p$ -value for the background-only hypothesis (left scale), and the corresponding significance,  $z$ , of a peak in  $\pi^+\pi^-\gamma(1S)$  (right scale), as a function of the mass of a hypothetical  $X_b$  parent state. Also shown are the expected values for the case of a signal with relative production rates  $(\sigma\mathcal{B})/(\sigma\mathcal{B})_{2S}$  of 3% (red, long-dashed) and 6.56% (blue, dashed curve).

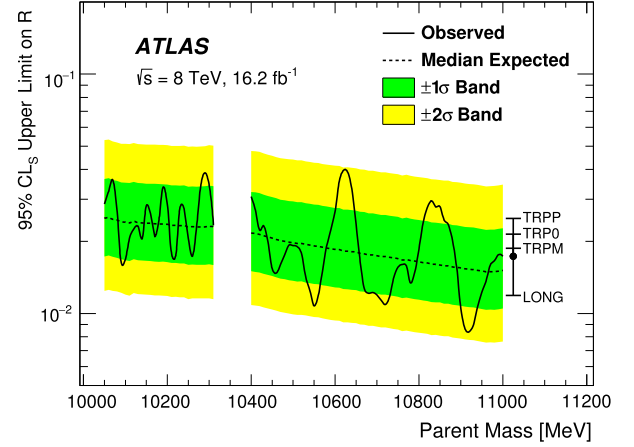
falling in each of the analysis bins. From Eq. (2), the upper limit on  $R$  is proportional to the inverse fitted  $\gamma(2S)$  yield,  $N_{2S}^{-1}$ , and the ratios  $\mathcal{A}_{2S}/\mathcal{A}$  and  $\epsilon_{2S}/\epsilon$ . For each source of systematic uncertainty, the impact on these factors is quantified to find the maximum shift across the mass range. These are then summed in quadrature and included in the fit as Gaussian-constrained nuisance parameters.

The  $X(3872) \rightarrow \pi^+\pi^-J/\psi$  dipion mass distribution favours high mass [6,9]; for a potential hidden-beauty counterpart this distribution is unknown. For  $\psi(2S) \rightarrow \pi^+\pi^-J/\psi$  [42], and both  $\gamma(2S)$  [39] and  $\gamma(4S) \rightarrow \pi^+\pi^-\gamma(1S)$  [43,44], the dipion mass distributions are concentrated near the upper boundary; those for  $Y(4260) \rightarrow \pi^+\pi^-J/\psi$  [45] and  $\gamma(3S) \rightarrow \pi^+\pi^-\gamma(1S)$  [40] are double-humped. The results quoted here assume decay according to three-body phase space;  $\gamma(2S)$ - and  $\gamma(3S)$ -like distributions change the splitting functions by up to 35%, decrease the efficiency ratio by up to 17%, and produce modest changes in other parameters.

The next largest contribution is due to the linear extrapolation of the acceptance between the  $\gamma(2S)$  and  $\gamma(3S)$  values. Alternative extrapolations between the  $\gamma(1S)$  and  $\gamma(2S)$ , and between  $\gamma(1S)$  and  $\gamma(3S)$ , are also tried; the greatest change in the acceptance ratio, 12%, is assigned as the uncertainty.

The parameters of the efficiency, the splitting functions, and the widths of the narrow signal components  $\sigma_b$  and  $\sigma_{ec}$  as functions of mass, are varied by the uncertainties on their fitted values; alternative functional forms are also tried. In each case, the largest deviation is assigned as the systematic uncertainty. The use of production weights (described in Section 4) relies on assumptions regarding rapidity dependence, and evolution from  $\sqrt{s} = 7$  TeV to 8 TeV. Removing these weights produces a  $\sim 1\%$  change in efficiency ratio (most of the differences cancel), but changes the values of the splitting functions by up to 8%.

Data versus simulation differences in the  $\gamma(2S)$  width parameters in the barrel and endcap (1.9% and 4.2%, respectively) are incorporated as a source of uncertainty, as is the statistical uncertainty on the averages used for signal shape parameters  $f$  and  $r$  (0.5–1.4%). The background shape model is also altered, allowing a third-order term comparable in size to typical values of the second-order terms. Finally, uncertainties on  $N_{2S}$  and the barrel/endcap scaling factor are assigned based on uncertainties from the  $\gamma(2S)$  fits.



**Fig. 4.** Observed 95%  $CL_s$  upper limits (solid line) on the relative production rate  $R = (\sigma\mathcal{B})/(\sigma\mathcal{B})_{2S}$  of a hypothetical  $X_b$  parent state decaying isotropically to  $\pi^+\pi^-\gamma(1S)$ , as a function of mass. The median expectation (dashed) and the corresponding  $\pm 1\sigma$  and  $\pm 2\sigma$  bands (green and yellow respectively) are also shown. The bar on the right shows typical shifts under alternative  $X_b$  spin-alignment scenarios, relative to the isotropic ('FLAT') case shown with the solid point.

### 7.3. Upper limit calculation

Upper limits are evaluated at the 95% confidence level using the  $CL_s$  method by implementing asymptotic formulae for the  $\tilde{q}_\mu$  statistic [41]. The results (Fig. 4, solid line) range between  $R = 0.8\%$  and 4.0%. Median expected upper limits assuming background only (dashed line), and corresponding  $\pm 1\sigma$  and  $\pm 2\sigma$  bands are also shown. These limits include the effect of systematic uncertainties: their inclusion increased the observed limits by up to 13% and inflated the  $\pm 1\sigma$  band by 9.5–25%, depending on the  $X_b$  mass.

As a check, upper limits are recalculated with modified fitting ranges ( $m \pm 7\sigma_{ec}$  and  $m \pm 9\sigma_{ec}$ ) and doubled bin widths in the  $\pi^+\pi^-\gamma(1S)$  mass distributions: shifts are small compared to the  $\pm 1\sigma$  bands. If an  $\gamma(2S)$ -like  $m_{\pi^+\pi^-}$  distribution is assumed (cf. CMS [25]), expected upper limits increase: the fractional change is +17% at 10.1 GeV, and  $\sim +5\%$  for  $m > 10.4$  GeV.

These results exclude  $X_b$  states with  $R = 6.56\%$  for masses 10.05–10.31 GeV and 10.40–11.00 GeV. The expected upper limits are more restrictive than those from CMS above  $m \sim 10.1$  GeV, and improve as a function of mass; the discrimination in  $(p_T, \cos\theta^*)$ , exploited by the binning method, becomes increasingly important as mass increases.

If an  $X_b$  state exists and lies within the range of masses to which this analysis is sensitive, its production cross section and/or its branching fraction must be lower, relative to the  $\gamma(2S)$ , than that of the  $X(3872)$  relative to the  $\psi(2S)$ . There are arguments that the decay  $X_b \rightarrow \pi^+\pi^-\gamma(1S)$  should be suppressed, in the absence of the strong isospin-violating effects that are present for  $X(3872) \rightarrow \pi^+\pi^-J/\psi$  [46,47]. In this case the  $X_b$  would have more prominent decays to  $\pi^+\pi^-\chi_{b1}$ ,  $\pi^+\pi^-\pi^0\gamma(1S)$ , and other final states which are relatively difficult to reconstruct.

All results to this point assume that any hypothetical  $X_b$  production is unpolarised. Angular distributions of the  $\gamma(1S, 2S, 3S)$  states in  $pp$  collisions are consistent with unpolarised production [31], but the  $X_b$  spin-alignment is unknown and can have a strong impact on the efficiency ratio, acceptance ratio, and bin splitting fractions. Rather than including this as a systematic uncertainty, upper limits are recalculated under longitudinal ('LONG') and three transverse ('TRPP', 'TRP0', 'TRPM') spin-alignment scenarios [48]. Shifts in the upper limits (either up or down) depend only weakly on mass; the shift is smaller at large masses. In Fig. 4 the effect of each hypothesis is represented by a single number,

chosen as the maximum difference in the median expected significance from the unpolarised ('FLAT') case.

## 8. Results for the $\Upsilon(1^3D_J)$ triplet, $\Upsilon(10860)$ , and $\Upsilon(11020)$

The search described above does not account for the closely spaced  $\Upsilon(1^3D_J)$  triplet or the broad  $\Upsilon(10860)$  and  $\Upsilon(11020)$ . To fit for the  $\Upsilon(1^3D_J)$ , two extra peaks are added to the signal model. CLEO [18] and BaBar [19] have measured the  $\Upsilon(1^3D_2)$  mass, with an average of  $(10163.7 \pm 1.4)$  MeV, but the mass splitting within the triplet is unknown. Averaging over several models [49] leads to triplet masses 10156, 10164, and 10170 MeV (at 1 MeV precision). A fit is performed using these values, assuming independent normalisations but common signal shapes and bin splitting fractions. A significance of  $z = 0.12$  is found, with fitted yields  $-1000 \pm 3100$ ,  $600 \pm 1800$ , and  $800 \pm 2300$  for  $J = 1, 2$ , and 3. Reasonable changes to the mass splittings do not appreciably increase the significance, so there appears to be no evidence for  $\Upsilon(1^3D_J)$  production. Assuming that  $J = 2$  production dominates, or that the mass splitting is larger than the experimental resolution, the upper limit on  $R$  can be read from Fig. 4; combined with the measured  $\Upsilon(1^3D_2) \rightarrow \pi^+\pi^-\Upsilon(1S)$  branching fraction [19], this yields an upper limit on the relative cross section  $\sigma(pp \rightarrow \Upsilon(1^3D_2))/\sigma(pp \rightarrow \Upsilon(2S)) \leq 0.55$ .

The signal model for  $\Upsilon(10860)$  and  $\Upsilon(11020)$  is described in Section 5. Due to the large natural widths of these states, the fitting range is extended to 10.498–11.198 GeV and the background polynomial order increased to three. Significances  $z = 0.6$  and  $z = 0.3$  are found for  $\Upsilon(10860)$  and  $\Upsilon(11020)$ , for masses and widths fixed to world-average values [23]. As these parameters have large uncertainties, the significance is also calculated in a grid of  $m \pm 20$  MeV and  $\Gamma \pm \Delta\Gamma$ , where  $\Delta\Gamma$  is the uncertainty on the world-average width [23]. The largest significance for the  $\Upsilon(10860)$  is  $z = 1.1$  at  $m = 10856$  MeV and  $\Gamma = 55$  MeV. For the  $\Upsilon(11020)$ , the largest significance is  $z = 0.6$  at  $m = 11039$  MeV and  $\Gamma = 95$  MeV. Thus, no evidence for  $\Upsilon(10860)$  or  $\Upsilon(11020)$  production is found.

## 9. Conclusions

A search for a hidden-beauty analogue of the  $X(3872)$  is conducted by reconstructing  $\pi^+\pi^-\Upsilon(1S)(\rightarrow \mu^+\mu^-)$  events in  $16.2 \text{ fb}^{-1}$  of  $pp$  collision data recorded at  $\sqrt{s} = 8$  TeV by ATLAS at the LHC. To optimise the sensitivity of the search, the analysis is performed in eight bins of rapidity, transverse momentum, and the angle (in the rest frame of the parent state) between the dipion system and the laboratory-frame momentum of the parent. At each mass, the presence of a signal is tested by performing simultaneous fits to the nearby  $\pi^+\pi^-\Upsilon(1S)$  mass spectrum in these bins; no evidence for new narrow states is found for masses 10.05–10.31 GeV and 10.40–11.00 GeV. Upper limits are also set on the ratio  $R = [\sigma(pp \rightarrow X_b)\mathcal{B}(X_b \rightarrow \pi^+\pi^-\Upsilon(1S))]/[\sigma(pp \rightarrow \Upsilon(2S))\mathcal{B}(\Upsilon(2S) \rightarrow \pi^+\pi^-\Upsilon(1S))]$ , with results ranging from 0.8% to 4.0% depending on the  $X_b$  mass. The analogous ratio for the  $X(3872)$  is 6.56%: a value this large is excluded for all  $X_b$  masses considered. Separate fits to the  $\Upsilon(1^3D_J)$  triplet,  $\Upsilon(10860)$ , and  $\Upsilon(11020)$  also reveal no significant signals, and a  $\text{CL}_S$  upper limit of 0.55 is set on  $\sigma(pp \rightarrow \Upsilon(1^3D_2))/\sigma(pp \rightarrow \Upsilon(2S))$ .

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G. Aad<sup>84</sup>, B. Abbott<sup>112</sup>, J. Abdallah<sup>152</sup>, S. Abdel Khalek<sup>116</sup>, O. Abdinov<sup>11</sup>, R. Aben<sup>106</sup>, B. Abi<sup>113</sup>, M. Abolins<sup>89</sup>, O.S. AbouZeid<sup>159</sup>, H. Abramowicz<sup>154</sup>, H. Abreu<sup>153</sup>, R. Abreu<sup>30</sup>, Y. Abulaiti<sup>147a,147b</sup>, B.S. Acharya<sup>165a,165b,a</sup>, L. Adamczyk<sup>38a</sup>, D.L. Adams<sup>25</sup>, J. Adelman<sup>177</sup>, S. Adomeit<sup>99</sup>, T. Adye<sup>130</sup>, T. Agatonovic-Jovin<sup>13a</sup>, J.A. Aguilar-Saavedra<sup>125a,125f</sup>, M. Agustoni<sup>17</sup>, S.P. Ahlen<sup>22</sup>, F. Ahmadov<sup>64,b</sup>, G. Aielli<sup>134a,134b</sup>, H. Akerstedt<sup>147a,147b</sup>, T.P.A. Åkesson<sup>80</sup>, G. Akimoto<sup>156</sup>, A.V. Akimov<sup>95</sup>, G.L. Alberghi<sup>20a,20b</sup>, J. Albert<sup>170</sup>, S. Albrand<sup>55</sup>, M.J. Alconada Verzini<sup>70</sup>, M. Aleksa<sup>30</sup>, I.N. Aleksandrov<sup>64</sup>, C. Alexa<sup>26a</sup>, G. Alexander<sup>154</sup>, G. Alexandre<sup>49</sup>, T. Alexopoulos<sup>10</sup>, M. Alhroob<sup>165a,165c</sup>, G. Alimonti<sup>90a</sup>, L. Alio<sup>84</sup>, J. Alison<sup>31</sup>, B.M.M. Allbrooke<sup>18</sup>, L.J. Allison<sup>71</sup>, P.P. Allport<sup>73</sup>, J. Almond<sup>83</sup>, A. Aloisio<sup>103a,103b</sup>, A. Alonso<sup>36</sup>, F. Alonso<sup>70</sup>, C. Alpigiani<sup>75</sup>, A. Altheimer<sup>35</sup>, B. Alvarez Gonzalez<sup>89</sup>, M.G. Alviggi<sup>103a,103b</sup>, K. Amako<sup>65</sup>, Y. Amaral Coutinho<sup>24a</sup>, C. Amelung<sup>23</sup>, D. Amidei<sup>88</sup>, S.P. Amor Dos Santos<sup>125a,125c</sup>, A. Amorim<sup>125a,125b</sup>, S. Amoroso<sup>48</sup>, N. Amram<sup>154</sup>, G. Amundsen<sup>23</sup>, C. Anastopoulos<sup>140</sup>, L.S. Ancu<sup>49</sup>, N. Andari<sup>30</sup>, T. Andeen<sup>35</sup>, C.F. Anders<sup>58b</sup>, G. Anders<sup>30</sup>, K.J. Anderson<sup>31</sup>, A. Andreazza<sup>90a,90b</sup>, V. Andrei<sup>58a</sup>, X.S. Anduaga<sup>70</sup>, S. Angelidakis<sup>9</sup>, I. Angelozzi<sup>106</sup>, P. Anger<sup>44</sup>, A. Angerami<sup>35</sup>, F. Anghinolfi<sup>30</sup>, A.V. Anisenkov<sup>108,c</sup>, N. Anjos<sup>12</sup>, A. Annovi<sup>47</sup>, A. Antonaki<sup>9</sup>, M. Antonelli<sup>47</sup>, A. Antonov<sup>97</sup>, J. Antos<sup>145b</sup>, F. Anulli<sup>133a</sup>, M. Aoki<sup>65</sup>, L. Aperio Bella<sup>18</sup>, R. Apolle<sup>119,d</sup>, G. Arabidze<sup>89</sup>, I. Aracena<sup>144</sup>, Y. Arai<sup>65</sup>, J.P. Araque<sup>125a</sup>, A.T.H. Arce<sup>45</sup>, J.-F. Arguin<sup>94</sup>, S. Argyropoulos<sup>42</sup>, M. Arik<sup>19a</sup>, A.J. Armbruster<sup>30</sup>, O. Arnaez<sup>30</sup>, V. Arnal<sup>81</sup>, H. Arnold<sup>48</sup>, M. Arratia<sup>28</sup>, O. Arslan<sup>21</sup>, A. Artamonov<sup>96</sup>,

G. Artoni<sup>23</sup>, S. Asai<sup>156</sup>, N. Asbah<sup>42</sup>, A. Ashkenazi<sup>154</sup>, B. Åsman<sup>147a,147b</sup>, L. Asquith<sup>6</sup>, K. Assamagan<sup>25</sup>, R. Astalos<sup>145a</sup>, M. Atkinson<sup>166</sup>, N.B. Atlay<sup>142</sup>, B. Auerbach<sup>6</sup>, K. Augsten<sup>127</sup>, M. Auresseau<sup>146b</sup>, G. Avolio<sup>30</sup>, G. Azuelos<sup>94,e</sup>, Y. Azuma<sup>156</sup>, M.A. Baak<sup>30</sup>, A.E. Baas<sup>58a</sup>, C. Bacci<sup>135a,135b</sup>, H. Bachacou<sup>137</sup>, K. Bachas<sup>155</sup>, M. Backes<sup>30</sup>, M. Backhaus<sup>30</sup>, J. Backus Mayes<sup>144</sup>, E. Badescu<sup>26a</sup>, P. Bagiachi<sup>133a,133b</sup>, P. Bagnaia<sup>133a,133b</sup>, Y. Bai<sup>33a</sup>, T. Bain<sup>35</sup>, J.T. Baines<sup>130</sup>, O.K. Baker<sup>177</sup>, P. Balek<sup>128</sup>, F. Balli<sup>137</sup>, E. Banas<sup>39</sup>, Sw. Banerjee<sup>174</sup>, A.A.E. Bannoura<sup>176</sup>, V. Bansal<sup>170</sup>, H.S. Bansil<sup>18</sup>, L. Barak<sup>173</sup>, S.P. Baranov<sup>95</sup>, E.L. Barberio<sup>87</sup>, D. Barberis<sup>50a,50b</sup>, M. Barbero<sup>84</sup>, T. Barillari<sup>100</sup>, M. Barisonzi<sup>176</sup>, T. Barklow<sup>144</sup>, N. Barlow<sup>28</sup>, B.M. Barnett<sup>130</sup>, R.M. Barnett<sup>15</sup>, Z. Barnovska<sup>5</sup>, A. Baroncelli<sup>135a</sup>, G. Barone<sup>49</sup>, A.J. Barr<sup>119</sup>, F. Barreiro<sup>81</sup>, J. Barreiro Guimarães da Costa<sup>57</sup>, R. Bartoldus<sup>144</sup>, A.E. Barton<sup>71</sup>, P. Bartos<sup>145a</sup>, V. Bartsch<sup>150</sup>, A. Bassalat<sup>116</sup>, A. Basye<sup>166</sup>, R.L. Bates<sup>53</sup>, J.R. Batley<sup>28</sup>, M. Battaglia<sup>138</sup>, M. Battistin<sup>30</sup>, F. Bauer<sup>137</sup>, H.S. Bawa<sup>144,f</sup>, M.D. Beattie<sup>71</sup>, T. Beau<sup>79</sup>, P.H. Beauchemin<sup>162</sup>, R. Beccherle<sup>123a,123b</sup>, P. Bechtel<sup>21</sup>, H.P. Beck<sup>17</sup>, K. Becker<sup>176</sup>, S. Becker<sup>99</sup>, M. Beckingham<sup>171</sup>, C. Becot<sup>116</sup>, A.J. Beddall<sup>19c</sup>, A. Beddall<sup>19c</sup>, S. Bedikian<sup>177</sup>, V.A. Bednyakov<sup>64</sup>, C.P. Bee<sup>149</sup>, L.J. Beemster<sup>106</sup>, T.A. Beermann<sup>176</sup>, M. Begel<sup>25</sup>, K. Behr<sup>119</sup>, C. Belanger-Champagne<sup>86</sup>, P.J. Bell<sup>49</sup>, W.H. Bell<sup>49</sup>, G. Bella<sup>154</sup>, L. Bellagamba<sup>20a</sup>, A. Bellerive<sup>29</sup>, M. Bellomo<sup>85</sup>, K. Belotskiy<sup>97</sup>, O. Beltramello<sup>30</sup>, O. Benary<sup>154</sup>, D. Benchekroun<sup>136a</sup>, K. Bendtz<sup>147a,147b</sup>, N. Benekos<sup>166</sup>, Y. Benhammou<sup>154</sup>, E. Benhar Noccioli<sup>49</sup>, J.A. Benitez Garcia<sup>160b</sup>, D.P. Benjamin<sup>45</sup>, J.R. Bensinger<sup>23</sup>, K. Benslama<sup>131</sup>, S. Bentvelsen<sup>106</sup>, D. Berge<sup>106</sup>, E. Bergeaas Kuutmann<sup>167</sup>, N. Berger<sup>5</sup>, F. Berghaus<sup>170</sup>, J. Beringer<sup>15</sup>, C. Bernard<sup>22</sup>, P. Bernat<sup>77</sup>, C. Bernius<sup>78</sup>, F.U. Bernlochner<sup>170</sup>, T. Berry<sup>76</sup>, P. Berta<sup>128</sup>, C. Bertella<sup>84</sup>, G. Bertoli<sup>147a,147b</sup>, F. Bertolucci<sup>123a,123b</sup>, C. Bertsche<sup>112</sup>, D. Bertsche<sup>112</sup>, M.I. Besana<sup>90a</sup>, G.J. Besjes<sup>105</sup>, O. Bessidskaia<sup>147a,147b</sup>, M. Bessner<sup>42</sup>, N. Besson<sup>137</sup>, C. Betancourt<sup>48</sup>, S. Bethke<sup>100</sup>, W. Bhimji<sup>46</sup>, R.M. Bianchi<sup>124</sup>, L. Bianchini<sup>23</sup>, M. Bianco<sup>30</sup>, O. Biebel<sup>99</sup>, S.P. Bieniek<sup>77</sup>, K. Bierwagen<sup>54</sup>, J. Biesiada<sup>15</sup>, M. Biglietti<sup>135a</sup>, J. Bilbao De Mendizabal<sup>49</sup>, H. Bilokon<sup>47</sup>, M. Bindi<sup>54</sup>, S. Binet<sup>116</sup>, A. Bingul<sup>19c</sup>, C. Bini<sup>133a,133b</sup>, C.W. Black<sup>151</sup>, J.E. Black<sup>144</sup>, K.M. Black<sup>22</sup>, D. Blackburn<sup>139</sup>, R.E. Blair<sup>6</sup>, J.-B. Blanchard<sup>137</sup>, T. Blazek<sup>145a</sup>, I. Bloch<sup>42</sup>, C. Blocker<sup>23</sup>, W. Blum<sup>82,\*</sup>, U. Blumenschein<sup>54</sup>, G.J. Bobbink<sup>106</sup>, V.S. Bobrovnikov<sup>108,c</sup>, S.S. Bocchetta<sup>80</sup>, A. Bocci<sup>45</sup>, C. Bock<sup>99</sup>, C.R. Boddy<sup>119</sup>, M. Boehler<sup>48</sup>, T.T. Boek<sup>176</sup>, J.A. Bogaerts<sup>30</sup>, A.G. Bogdanchikov<sup>108</sup>, A. Bogouch<sup>91,\*</sup>, C. Boehm<sup>147a</sup>, J. Bohm<sup>126</sup>, V. Boisvert<sup>76</sup>, T. Bold<sup>38a</sup>, V. Boldea<sup>26a</sup>, A.S. Boldyrev<sup>98</sup>, M. Bomben<sup>79</sup>, M. Bona<sup>75</sup>, M. Boonekamp<sup>137</sup>, A. Borisov<sup>129</sup>, G. Borissov<sup>71</sup>, M. Borri<sup>83</sup>, S. Borroni<sup>42</sup>, J. Bortfeldt<sup>99</sup>, V. Bortolotto<sup>135a,135b</sup>, K. Bos<sup>106</sup>, D. Boscherini<sup>20a</sup>, M. Bosman<sup>12</sup>, H. Boterenbrood<sup>106</sup>, J. Boudreau<sup>124</sup>, J. Bouffard<sup>2</sup>, E.V. Bouhova-Thacker<sup>71</sup>, D. Boumediene<sup>34</sup>, C. Bourdarios<sup>116</sup>, N. Bousson<sup>113</sup>, S. Boutouil<sup>136d</sup>, A. Boveia<sup>31</sup>, J. Boyd<sup>30</sup>, I.R. Boyko<sup>64</sup>, I. Bozic<sup>13a</sup>, J. Bracinik<sup>18</sup>, A. Brandt<sup>8</sup>, G. Brandt<sup>15</sup>, O. Brandt<sup>58a</sup>, U. Bratzler<sup>157</sup>, B. Brau<sup>85</sup>, J.E. Brau<sup>115</sup>, H.M. Braun<sup>176,\*</sup>, S.F. Brazzale<sup>165a,165c</sup>, B. Brelier<sup>159</sup>, K. Brendlinger<sup>121</sup>, A.J. Brennan<sup>87</sup>, R. Brenner<sup>167</sup>, S. Bressler<sup>173</sup>, K. Bristow<sup>146c</sup>, T.M. Bristow<sup>46</sup>, D. Britton<sup>53</sup>, F.M. Brochu<sup>28</sup>, I. Brock<sup>21</sup>, R. Brock<sup>89</sup>, C. Bromberg<sup>89</sup>, J. Bronner<sup>100</sup>, G. Brooijmans<sup>35</sup>, T. Brooks<sup>76</sup>, W.K. Brooks<sup>32b</sup>, J. Brosamer<sup>15</sup>, E. Brost<sup>115</sup>, J. Brown<sup>55</sup>, P.A. Bruckman de Renstrom<sup>39</sup>, D. Bruncko<sup>145b</sup>, R. Bruneliere<sup>48</sup>, S. Brunet<sup>60</sup>, A. Bruni<sup>20a</sup>, G. Bruni<sup>20a</sup>, M. Bruschi<sup>20a</sup>, L. Bryngemark<sup>80</sup>, T. Buanes<sup>14</sup>, Q. Buat<sup>143</sup>, F. Bucci<sup>49</sup>, P. Buchholz<sup>142</sup>, R.M. Buckingham<sup>119</sup>, A.G. Buckley<sup>53</sup>, S.I. Buda<sup>26a</sup>, I.A. Budagov<sup>64</sup>, F. Buehrer<sup>48</sup>, L. Bugge<sup>118</sup>, M.K. Bugge<sup>118</sup>, O. Bulekov<sup>97</sup>, A.C. Bundock<sup>73</sup>, H. Burckhart<sup>30</sup>, S. Burdin<sup>73</sup>, B. Burghgrave<sup>107</sup>, S. Burke<sup>130</sup>, I. Burmeister<sup>43</sup>, E. Busato<sup>34</sup>, D. Büscher<sup>48</sup>, V. Büscher<sup>82</sup>, P. Bussey<sup>53</sup>, C.P. Buszello<sup>167</sup>, B. Butler<sup>57</sup>, J.M. Butler<sup>22</sup>, A.I. Butt<sup>3</sup>, C.M. Buttar<sup>53</sup>, J.M. Butterworth<sup>77</sup>, P. Butti<sup>106</sup>, W. Buttinger<sup>28</sup>, A. Buzatu<sup>53</sup>, M. Byszewski<sup>10</sup>, S. Cabrera Urbán<sup>168</sup>, D. Caforio<sup>20a,20b</sup>, O. Cakir<sup>4a</sup>, P. Calafiura<sup>15</sup>, A. Calandri<sup>137</sup>, G. Calderini<sup>79</sup>, P. Calfayan<sup>99</sup>, R. Calkins<sup>107</sup>, L.P. Caloba<sup>24a</sup>, D. Calvet<sup>34</sup>, S. Calvet<sup>34</sup>, R. Camacho Toro<sup>49</sup>, S. Camarda<sup>42</sup>, D. Cameron<sup>118</sup>, L.M. Caminada<sup>15</sup>, R. Caminal Armadans<sup>12</sup>, S. Campana<sup>30</sup>, M. Campanelli<sup>77</sup>, A. Campoverde<sup>149</sup>, V. Canale<sup>103a,103b</sup>, A. Canepa<sup>160a</sup>, M. Cano Bret<sup>75</sup>, J. Cantero<sup>81</sup>, R. Cantrill<sup>125a</sup>, T. Cao<sup>40</sup>, M.D.M. Capeans Garrido<sup>30</sup>, I. Caprini<sup>26a</sup>, M. Caprini<sup>26a</sup>, M. Capua<sup>37a,37b</sup>, R. Caputo<sup>82</sup>, R. Cardarelli<sup>134a</sup>, T. Carli<sup>30</sup>, G. Carlino<sup>103a</sup>, L. Carminati<sup>90a,90b</sup>, S. Caron<sup>105</sup>, E. Carquin<sup>32a</sup>, G.D. Carrillo-Montoya<sup>146c</sup>, J.R. Carter<sup>28</sup>, J. Carvalho<sup>125a,125c</sup>, D. Casadei<sup>77</sup>, M.P. Casado<sup>12</sup>, M. Casolino<sup>12</sup>, E. Castaneda-Miranda<sup>146b</sup>, A. Castelli<sup>106</sup>, V. Castillo Gimenez<sup>168</sup>, N.F. Castro<sup>125a</sup>, P. Catastini<sup>57</sup>, A. Catinaccio<sup>30</sup>, J.R. Catmore<sup>118</sup>, A. Cattai<sup>30</sup>, G. Cattani<sup>134a,134b</sup>, J. Caudron<sup>82</sup>, V. Cavaliere<sup>166</sup>, D. Cavalli<sup>90a</sup>, M. Cavalli-Sforza<sup>12</sup>, V. Cavasinni<sup>123a,123b</sup>, F. Ceradini<sup>135a,135b</sup>, B.C. Cerio<sup>45</sup>, K. Cerny<sup>128</sup>, A.S. Cerqueira<sup>24b</sup>, A. Cerri<sup>150</sup>,



L. Cerrito<sup>75</sup>, F. Cerutti<sup>15</sup>, M. Cerv<sup>30</sup>, A. Cervelli<sup>17</sup>, S.A. Cetin<sup>19b</sup>, A. Chafaq<sup>136a</sup>, D. Chakraborty<sup>107</sup>, I. Chalupkova<sup>128</sup>, P. Chang<sup>166</sup>, B. Chapleau<sup>86</sup>, J.D. Chapman<sup>28</sup>, D. Charfeddine<sup>116</sup>, D.G. Charlton<sup>18</sup>, C.C. Chau<sup>159</sup>, C.A. Chavez Barajas<sup>150</sup>, S. Cheatham<sup>86</sup>, A. Chegwiddden<sup>89</sup>, S. Chekanov<sup>6</sup>, S.V. Chekulaev<sup>160a</sup>, G.A. Chelkov<sup>64.g</sup>, M.A. Chelstowska<sup>88</sup>, C. Chen<sup>63</sup>, H. Chen<sup>25</sup>, K. Chen<sup>149</sup>, L. Chen<sup>33d,h</sup>, S. Chen<sup>33c</sup>, X. Chen<sup>146c</sup>, Y. Chen<sup>66</sup>, Y. Chen<sup>35</sup>, H.C. Cheng<sup>88</sup>, Y. Cheng<sup>31</sup>, A. Cheplakov<sup>64</sup>, R. Cherkaoui El Moursli<sup>136e</sup>, V. Chernyatin<sup>25,\*</sup>, E. Cheu<sup>7</sup>, L. Chevalier<sup>137</sup>, V. Chiarella<sup>47</sup>, G. Chiefari<sup>103a,103b</sup>, J.T. Childers<sup>6</sup>, A. Chilingarov<sup>71</sup>, G. Chiodini<sup>72a</sup>, A.S. Chisholm<sup>18</sup>, R.T. Chislett<sup>77</sup>, A. Chitan<sup>26a</sup>, M.V. Chizhov<sup>64</sup>, S. Chouridou<sup>9</sup>, B.K.B. Chow<sup>99</sup>, D. Chromek-Burckhart<sup>30</sup>, M.L. Chu<sup>152</sup>, J. Chudoba<sup>126</sup>, J.J. Chwastowski<sup>39</sup>, L. Chytka<sup>114</sup>, G. Ciapetti<sup>133a,133b</sup>, A.K. Ciftci<sup>4a</sup>, R. Ciftci<sup>4a</sup>, D. Cinca<sup>53</sup>, V. Cindro<sup>74</sup>, A. Ciochio<sup>15</sup>, P. Cirkovic<sup>13b</sup>, Z.H. Citron<sup>173</sup>, M. Citterio<sup>90a</sup>, M. Ciubancan<sup>26a</sup>, A. Clark<sup>49</sup>, P.J. Clark<sup>46</sup>, R.N. Clarke<sup>15</sup>, W. Cleland<sup>124</sup>, J.C. Clemens<sup>84</sup>, C. Clement<sup>147a,147b</sup>, Y. Coadou<sup>84</sup>, M. Cobal<sup>165a,165c</sup>, A. Coccaro<sup>139</sup>, J. Cochran<sup>63</sup>, L. Coffey<sup>23</sup>, J.G. Cogan<sup>144</sup>, J. Coggeshall<sup>166</sup>, B. Cole<sup>35</sup>, S. Cole<sup>107</sup>, A.P. Colijn<sup>106</sup>, J. Collot<sup>55</sup>, T. Colombo<sup>58c</sup>, G. Colon<sup>85</sup>, G. Compostella<sup>100</sup>, P. Conde Muiño<sup>125a,125b</sup>, E. Coniavitis<sup>48</sup>, M.C. Conidi<sup>12</sup>, S.H. Connell<sup>146b</sup>, I.A. Connelly<sup>76</sup>, S.M. Consonni<sup>90a,90b</sup>, V. Consorti<sup>48</sup>, S. Constantinescu<sup>26a</sup>, C. Conta<sup>120a,120b</sup>, G. Conti<sup>57</sup>, F. Conventi<sup>103a,i</sup>, M. Cooke<sup>15</sup>, B.D. Cooper<sup>77</sup>, A.M. Cooper-Sarkar<sup>119</sup>, N.J. Cooper-Smith<sup>76</sup>, K. Copic<sup>15</sup>, T. Cornelissen<sup>176</sup>, M. Corradi<sup>20a</sup>, F. Corriveau<sup>86.j</sup>, A. Corso-Radu<sup>164</sup>, A. Cortes-Gonzalez<sup>12</sup>, G. Cortiana<sup>100</sup>, G. Costa<sup>90a</sup>, M.J. Costa<sup>168</sup>, D. Costanzo<sup>140</sup>, D. Côté<sup>8</sup>, G. Cottin<sup>28</sup>, G. Cowan<sup>76</sup>, B.E. 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<sup>1</sup> Department of Physics, University of Adelaide, Adelaide, Australia

<sup>2</sup> Physics Department, SUNY Albany, Albany, NY, United States

<sup>3</sup> Department of Physics, University of Alberta, Edmonton, AB, Canada

<sup>4</sup> (a) Department of Physics, Ankara University, Ankara; (b) Department of Physics, Gazi University, Ankara; (c) Division of Physics, TOBB University of Economics and Technology, Ankara;

(d) Turkish Atomic Energy Authority, Ankara, Turkey

<sup>5</sup> LAPP, CNRS/IN2P3 and Université de Savoie, Annecy-le-Vieux, France

<sup>6</sup> High Energy Physics Division, Argonne National Laboratory, Argonne, IL, United States

<sup>7</sup> Department of Physics, University of Arizona, Tucson, AZ, United States

<sup>8</sup> Department of Physics, The University of Texas at Arlington, Arlington, TX, United States

<sup>9</sup> Physics Department, University of Athens, Athens, Greece

<sup>10</sup> Physics Department, National Technical University of Athens, Zografou, Greece

<sup>11</sup> Institute of Physics, Azerbaijan Academy of Sciences, Baku, Azerbaijan

<sup>12</sup> Institut de Física d'Altes Energies and Departament de Física de la Universitat Autònoma de Barcelona, Barcelona, Spain

<sup>13</sup> (a) Institute of Physics, University of Belgrade, Belgrade; (b) Vinca Institute of Nuclear Sciences, University of Belgrade, Belgrade, Serbia

<sup>14</sup> Department for Physics and Technology, University of Bergen, Bergen, Norway

<sup>15</sup> Physics Division, Lawrence Berkeley National Laboratory and University of California, Berkeley, CA, United States

<sup>16</sup> Department of Physics, Humboldt University, Berlin, Germany

<sup>17</sup> Albert Einstein Center for Fundamental Physics and Laboratory for High Energy Physics, University of Bern, Bern, Switzerland

<sup>18</sup> School of Physics and Astronomy, University of Birmingham, Birmingham, United Kingdom

<sup>19</sup> (a) Department of Physics, Bogazici University, Istanbul; (b) Department of Physics, Dogus University, Istanbul; (c) Department of Physics Engineering, Gaziantep University, Gaziantep, Turkey

<sup>20</sup> (a) INFN Sezione di Bologna; (b) Dipartimento di Fisica e Astronomia, Università di Bologna, Bologna, Italy

<sup>21</sup> Physikalisches Institut, University of Bonn, Bonn, Germany

<sup>22</sup> Department of Physics, Boston University, Boston, MA, United States

<sup>23</sup> Department of Physics, Brandeis University, Waltham, MA, United States

<sup>24</sup> (a) Universidade Federal do Rio De Janeiro COPPE/EE/IF, Rio de Janeiro; (b) Electrical Circuits Department, Federal University of Juiz de Fora (UFJF), Juiz de Fora; (c) Federal University of Sao Joao del Rei (UFSJ), Sao Joao del Rei; (d) Instituto de Física, Universidade de Sao Paulo, Sao Paulo, Brazil

<sup>25</sup> Physics Department, Brookhaven National Laboratory, Upton, NY, United States

<sup>26</sup> (a) National Institute of Physics and Nuclear Engineering, Bucharest; (b) National Institute for Research and Development of Isotopic and Molecular Technologies, Physics Department, Cluj Napoca; (c) University Politehnica Bucharest, Bucharest; (d) West University in Timisoara, Timisoara, Romania

<sup>27</sup> Departamento de Física, Universidad de Buenos Aires, Buenos Aires, Argentina

<sup>28</sup> Cavendish Laboratory, University of Cambridge, Cambridge, United Kingdom

<sup>29</sup> Department of Physics, Carleton University, Ottawa, ON, Canada

<sup>30</sup> CERN, Geneva, Switzerland

<sup>31</sup> Enrico Fermi Institute, University of Chicago, Chicago, IL, United States

<sup>32</sup> (a) Departamento de Física, Pontificia Universidad Católica de Chile, Santiago; (b) Departamento de Física, Universidad Técnica Federico Santa María, Valparaíso, Chile

<sup>33</sup> (a) Institute of High Energy Physics, Chinese Academy of Sciences, Beijing; (b) Department of Modern Physics, University of Science and Technology of China, Anhui; (c) Department of Physics, Nanjing University, Jiangsu; (d) School of Physics, Shandong University, Shandong; (e) Physics Department, Shanghai Jiao Tong University, Shanghai, China

<sup>34</sup> Laboratoire de Physique Corpusculaire, Clermont Université and Université Blaise Pascal and CNRS/IN2P3, Clermont-Ferrand, France

<sup>35</sup> Nevis Laboratory, Columbia University, Irvington, NY, United States

<sup>36</sup> Niels Bohr Institute, University of Copenhagen, Copenhagen, Denmark

<sup>37</sup> (a) INFN Gruppo Collegato di Cosenza, Laboratori Nazionali di Frascati; (b) Dipartimento di Fisica, Università della Calabria, Rende, Italy



- 38 <sup>(a)</sup> AGH University of Science and Technology, Faculty of Physics and Applied Computer Science, Krakow; <sup>(b)</sup> Marian Smoluchowski Institute of Physics, Jagiellonian University, Krakow, Poland
- 39 The Henryk Niewodniczanski Institute of Nuclear Physics, Polish Academy of Sciences, Krakow, Poland
- 40 Physics Department, Southern Methodist University, Dallas, TX, United States
- 41 Physics Department, University of Texas at Dallas, Richardson, TX, United States
- 42 DESY, Hamburg and Zeuthen, Germany
- 43 Institut für Experimentelle Physik IV, Technische Universität Dortmund, Dortmund, Germany
- 44 Institut für Kern- und Teilchenphysik, Technische Universität Dresden, Dresden, Germany
- 45 Department of Physics, Duke University, Durham, NC, United States
- 46 SUPA – School of Physics and Astronomy, University of Edinburgh, Edinburgh, United Kingdom
- 47 INFN Laboratori Nazionali di Frascati, Frascati, Italy
- 48 Fakultät für Mathematik und Physik, Albert-Ludwigs-Universität, Freiburg, Germany
- 49 Section de Physique, Université de Genève, Geneva, Switzerland
- 50 <sup>(a)</sup> INFN Sezione di Genova; <sup>(b)</sup> Dipartimento di Fisica, Università di Genova, Genova, Italy
- 51 <sup>(a)</sup> E. Andronikashvili Institute of Physics, Iv. Javakishvili Tbilisi State University, Tbilisi; <sup>(b)</sup> High Energy Physics Institute, Tbilisi State University, Tbilisi, Georgia
- 52 II Physikalisches Institut, Justus-Liebig-Universität Giessen, Giessen, Germany
- 53 SUPA – School of Physics and Astronomy, University of Glasgow, Glasgow, United Kingdom
- 54 II Physikalisches Institut, Georg-August-Universität, Göttingen, Germany
- 55 Laboratoire de Physique Subatomique et de Cosmologie, Université Grenoble-Alpes, CNRS/IN2P3, Grenoble, France
- 56 Department of Physics, Hampton University, Hampton, VA, United States
- 57 Laboratory for Particle Physics and Cosmology, Harvard University, Cambridge, MA, United States
- 58 <sup>(a)</sup> Kirchhoff-Institut für Physik, Ruprecht-Karls-Universität Heidelberg, Heidelberg; <sup>(b)</sup> Physikalisches Institut, Ruprecht-Karls-Universität Heidelberg, Heidelberg; <sup>(c)</sup> ZITI Institut für technische Informatik, Ruprecht-Karls-Universität Heidelberg, Mannheim, Germany
- 59 Faculty of Applied Information Science, Hiroshima Institute of Technology, Hiroshima, Japan
- 60 Department of Physics, Indiana University, Bloomington, IN, United States
- 61 Institut für Astro- und Teilchenphysik, Leopold-Franzens-Universität, Innsbruck, Austria
- 62 University of Iowa, Iowa City, IA, United States
- 63 Department of Physics and Astronomy, Iowa State University, Ames, IA, United States
- 64 Joint Institute for Nuclear Research, JINR Dubna, Dubna, Russia
- 65 KEK, High Energy Accelerator Research Organization, Tsukuba, Japan
- 66 Graduate School of Science, Kobe University, Kobe, Japan
- 67 Faculty of Science, Kyoto University, Kyoto, Japan
- 68 Kyoto University of Education, Kyoto, Japan
- 69 Department of Physics, Kyushu University, Fukuoka, Japan
- 70 Instituto de Física La Plata, Universidad Nacional de La Plata and CONICET, La Plata, Argentina
- 71 Physics Department, Lancaster University, Lancaster, United Kingdom
- 72 <sup>(a)</sup> INFN Sezione di Lecce; <sup>(b)</sup> Dipartimento di Matematica e Fisica, Università del Salento, Lecce, Italy
- 73 Oliver Lodge Laboratory, University of Liverpool, Liverpool, United Kingdom
- 74 Department of Physics, Jožef Stefan Institute and University of Ljubljana, Ljubljana, Slovenia
- 75 School of Physics and Astronomy, Queen Mary University of London, London, United Kingdom
- 76 Department of Physics, Royal Holloway University of London, Surrey, United Kingdom
- 77 Department of Physics and Astronomy, University College London, London, United Kingdom
- 78 Louisiana Tech University, Ruston, LA, United States
- 79 Laboratoire de Physique Nucléaire et de Hautes Energies, UPMC and Université Paris-Diderot and CNRS/IN2P3, Paris, France
- 80 Fysiska institutionen, Lunds universitet, Lund, Sweden
- 81 Departamento de Física Teórica C-15, Universidad Autónoma de Madrid, Madrid, Spain
- 82 Institut für Physik, Universität Mainz, Mainz, Germany
- 83 School of Physics and Astronomy, University of Manchester, Manchester, United Kingdom
- 84 CPPM, Aix-Marseille Université and CNRS/IN2P3, Marseille, France
- 85 Department of Physics, University of Massachusetts, Amherst, MA, United States
- 86 Department of Physics, McGill University, Montreal, QC, Canada
- 87 School of Physics, University of Melbourne, Victoria, Australia
- 88 Department of Physics, The University of Michigan, Ann Arbor, MI, United States
- 89 Department of Physics and Astronomy, Michigan State University, East Lansing, MI, United States
- 90 <sup>(a)</sup> INFN Sezione di Milano; <sup>(b)</sup> Dipartimento di Fisica, Università di Milano, Milano, Italy
- 91 B.I. Stepanov Institute of Physics, National Academy of Sciences of Belarus, Minsk, Belarus
- 92 National Scientific and Educational Centre for Particle and High Energy Physics, Minsk, Belarus
- 93 Department of Physics, Massachusetts Institute of Technology, Cambridge, MA, United States
- 94 Group of Particle Physics, University of Montreal, Montreal, QC, Canada
- 95 P.N. Lebedev Institute of Physics, Academy of Sciences, Moscow, Russia
- 96 Institute for Theoretical and Experimental Physics (ITEP), Moscow, Russia
- 97 National Research Nuclear University MEPhI, Moscow, Russia
- 98 D.V. Skobel'syn Institute of Nuclear Physics, M.V. Lomonosov Moscow State University, Moscow, Russia
- 99 Fakultät für Physik, Ludwig-Maximilians-Universität München, München, Germany
- 100 Max-Planck-Institut für Physik (Werner-Heisenberg-Institut), München, Germany
- 101 Nagasaki Institute of Applied Science, Nagasaki, Japan
- 102 Graduate School of Science and Kobayashi–Maskawa Institute, Nagoya University, Nagoya, Japan
- 103 <sup>(a)</sup> INFN Sezione di Napoli; <sup>(b)</sup> Dipartimento di Fisica, Università di Napoli, Napoli, Italy
- 104 Department of Physics and Astronomy, University of New Mexico, Albuquerque, NM, United States
- 105 Institute for Mathematics, Astrophysics and Particle Physics, Radboud University Nijmegen/Nikhef, Nijmegen, Netherlands
- 106 Nikhef National Institute for Subatomic Physics and University of Amsterdam, Amsterdam, Netherlands
- 107 Department of Physics, Northern Illinois University, DeKalb, IL, United States
- 108 Budker Institute of Nuclear Physics, SB RAS, Novosibirsk, Russia
- 109 Department of Physics, New York University, New York, NY, United States
- 110 Ohio State University, Columbus, OH, United States
- 111 Faculty of Science, Okayama University, Okayama, Japan
- 112 Homer L. Dodge Department of Physics and Astronomy, University of Oklahoma, Norman, OK, United States
- 113 Department of Physics, Oklahoma State University, Stillwater, OK, United States
- 114 Palacký University, RCPTM, Olomouc, Czech Republic

- <sup>115</sup> Center for High Energy Physics, University of Oregon, Eugene, OR, United States
- <sup>116</sup> LAL, Université Paris-Sud and CNRS/IN2P3, Orsay, France
- <sup>117</sup> Graduate School of Science, Osaka University, Osaka, Japan
- <sup>118</sup> Department of Physics, University of Oslo, Oslo, Norway
- <sup>119</sup> Department of Physics, Oxford University, Oxford, United Kingdom
- <sup>120</sup> <sup>(a)</sup> INFN Sezione di Pavia; <sup>(b)</sup> Dipartimento di Fisica, Università di Pavia, Pavia, Italy
- <sup>121</sup> Department of Physics, University of Pennsylvania, Philadelphia, PA, United States
- <sup>122</sup> Petersburg Nuclear Physics Institute, Gatchina, Russia
- <sup>123</sup> <sup>(a)</sup> INFN Sezione di Pisa; <sup>(b)</sup> Dipartimento di Fisica E. Fermi, Università di Pisa, Pisa, Italy
- <sup>124</sup> Department of Physics and Astronomy, University of Pittsburgh, Pittsburgh, PA, United States
- <sup>125</sup> <sup>(a)</sup> Laboratório de Instrumentação e Física Experimental de Partículas – LIP, Lisboa; <sup>(b)</sup> Faculdade de Ciências, Universidade de Lisboa, Lisboa; <sup>(c)</sup> Department of Physics, University of Coimbra, Coimbra; <sup>(d)</sup> Centro de Física Nuclear da Universidade de Lisboa, Lisboa; <sup>(e)</sup> Departamento de Física, Universidade do Minho, Braga; <sup>(f)</sup> Departamento de Física Teórica y del Cosmos and CAFPE, Universidad de Granada, Granada (Spain); <sup>(g)</sup> Dep Física and CEFITEC de Faculdade de Ciências e Tecnologia, Universidade Nova de Lisboa, Caparica, Portugal
- <sup>126</sup> Institute of Physics, Academy of Sciences of the Czech Republic, Praha, Czech Republic
- <sup>127</sup> Czech Technical University in Prague, Praha, Czech Republic
- <sup>128</sup> Faculty of Mathematics and Physics, Charles University in Prague, Praha, Czech Republic
- <sup>129</sup> State Research Center Institute for High Energy Physics, Protvino, Russia
- <sup>130</sup> Particle Physics Department, Rutherford Appleton Laboratory, Didcot, United Kingdom
- <sup>131</sup> Physics Department, University of Regina, Regina, SK, Canada
- <sup>132</sup> Ritsumeikan University, Kusatsu, Shiga, Japan
- <sup>133</sup> <sup>(a)</sup> INFN Sezione di Roma; <sup>(b)</sup> Dipartimento di Fisica, Sapienza Università di Roma, Roma, Italy
- <sup>134</sup> <sup>(a)</sup> INFN Sezione di Roma Tor Vergata; <sup>(b)</sup> Dipartimento di Fisica, Università di Roma Tor Vergata, Roma, Italy
- <sup>135</sup> <sup>(a)</sup> INFN Sezione di Roma Tre; <sup>(b)</sup> Dipartimento di Matematica e Fisica, Università Roma Tre, Roma, Italy
- <sup>136</sup> <sup>(a)</sup> Faculté des Sciences Ain Chock, Réseau Universitaire de Physique des Hautes Energies – Université Hassan II, Casablanca; <sup>(b)</sup> Centre National de l’Energie des Sciences Techniques Nucleaires, Rabat; <sup>(c)</sup> Faculté des Sciences Semlalia, Université Cadi Ayyad, LPHEA – Marrakech; <sup>(d)</sup> Faculté des Sciences, Université Mohamed Premier and LPTPM, Oujda; <sup>(e)</sup> Faculté des sciences, Université Mohammed V – Agdal, Rabat, Morocco
- <sup>137</sup> DSM/IRFU (Institut de Recherches sur les Lois Fondamentales de l’Univers), CEA Saclay (Commissariat à l’Energie Atomique et aux Energies Alternatives), Gif-sur-Yvette, France
- <sup>138</sup> Santa Cruz Institute for Particle Physics, University of California Santa Cruz, Santa Cruz, CA, United States
- <sup>139</sup> Department of Physics, University of Washington, Seattle, WA, United States
- <sup>140</sup> Department of Physics and Astronomy, University of Sheffield, Sheffield, United Kingdom
- <sup>141</sup> Department of Physics, Shinshu University, Nagano, Japan
- <sup>142</sup> Fachbereich Physik, Universität Siegen, Siegen, Germany
- <sup>143</sup> Department of Physics, Simon Fraser University, Burnaby, BC, Canada
- <sup>144</sup> SLAC National Accelerator Laboratory, Stanford, CA, United States
- <sup>145</sup> <sup>(a)</sup> Faculty of Mathematics, Physics & Informatics, Comenius University, Bratislava; <sup>(b)</sup> Department of Subnuclear Physics, Institute of Experimental Physics of the Slovak Academy of Sciences, Kosice, Slovak Republic
- <sup>146</sup> <sup>(a)</sup> Department of Physics, University of Cape Town, Cape Town; <sup>(b)</sup> Department of Physics, University of Johannesburg, Johannesburg; <sup>(c)</sup> School of Physics, University of the Witwatersrand, Johannesburg, South Africa
- <sup>147</sup> <sup>(a)</sup> Department of Physics, Stockholm University; <sup>(b)</sup> The Oskar Klein Centre, Stockholm, Sweden
- <sup>148</sup> Physics Department, Royal Institute of Technology, Stockholm, Sweden
- <sup>149</sup> Departments of Physics & Astronomy and Chemistry, Stony Brook University, Stony Brook, NY, United States
- <sup>150</sup> Department of Physics and Astronomy, University of Sussex, Brighton, United Kingdom
- <sup>151</sup> School of Physics, University of Sydney, Sydney, Australia
- <sup>152</sup> Institute of Physics, Academia Sinica, Taipei, Taiwan
- <sup>153</sup> Department of Physics, Technion: Israel Institute of Technology, Haifa, Israel
- <sup>154</sup> Raymond and Beverly Sackler School of Physics and Astronomy, Tel Aviv University, Tel Aviv, Israel
- <sup>155</sup> Department of Physics, Aristotle University of Thessaloniki, Thessaloniki, Greece
- <sup>156</sup> International Center for Elementary Particle Physics and Department of Physics, The University of Tokyo, Tokyo, Japan
- <sup>157</sup> Graduate School of Science and Technology, Tokyo Metropolitan University, Tokyo, Japan
- <sup>158</sup> Department of Physics, Tokyo Institute of Technology, Tokyo, Japan
- <sup>159</sup> Department of Physics, University of Toronto, Toronto, ON, Canada
- <sup>160</sup> <sup>(a)</sup> TRIUMF, Vancouver, BC; <sup>(b)</sup> Department of Physics and Astronomy, York University, Toronto, ON, Canada
- <sup>161</sup> Faculty of Pure and Applied Sciences, University of Tsukuba, Tsukuba, Japan
- <sup>162</sup> Department of Physics and Astronomy, Tufts University, Medford, MA, United States
- <sup>163</sup> Centro de Investigaciones, Universidad Antonio Narino, Bogota, Colombia
- <sup>164</sup> Department of Physics and Astronomy, University of California Irvine, Irvine, CA, United States
- <sup>165</sup> <sup>(a)</sup> INFN Gruppo Collegato di Udine, Sezione di Trieste, Udine; <sup>(b)</sup> ICTP, Trieste; <sup>(c)</sup> Dipartimento di Chimica, Fisica e Ambiente, Università di Udine, Udine, Italy
- <sup>166</sup> Department of Physics, University of Illinois, Urbana, IL, United States
- <sup>167</sup> Department of Physics and Astronomy, University of Uppsala, Uppsala, Sweden
- <sup>168</sup> Instituto de Física Corpuscular (IFIC) and Departamento de Física Atómica, Molecular y Nuclear and Departamento de Ingeniería Electrónica and Instituto de Microelectrónica de Barcelona (IMB-CNM), University of Valencia and CSIC, Valencia, Spain
- <sup>169</sup> Department of Physics, University of British Columbia, Vancouver, BC, Canada
- <sup>170</sup> Department of Physics and Astronomy, University of Victoria, Victoria, BC, Canada
- <sup>171</sup> Department of Physics, University of Warwick, Coventry, United Kingdom
- <sup>172</sup> Waseda University, Tokyo, Japan
- <sup>173</sup> Department of Particle Physics, The Weizmann Institute of Science, Rehovot, Israel
- <sup>174</sup> Department of Physics, University of Wisconsin, Madison, WI, United States
- <sup>175</sup> Fakultät für Physik und Astronomie, Julius-Maximilians-Universität, Würzburg, Germany
- <sup>176</sup> Fachbereich C Physik, Bergische Universität Wuppertal, Wuppertal, Germany
- <sup>177</sup> Department of Physics, Yale University, New Haven, CT, United States
- <sup>178</sup> Yerevan Physics Institute, Yerevan, Armenia
- <sup>179</sup> Centre de Calcul de l’Institut National de Physique Nucléaire et de Physique des Particules (IN2P3), Villeurbanne, France

<sup>a</sup> Also at Department of Physics, King’s College London, London, United Kingdom.

<sup>b</sup> Also at Institute of Physics, Azerbaijan Academy of Sciences, Baku, Azerbaijan.

<sup>c</sup> Also at Novosibirsk State University, Novosibirsk, Russia.

<sup>d</sup> Also at Particle Physics Department, Rutherford Appleton Laboratory, Didcot, United Kingdom.

<sup>e</sup> Also at TRIUMF, Vancouver, BC, Canada.

- f* Also at Department of Physics, California State University, Fresno, CA, United States of America.
- g* Also at Tomsk State University, Tomsk, Russia.
- h* Also at CPPM, Aix-Marseille Université and CNRS/IN2P3, Marseille, France.
- i* Also at Università di Napoli Parthenope, Napoli, Italy.
- j* Also at Institute of Particle Physics (IPP), Canada.
- k* Also at Department of Physics, St. Petersburg State Polytechnical University, St. Petersburg, Russia.
- l* Also at Chinese University of Hong Kong, China.
- m* Also at Department of Financial and Management Engineering, University of the Aegean, Chios, Greece.
- n* Also at Louisiana Tech University, Ruston, LA, United States of America.
- o* Also at Institutio Catalana de Recerca i Estudis Avancats, ICREA, Barcelona, Spain.
- p* Also at Department of Physics, The University of Texas at Austin, Austin, TX, United States of America.
- q* Also at Institute of Theoretical Physics, Ilia State University, Tbilisi, Georgia.
- r* Also at CERN, Geneva, Switzerland.
- s* Also at Ochadai Academic Production, Ochanomizu University, Tokyo, Japan.
- t* Also at Manhattan College, New York, NY, United States of America.
- u* Also at Institute of Physics, Academia Sinica, Taipei, Taiwan.
- v* Also at LAL, Université Paris-Sud and CNRS/IN2P3, Orsay, France.
- w* Also at Academia Sinica Grid Computing, Institute of Physics, Academia Sinica, Taipei, Taiwan.
- x* Also at Laboratoire de Physique Nucléaire et de Hautes Energies, UPMC and Université Paris-Diderot and CNRS/IN2P3, Paris, France.
- y* Also at School of Physical Sciences, National Institute of Science Education and Research, Bhubaneswar, India.
- z* Also at Dipartimento di Fisica, Sapienza Università di Roma, Roma, Italy.
- aa* Also at Moscow Institute of Physics and Technology State University, Dolgoprudny, Russia.
- ab* Also at Section de Physique, Université de Genève, Geneva, Switzerland.
- ac* Also at International School for Advanced Studies (SISSA), Trieste, Italy.
- ad* Also at Department of Physics and Astronomy, University of South Carolina, Columbia, SC, United States of America.
- ae* Also at School of Physics and Engineering, Sun Yat-sen University, Guangzhou, China.
- af* Also at Faculty of Physics, M.V. Lomonosov Moscow State University, Moscow, Russia.
- ag* Also at National Research Nuclear University MEPhI, Moscow, Russia.
- ah* Also at Institute for Particle and Nuclear Physics, Wigner Research Centre for Physics, Budapest, Hungary.
- ai* Also at Department of Physics, Oxford University, Oxford, United Kingdom.
- aj* Also at Department of Physics, Nanjing University, Jiangsu, China.
- ak* Also at Institut für Experimentalphysik, Universität Hamburg, Hamburg, Germany.
- al* Also at Department of Physics, The University of Michigan, Ann Arbor, MI, United States of America.
- am* Also at Discipline of Physics, University of KwaZulu-Natal, Durban, South Africa.
- an* Also at University of Malaya, Department of Physics, Kuala Lumpur, Malaysia.
- \* Deceased.